

Osamu Seto, JGRG 22(2012)111413

“Curvaton induced modulated reheating”

---

**RESCEU SYMPOSIUM ON  
GENERAL RELATIVITY AND GRAVITATION**

**JGRG 22**

November 12-16 2012

Koshiba Hall, The University of Tokyo, Hongo, Tokyo, Japan



# *Curvaton induced modulated reheating*

Osamu Seto (Hokkai-Gakuen Univ.)

With Ki-Young Choi (APCTP and POSTECH)

Based on: corrected\* **Phys. Rev. D 85, 123528 (2012)**

\*Thanks to S.Yokoyama, T.Suyama and D.Wands

# Are you ready for this?

## **First year Planck observations: Full sky maps and cosmological implications.**

Authors.....

**Abstract** We present full sky microwave from the Planck first year sky survey. .... The age of the Universe.....Nonlinearity of local type  $f_{NL}^{local}(k_0=0.002 \text{ Mpc}^{-1})=15 \pm 9$  (95% CL) is detected. .... index  $n_s = 0.97 \pm 0.02$ ... The B-mode polarization is also detected. The corresponding tensor-to-scalar ratio is estimated as  $r_T(k_0=0.002 \text{ Mpc}^{-1})=0.10 \pm 0.03$  (95% CL).

# § Introduction

- Inflation

is an elegant solution to

horizon, flatness, monopole problem

and

provides the seed of density fluctuation of

adiabatic,

almost scale invariant, and **Gaussian**

as well as **tensor modes**.

# Non-Gaussianity in perturbation generation

- Inflaton  $\phi$  [Hawking, Guth and Pi,... 1982]

$$f_{\text{NL}} = \mathcal{O}(\epsilon, \eta) \quad [\text{Maldacena 2003}]$$

- Curvaton  $\sigma$  [Lyth and Wands, Moroi and Takahashi, Enqvist and Sloth 2001~2]

$$f_{\text{NL}} = \mathcal{O}(1/R) \quad [\text{Lyth et al 2003,...}]$$

- Modulated reheating by a light scalar  $\sigma$  [Kofman 2003, Dvali et al 2004]

$$\delta\Gamma_{\phi[\sigma]} \neq 0 \quad \implies \quad \delta T/T \neq 0$$

$$f_{\text{NL}} = f_{\text{NL}}(\Gamma, \partial\Gamma, \dots) \quad [\text{Zaldarriaga 2004,...}]$$

# System in perturbation generation

- Inflaton  $\phi$  [Hawking, Guth and Pi,... 1982]

Inflaton  $\phi$

- Curvaton  $\sigma$  [Lyth and Wands, Moroi and Takahashi, Enqvist and Sloth 2001~2]

Inflaton  $\phi$  + Curvaton  $\sigma$

- Modulated reheating by a light scalar  $\sigma$  [Kofman 2003, Dvali et al 2004]

Inflaton  $\phi$  + Light scalar  $\sigma$

# System in perturbation generation

- Inflaton  $\phi$  [Hawking, Guth and Pi,... 1982]

Inflaton  $\phi$

- Curvaton  $\sigma$  [Lyth and Wands, Moroi and Takahashi, Enqvist and Sloth 2001~2] “massive” + “decaying” + “non  $\phi$  interacting”

Inflaton  $\phi$  + Curvaton  $\sigma$

- Modulated reheating by a light scalar  $\sigma$  [Kofman 2003, Dvali et al 2004]  What's difference?

Inflaton  $\phi$  + Light scalar  $\sigma$

“massless” + “ $\phi$  interacting”

## § $\phi$ - $\sigma$ interaction

- Inflaton  $\phi$  - Curvaton/light scalar  $\sigma$  interactions may exist.

- If  $\phi$  and  $\sigma$  are singlet, there might be

$$\mathcal{L}_{\text{int}} = \phi \sigma \mathcal{O}_{\text{SM}} + \text{tiny } \phi^2 \sigma^2 + \dots, .$$

- ex.)

