### The boundary effect of anomaly-induced action

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Alcospa

- Scenario: Quantum field in curved spacetime
- Goal: Compute  $\langle T_{\mu\nu} \rangle$  of some system.
- Method: Anomaly-induced action
- (Our work) Improvement: Including the Boundary term in above action

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- Goal: Compute  $\langle T_{\mu\nu} \rangle$  of some system. (Why?)
- Method: Anomaly-induced action
- (Our work) Improvement: Including the Boundary term in above action

Ans: Because it is important!

- $\langle 0 | T_{\mu\nu} | 0 \rangle$  However, the concept of the vacuum states is always ambiguous.
- (?) (?) Different observers (frame) will realize different quantum states as "their vacuum state".

(Ex: Minkowski & Rindler vacuum? Boulware & Hartle-Hawking & Unruh?)

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$$S_{\rm anom} = \frac{1}{96\pi} \int d^2 x \sqrt{-g} \left[ g^{\alpha\beta} \nabla_{\alpha} \varphi \nabla_{\beta} \varphi - 2R\varphi \right] \quad \text{and} \quad \Box \varphi = -R$$

$$\blacktriangleright T_{\mu\nu}^{\rm anom} = -\frac{2}{\sqrt{-g}} \frac{\delta S_{\rm anom}}{\delta g^{\mu\nu}} = \frac{1}{24\pi} \left[ -\varphi_{;\mu\nu} + g_{\mu\nu} \Box \varphi - \frac{1}{2} \varphi_{;\mu} \varphi_{;\nu} + \frac{1}{4} g_{\mu\nu} \varphi_{;\alpha} \varphi^{;\alpha} \right]$$

People expected that this action can be variated to **get renormalized stress tensor of conformal theory** directly . However, **the correspondence between the solutions the and vacuum states are unclear**.

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Key idea: **boundary -> information of specific frames -> vacuum.** 

$$\Rightarrow S_{\text{anom}} = \frac{1}{96\pi} \int d^2 x \sqrt{-g} \left[ g^{\alpha\beta} \nabla_{\alpha} \varphi \nabla_{\beta} \varphi - 2R\varphi \right] \text{ and } \Box \varphi = -R \text{ and } n^{\nu} \nabla_{\nu} \varphi = 2K$$

$$(i) \quad \left\langle 0 \middle| \mathbf{T}_{\mu\nu} \middle| 0 \right\rangle \quad \text{VEV for what vacuum state is clear now. (Including excitation states.)}$$

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(ii) The extended version of this approach can be used in bounded (finite) region in any given manifold.

# Quantum fields in curved spacetime – Standard approach

Consider the simplest case, a scalar field in curved spacetime with the Lagrangian :

$$\mathcal{L} = \int \sqrt{-g(-\frac{1}{2}g^{\mu\nu}\nabla_{\mu}\phi\nabla_{\nu}\phi - \frac{1}{2}m^2\phi^2 - \xi \operatorname{R}\phi^2)}$$

The equation of motion:  $(\Box + m^2 + \xi R)\phi(\mathbf{x}) = 0$ 

Solving a complete solution set:  $\{u_k(x) \mid all \; k\}$ 

Quantize the scalar field: 
$$\phi(\mathbf{x}) = \int dk [\hat{\mathbf{a}}_k u_k(\mathbf{x}) + \hat{\mathbf{a}}_k^{\dagger} u_k^{*}(\mathbf{x})]$$

The vacuum state is thus defined by:  $\hat{a}_i | 0 \rangle = 0$ 

However, there is no unique way to choose a complete solution set, therefore the concept of **"vacuum" is ambiguous** in general.

Conformally coupled scalar field in n-dim:

$$\xi(n) = \frac{n-2}{4(n-1)}$$
 and m = 0.