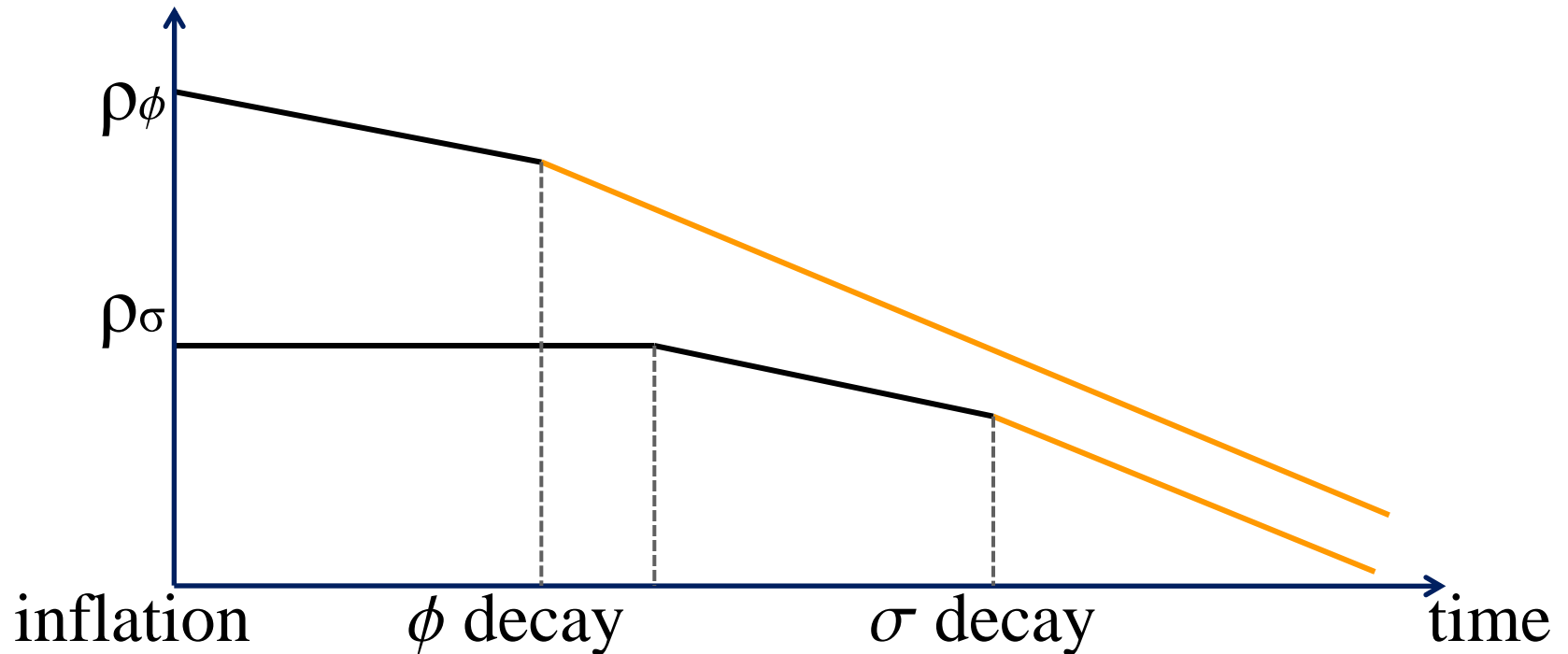
$$\mathcal{L}_{\text{int}} = \phi \sigma |\Phi|^2 \quad \text{then} \quad \Gamma_{\phi} \propto \sigma^2$$

- Modulated reheating by curvaton [Suyama et al 2010]



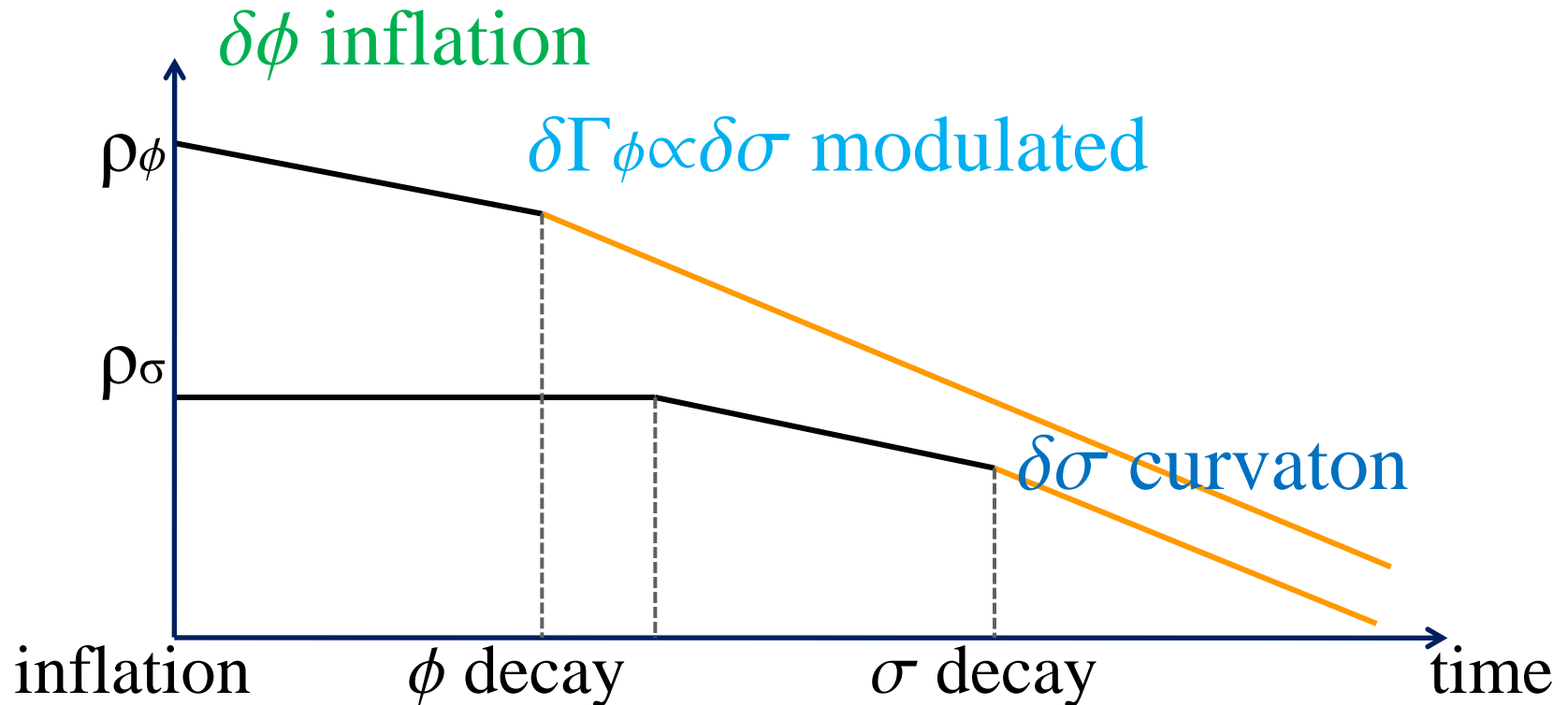
# § General analysis

- Inflaton  $\phi$  and another scalar  $\sigma$
- Thermal history after inflation



# § General analysis

- Inflaton  $\phi$  and another scalar  $\sigma$
- Thermal history after inflation



# § § Formula of perturbation

- After  $\sigma$  decay:

**inflaton-modulated mixed** [Zaldarriaga 2004, Ichikawa et al 2008] + **curvaton**

- On uniform  $\sigma$  density hypersurface

$$\zeta = (1 - R)\zeta_r + R\zeta_\sigma = \zeta_r + \frac{R}{3}S_\sigma$$

$$R \equiv \frac{3\rho_\sigma}{4\rho_r + 3\rho_\sigma} \Big|_{H=\Gamma_\sigma} \quad S_\sigma \equiv 3(\zeta_\sigma - \zeta_r)$$