### Quantum fields in curved spacetime – Renormalization

Compute the (one-loop) effective action  $S_{eff}$  by path integral :  $e^{-S_{eff}} = \int [D\phi] e^{-S[\phi]}$ 

Derive the expectation value of stress tensor

By the DeWitt-Schwinger expansion, this effective Lagrangian in **n**-dim can be expand

$$\text{r from S}_{\text{eff}}: -\frac{2}{\sqrt{-g}} \frac{\delta S_{eff}[g_{\mu\nu}]}{\delta g^{\mu\nu}} = \frac{\langle out, 0 | T_{\mu\nu} | 0, in \rangle}{\langle out, 0 | 0, in \rangle} = \langle T_{\mu\nu} \rangle$$

$$\text{led as} \quad \square \quad L_{eff}(x) \approx \frac{1}{2(4\pi)^{n/2}} \sum_{j=0}^{\infty} a_j(x) m^{n-2j} \Gamma(j - \frac{n}{2})$$

The singularities of the **gamma function** are simple poles at zero and the negative integers

$$L_{div} \xrightarrow{2-\dim} \begin{bmatrix} a_0(x) = 1, \\ a_1(x) = (\frac{1}{6} - \xi)R, \\ a_2(x) = \frac{1}{2}(\frac{1}{6} - \xi)^2 R^2 + \frac{1}{180}(R_{\mu\nu\alpha\beta}R^{\mu\nu\alpha\beta} - R_{\mu\nu}R^{\mu\nu}) - \frac{1}{6}(\frac{1}{5} - \xi)\Box R \end{bmatrix}$$

PS: The UV divergent behavior here is independent of choice of quantum state.

• In conformal theory, the stress tensor is traceless in general.

$$S[\bar{g}_{\mu\nu}(x)] = S[g_{\mu\nu}(x)] + \int d^n x \frac{\delta S[g_{\mu\nu}]}{\delta \bar{g}^{\alpha\beta}} \delta \bar{g}^{\alpha\beta}(x) = S[g_{\mu\nu}(x)] - \int d^n x \sqrt{-\bar{g}} T^{\lambda}_{\lambda}[\bar{g}_{\mu\nu}(x)] \delta \sigma$$
$$\Longrightarrow T^{\lambda}_{\lambda}[g_{\mu\nu}] = -\frac{1}{\sqrt{-g}} \frac{\delta S[\bar{g}_{\mu\nu}]}{\delta \sigma}|_{\sigma=0}$$

 However, after renormalization, due to the counter terms the conformal symmetry is broken and it results in an nonzero trace of stress tensor, which is called conformal anomaly.

• For example, in 2-dim: 
$$\langle T^{\mu}_{\mu} \rangle_{ren} = -\frac{R}{24\pi}$$

• By using **conformal symmetry** and **the counter terms**, the anomalyinduced action can be built.

2-dim Anomaly-induced action (**Polyakov action** [3]):

$$S_{\text{anom}} = \frac{1}{96\pi} \int d^2 x \sqrt{-g} \left[ g^{\alpha\beta} \nabla_{\alpha} \varphi \nabla_{\beta} \varphi - 2R\varphi \right]$$
  
where the auxiliary scalar field should satisfy the field equation:  $\Box \varphi = -R$   
 $\downarrow T_{\mu\nu}^{\text{anom}} = -\frac{2}{\sqrt{-g}} \frac{\delta S_{\text{anom}}}{\delta g^{\mu\nu}} = \frac{1}{24\pi} \left[ -\varphi_{;\mu\nu} + g_{\mu\nu} \Box \varphi - \frac{1}{2} \varphi_{;\mu} \varphi_{;\nu} + \frac{1}{4} g_{\mu\nu} \varphi_{;\alpha} \varphi^{;\alpha} \right]$   
 $\downarrow g^{\alpha\beta} T_{\alpha\beta}^{\text{anom}} = \frac{1}{24\pi} \Box \varphi = -\frac{1}{24\pi} R$ 

 People expected and checked that the stress tensor from above action can be used to solve renormalized stress tensor corresponding to various quantum states. [1][2]

[1] R. Balbinot, A. Fabbri and I. L. Shapiro, Nucl. Phys. B 559, 301 (1999).

[2] E. Mottola and R. Vaulin, Phys. Rev. D 74, 064004 (2006).