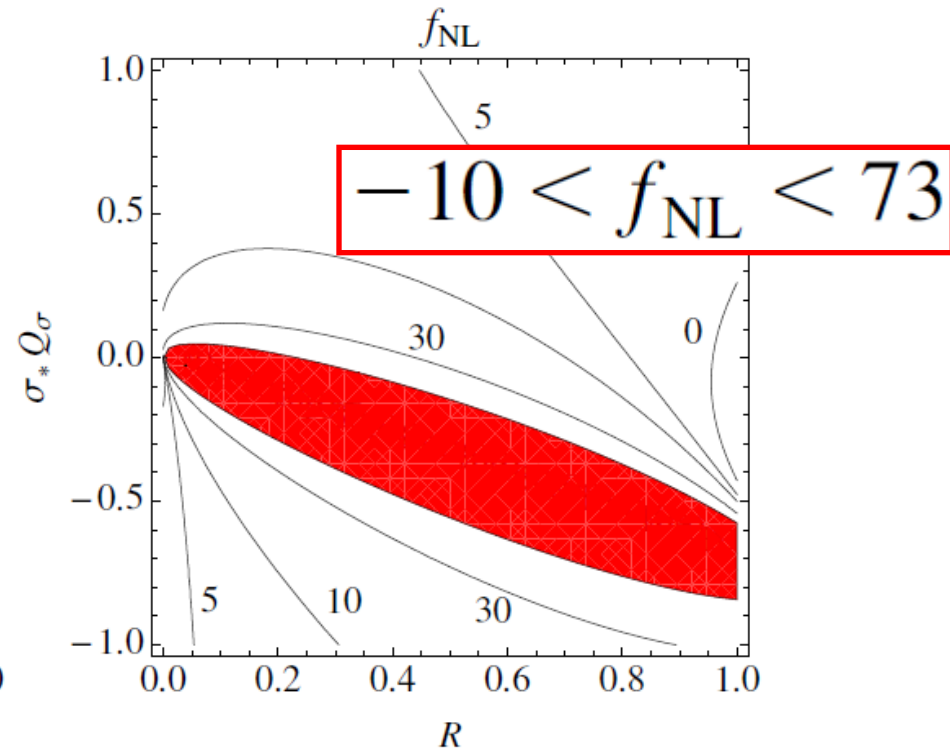
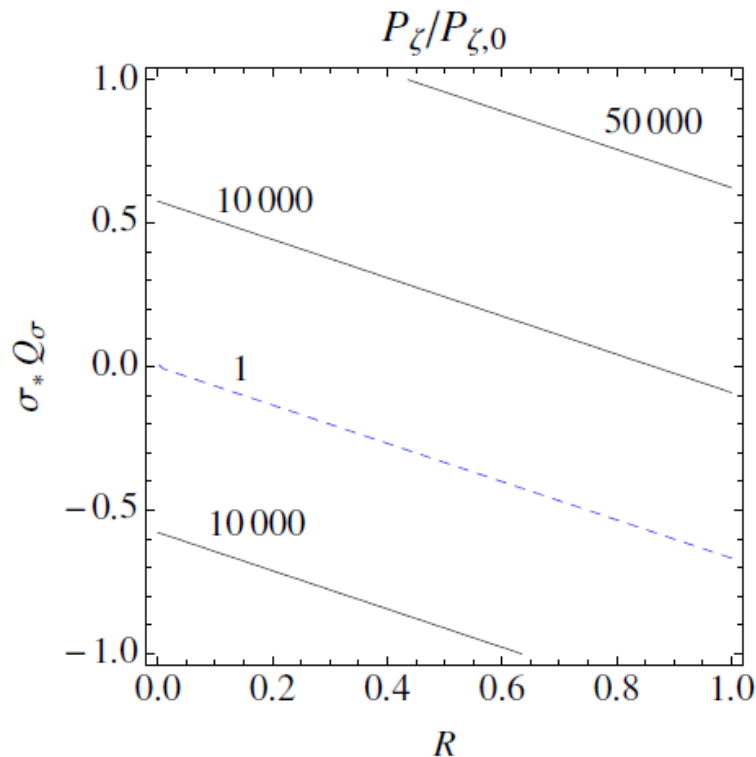
- Power spectrum

$Q$  is a function of  $\Gamma_\phi(\sigma)/H_c$

$$\mathcal{P}_\zeta = \frac{1}{2M_{\text{Pl}}^2\epsilon_*} \mathcal{P}_{\delta\phi_*} + \left( Q_\sigma + \frac{2R}{3\sigma_*} \right)^2 \mathcal{P}_{\delta\sigma_*}$$

# § § Resultant perturbation

- Important parameters:  $\sigma^*$ ,  $R$ ,  $\{\sigma^*Q_\sigma, \sigma^{*2}Q_{\sigma\sigma}, \dots\}$   
 $\phi$ - $\sigma$  ratio      Modulated

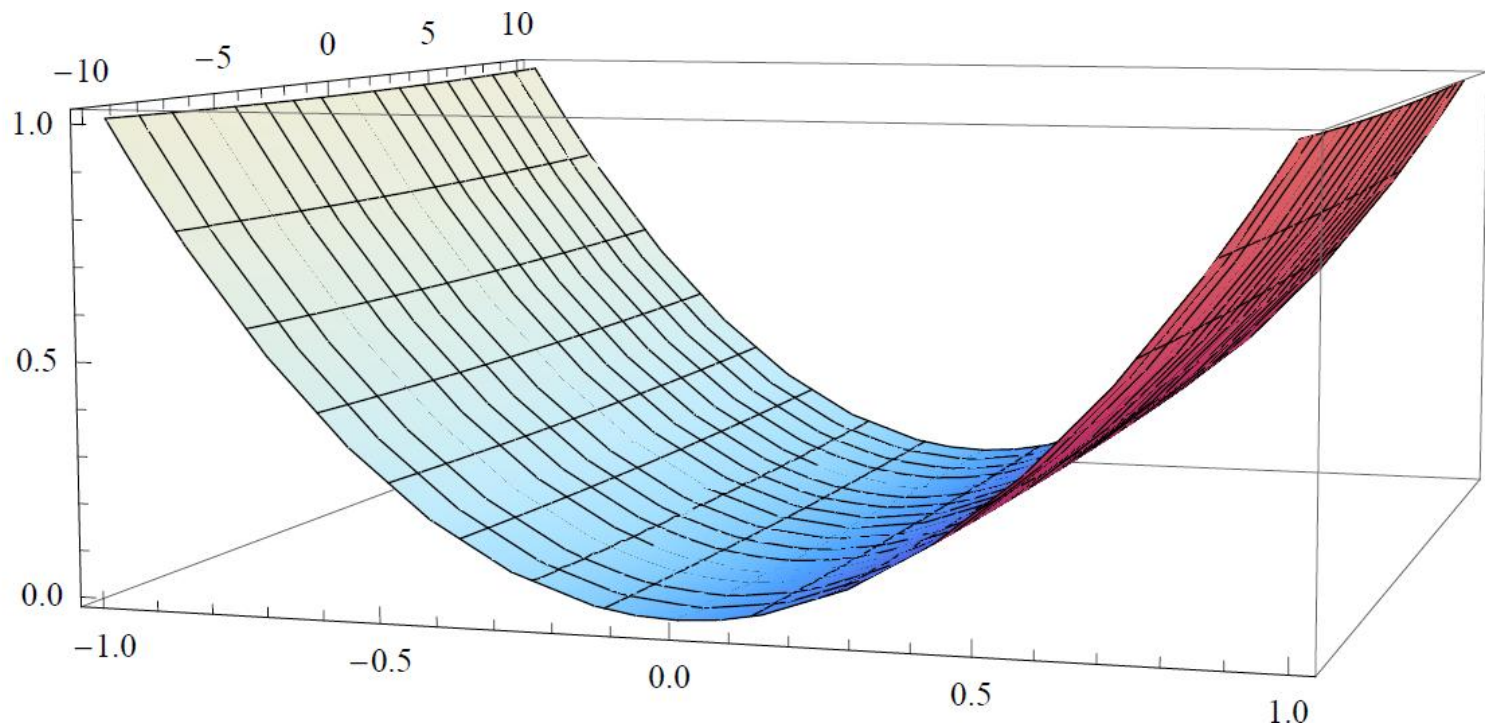


- Small  $\sigma^*$  case : mainly  $\sigma$  originated

# § Simplest case

- Potential

$$V[\phi, \sigma] = \frac{1}{2}m_\phi^2 \phi^2 + \frac{1}{2}m_\sigma^2 \sigma^2$$