[3] A. M. Polyakov, Phys. Lett. B 103, 207 (1981).

Consider the Wess-Zumino action, which was first found by Polyakov:

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$$\begin{aligned} \nabla_{WZ}[\bar{g},\sigma] &= S_{ren}[\bar{g}] - S_{ren}[g] \\ &= \left[ S_{eff}[\bar{g}] - S_{eff}[g] \right] - \left[ S_{ct}[\bar{g}] - S_{ct}[g] \right] \\ &= 0 - \left[ S_{ct}[\bar{g}] - S_{ct}[g] \right] \\ &= \frac{1}{24\pi} \lim_{n \to 2} \left[ \frac{\int d^2 x \sqrt{-\bar{g}} \bar{R} - \int d^2 x \sqrt{-g} R}{n-2} \right] \\ &= S_{anom}[\bar{g}] - S_{anom}[g] \end{aligned}$$

Recall the divergent term in 2-dim come from 
$$a_1$$
:  

$$S_{ct} = \frac{-1}{24\pi} \lim_{n \to 2} \int d^2x \sqrt{-g} \frac{R}{(n-2)}$$

$$S_{\text{anom}}[g] = \frac{1}{96\pi} \int d^2x \sqrt{-g} \int d^2x' \sqrt{-g'} RD_2(x, x') R' \quad \text{where } \Box D_2(x, x') = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} N = -\frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta(x - x')}{\sqrt{-g'}} \prod_{11} \frac{\delta$$

Take the Schwarzschild metric as an example:  $ds^2 = -(1 - \frac{2M}{r})dt^2 + (1 - \frac{2M}{r})^{-1}dr^2$ 

$$\Box \varphi = -R \quad \clubsuit \quad \text{(The most general solutions which result in stationary stress tensor)} \\ \varphi = At - ln(1 - \frac{2M}{r}) + B\frac{1}{2M}r^* \qquad \text{where } r^* \equiv r + 2M\ln(\frac{r}{2M} - 1)$$

(i) A = B = 0 \Rightarrow Boulware vacuum solution: 
$$T_{\mu\nu} = \frac{1}{24\pi} \begin{pmatrix} \frac{-4Mr+7M^2}{r^4} & 0\\ 0 & -\frac{M^2}{r^2(r-2M)^2} \end{pmatrix} = \langle B | \hat{T}_{\mu\nu} | B \rangle$$

(ii) A = 0, B = 1/(2M)

Hartle-Hawking vacuum solution:

$$\begin{bmatrix} T_{tt} = \frac{1}{24\pi} \left( \frac{7M^2}{r^4} - \frac{4M}{r^3} + \frac{1}{16M^2} \right) \\ T_{rr} = \frac{1}{24\pi} \left( 1 - \frac{2M}{r} \right)^{-2} \left( \frac{1}{16M^2} - \frac{M^2}{r^4} \right) = \left\langle \mathbf{H} \right| \hat{\mathbf{T}}_{\mu\nu} \left| \mathbf{H} \right\rangle \\ T_{tr} = T_{rt} = 0 \end{bmatrix}$$

(iii) A = -B = -1/(4M)

#### Unruh vacuum solution:

$$\begin{cases} T_{tt} = \frac{1}{24\pi} \left( \frac{1}{32M^2} + \frac{7M^2}{r^4} - \frac{4M}{r^3} \right) \\ T_{rr} = \frac{1}{24\pi} \left( 1 - \frac{2M}{r} \right)^{-2} \left( \frac{-M^2}{r^4} + \frac{1}{32M^2} \right) = \left\langle \mathbf{U} \middle| \hat{\mathbf{T}}_{\mu\nu} \middle| \mathbf{U} \right\rangle \\ T_{tr} = T_{rt} = \frac{1}{24\pi} \left[ \frac{-1}{32M^2} \left( 1 - \frac{2M}{r} \right)^{-1} \right] \end{cases}$$

$$= 12$$

- Introduce the boundary part of divergent terms.
- Recalculate the anomaly-induced action with boundary effect.