# § § Power spectrum

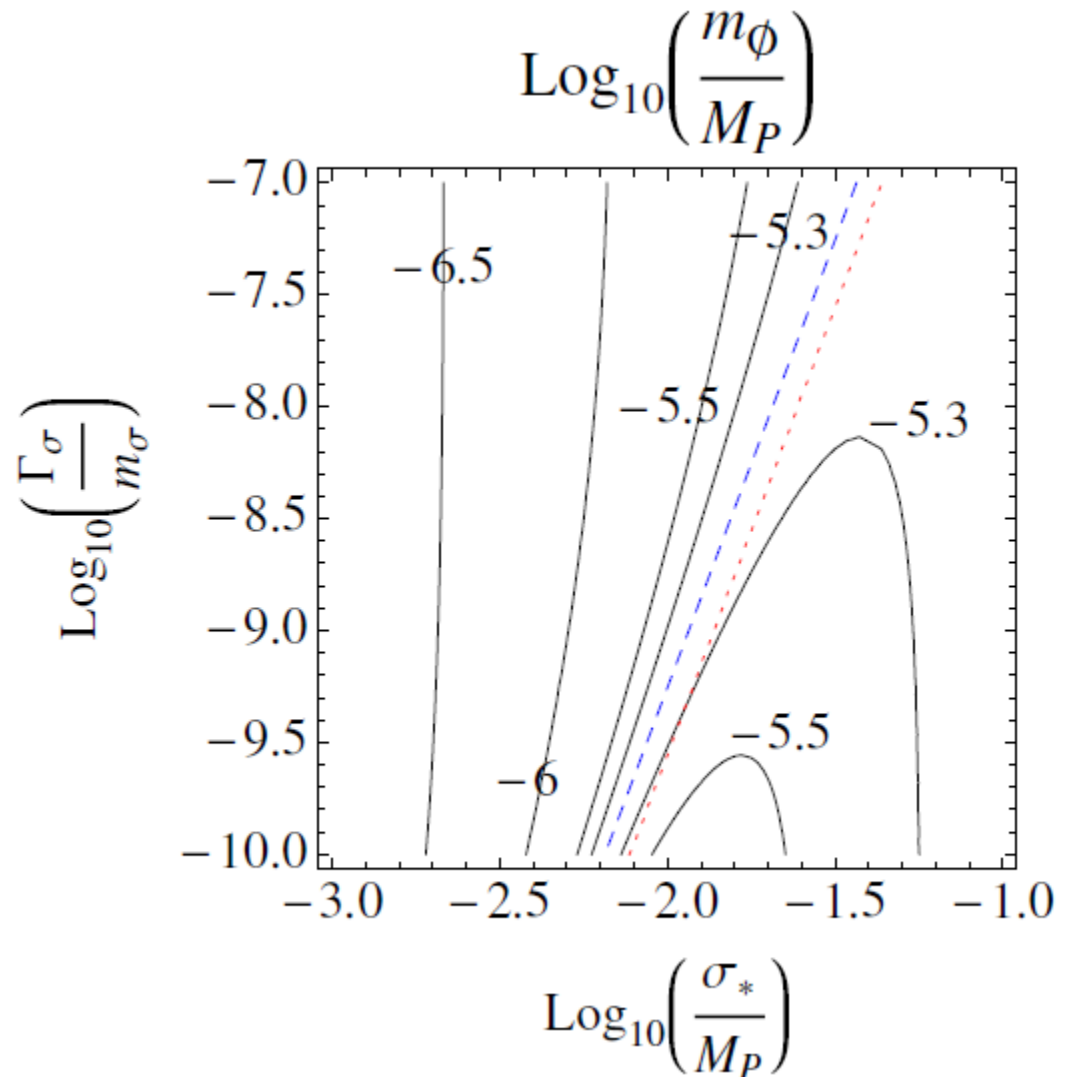
- Power spectrum

$$\mathcal{P}_\zeta \simeq \frac{1}{6(2\pi)^2} \frac{m_\phi^2}{M_P^2} (2N_{\text{inf}} + 1)^2 + \left\{ Q_\sigma \sigma_* + \frac{2R}{3} \right\}^2 \frac{m_\phi^2}{3(2\pi\sigma_*)^2} (2N_{\text{inf}} + 1)$$

# § § Power spectrum

- Power spectrum

$$\mathcal{P}_\zeta \simeq \frac{1}{6(2\pi)^2} \frac{m_\phi^2}{M_P^2} (2N_{\text{inf}})$$



# § § Tensor and nonlinearity

- Model I
- Interaction

$$\mathcal{L}_{\text{int}} = \lambda |\Phi|^2 \phi \sigma$$

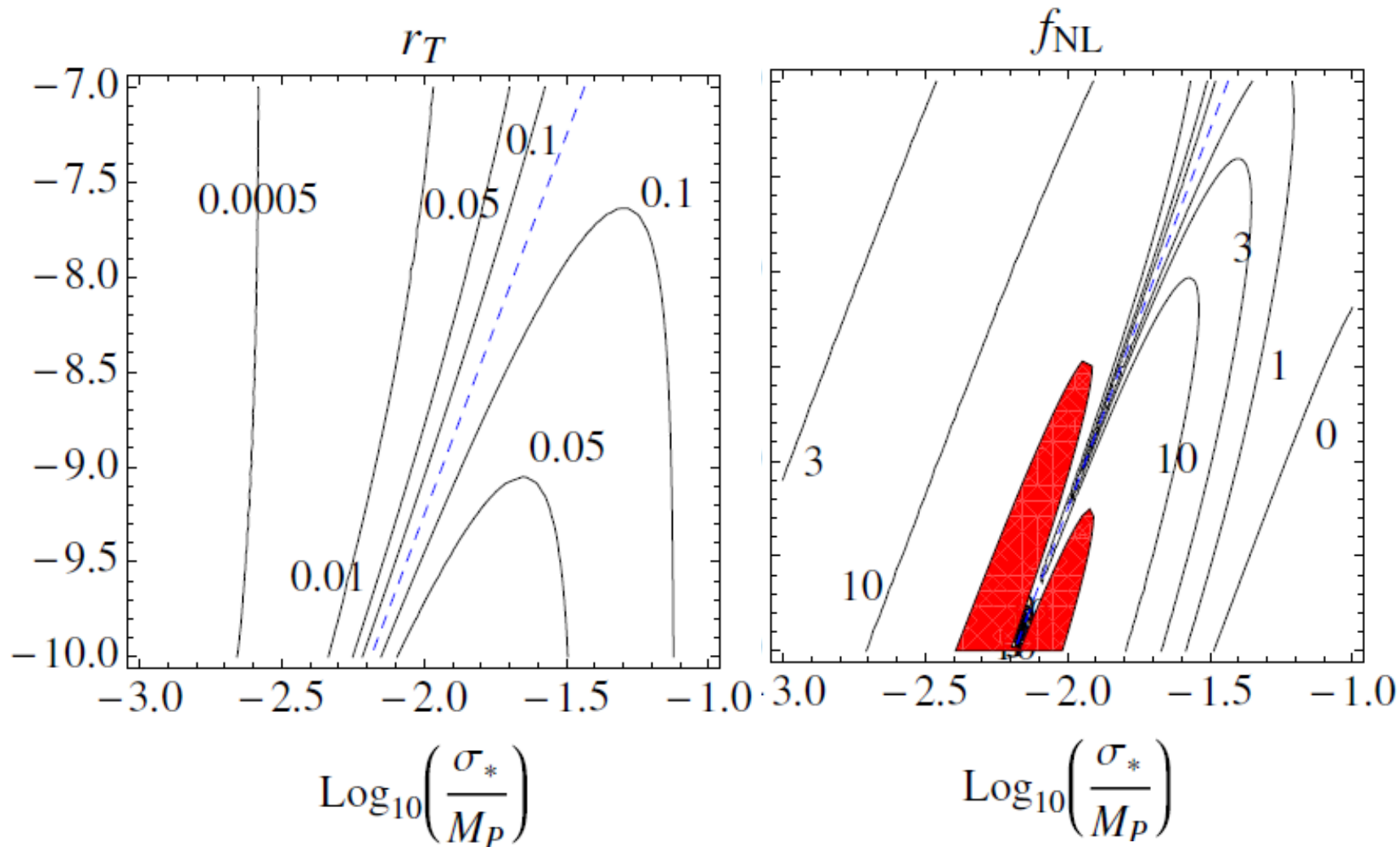
- Decay rate

$$\Gamma_{\phi}^{\text{CD}}(\sigma) = \frac{1}{8\pi m_{\phi}} \lambda^2 \sigma^2$$