$$S_{anom}[g] = \frac{1}{96\pi} \{ \int_{\mathcal{M}} d^2x \sqrt{-g} (-\varphi \Box \varphi - 2\varphi R) + \int_{\Sigma} d^1x \sqrt{\gamma} (\varphi n_\mu \nabla^\mu \varphi - 4\varphi K) \}$$

where the auxiliary scalar field should satisfy the original equation and an additional boundary constraint:

$$\begin{bmatrix} \Box \varphi = -R \\ n^{\nu} \nabla_{\nu} \varphi = 2K \end{bmatrix}, \qquad x \in \Sigma$$

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Introducing the boundary part of divergent term: [4][5]

$$S_{ct} = \frac{-1}{24\pi} \lim_{n \to 2} \int d^2 x \sqrt{-g} \frac{R}{(n-2)} \longrightarrow \frac{-1}{24\pi} \frac{\int_{\mathcal{M}} d^2 x \sqrt{-g} R + 2\int_{\Sigma} d^1 x \sqrt{-\gamma} K}{n-2}$$

$$\Gamma_{WZ}[\bar{g},\sigma] = S_{ren}[\bar{g}] - S_{ren}[g] = \left[S_{eff}[\bar{g}] - S_{eff}[g]\right] - \left[S_{ct}[\bar{g}] - S_{ct}[g]\right] = 0 - \left[S_{ct}[\bar{g}] - S_{ct}[g]\right] = \frac{1}{24\pi} \lim_{n \to 2} \left[\frac{(\int d^2x \sqrt{-\bar{g}}\bar{R} + 2\int d^1x \sqrt{-\bar{\gamma}}\bar{K}) - (\int d^2x \sqrt{-g}R + 2\int d^1x \sqrt{-\gamma}K)}{n-2}\right] = S_{anom}[\bar{g}] - S_{anom}[g]$$

[4] T. P. Branson and P. B. Gilkey, Commun. Part. Di. Eq. 15, 245 (1990).
[5] I. G. Avramidi, Phys. Atom. Nucl. 56, 138 (1993) [Yad. Fiz. 56N1, 243 (1993)].

#### Schwarzschild metric and its various coordinates

$$\begin{split} ds^2 &= -(1 - \frac{2M}{r})dt^2 + (1 - \frac{2M}{r})^{-1}dr^2 = (1 - \frac{2M}{r})(-dt^2 + dr^{*2}) = -(1 - \frac{2M}{r})dudu \\ &= -\frac{32M^3}{r}e^{-\frac{r}{2M}}dUdV = -\frac{32M^3}{r}e^{-\frac{r}{2M}}(-dT_H{}^2 + dR_H{}^2) \\ &= -\frac{8M^2}{r}(\frac{r}{2M} - 1)^{\frac{1}{2}}e^{\frac{t-r}{2M}}dUdv = \frac{8M}{r}(\frac{r}{2M} - 1)^{\frac{1}{2}}e^{\frac{t-r}{2M}}(-dT_U{}^2 + dR_U{}^2), \end{split}$$
where  $r^* := r + 2M\ln(\frac{r}{2M} - 1), \quad u := t - r^*, \quad v := t + r^*, \\ U := -e^{\frac{-u}{4M}}, \quad V := e^{\frac{v}{4M}}, \quad T_H := \frac{1}{2}(V + U), \quad R_H := \frac{1}{2}(V - U), \\ T_U := \frac{1}{2}(v + U), \quad R_U := \frac{1}{2}(v - U). \end{split}$ 

- The vacuum states depend on observer (frame) choosing.
- Key idea: introduce **boundaries corresponding to specific frames** can naturally give us the corresponding vacuum solutions.
- For example, we consider Schwarzschild black hole metric again with the appropriate choices of boundaries corresponding to different vacuum states.



The dashed lines are  $(T_H, R_H)$  coordinate.

The dashed lines are (t,r) coordinate.

The dashed lines are (T<sub>U</sub>,R<sub>U</sub>) coordinate.

At first, consider "Boulware" boundary:

$$\Box \varphi = -R$$

$$n^{\nu} \nabla_{\nu} \varphi = 2K \quad \Rightarrow \quad \varphi = \ln(1 - \frac{2M}{r}) + A_2 t + \int d\omega \, c(\omega) \cos[\omega r^*] e^{i\omega t}$$

The most general solutions which result in "stationary" stress tensor is  $\varphi = \ln(1 - \frac{2M}{r}) + A_2 t$   $\downarrow T_{\mu\nu} = \frac{1}{24\pi} \left( \begin{array}{c} \frac{-4Mr + 7M^2}{r^4} + \frac{A_2^2}{4} & 0 \\ 0 & -\frac{M^2}{r^2(r-2M)^2} + \frac{A_2^2 r^2}{4(r-2M)^2} \end{array} \right)$ 

> When  $A_2=0$ , the result is indeed the Boulware vacuum solution.  $A_2$  characterizes the thermal excitation based on the Boulware vacuum.

Similarly, by considering the "Hartle-Hawking" and "Unruh" boundaries, we will get

The Hartle-Hawking vacuum solution and the corresponding thermal excitation states (which is parametrized by A<sub>H</sub>):

$$\begin{bmatrix} T_{tt} = \frac{1}{24\pi} \left[ \left( \frac{7M^2}{r^4} - \frac{4M}{r^3} + \frac{1}{16M^2} \right) + \frac{A_H^2}{64M^2} \left( R_H^2 + T_H^2 \right) \right] \\ T_{rr} = \frac{1}{24\pi} \left[ \left( 1 - \frac{2M}{r} \right)^{-2} \left( \frac{1}{16M^2} - \frac{M^2}{r^4} \right) + \frac{A_H^2}{64M^2} \left( 1 - \frac{2M}{r} \right)^{-2} \left( R_H^2 + T_H^2 \right) \right] \\ T_{tr} = T_{rt} = \frac{1}{24\pi} \left[ \frac{A_H^2}{32M^2} \left( 1 - \frac{2M}{r} \right)^{-1} \left( T_H R_H \right) \right]$$

The Unruh vacuum solution and the corresponding thermal excitation states (which is parametrized by A<sub>u</sub>):

$$\begin{bmatrix} T_{tt} = \frac{1}{24\pi} \left[ \left( \frac{1}{32M^2} + \frac{7M^2}{r^4} - \frac{4M}{r^3} \right) + \frac{A_U^2}{8} \left( 1 + \frac{r - 2M}{32M^3} e^{\frac{r - t}{2M}} \right) \right] \\ T_{rr} = \frac{1}{24\pi} \left[ \left( 1 - \frac{2M}{r} \right)^{-2} \left( \frac{-M^2}{r^4} + \frac{1}{32M^2} \right) + \frac{A_U^2}{8} \left( \left( 1 - \frac{2M}{r} \right)^{-2} + \left( 1 - \frac{2M}{r} \right)^{-1} \frac{r}{32M^3} e^{\frac{r - t}{2M}} \right) \right] \\ T_{tr} = T_{rt} = \frac{1}{24\pi} \left[ \frac{-1}{32M^2} \left( 1 - \frac{2M}{r} \right)^{-1} + \frac{A_U^2}{8} \left( \left( 1 - \frac{2M}{r} \right)^{-1} - \frac{r}{32M^3} e^{\frac{r - t}{2M}} \right) \right]$$

Note: When A2 = 1/(2M), the result in previous page is the same as that of the Hartle-Hawking vacuum state. This fact is consistent with the usual understanding about Hawking temperature.

- Anomaly-induced action is a useful approach to solve  $\langle T_{\mu\nu} \rangle$
- However, the **correspondence** between its solutions and the vacuum states which people are interested in **is missing**.
- Take **boundary** effect into consideration **introduce** some **information related to frame** (and thus the corresponding vaccum).
- The extended version of anomaly-induced action now really become an independent tool to solve  $\langle 0|T_{\mu\nu}|0\rangle$ (The original one is unable to do this.)

### Thank you