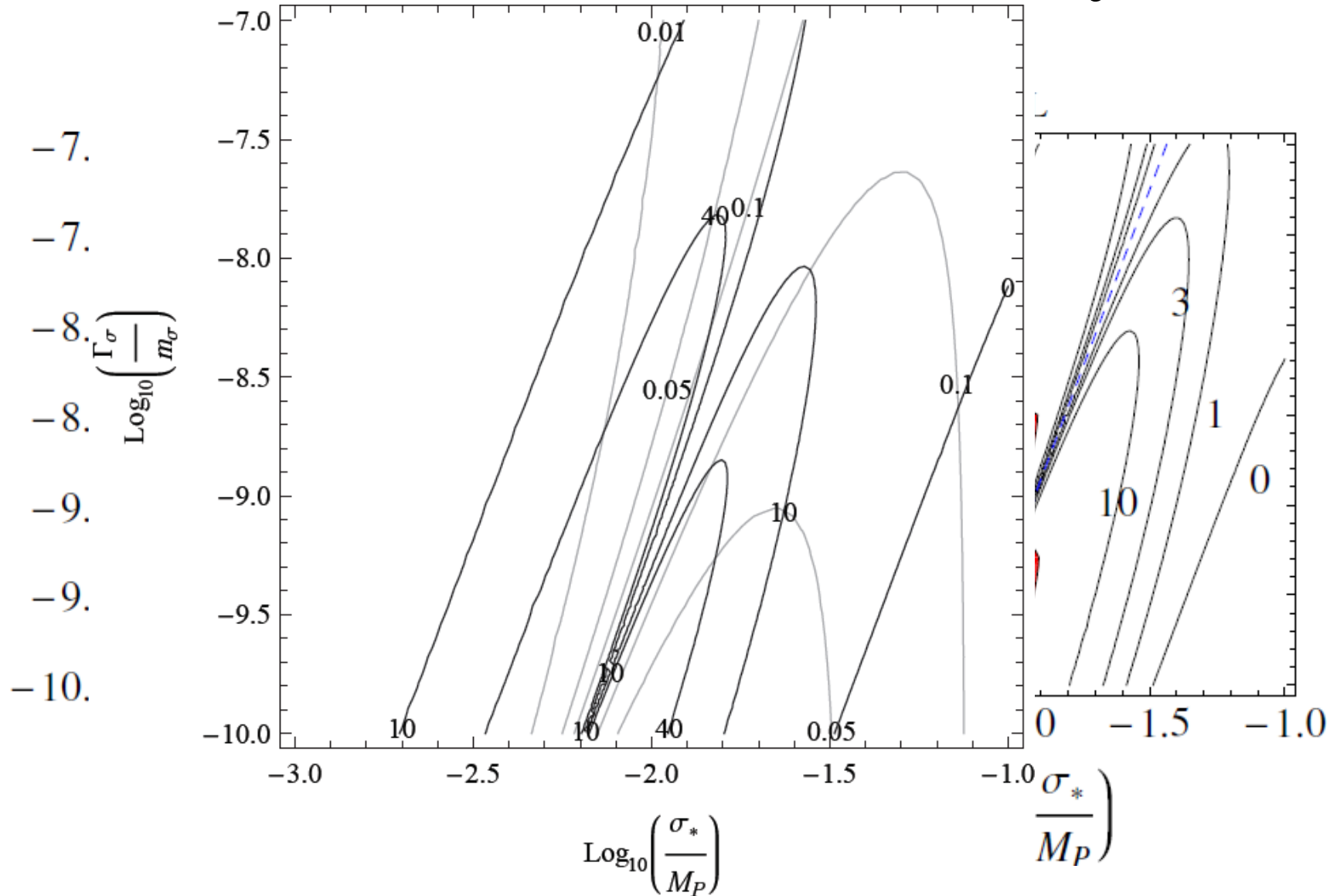
$$(Q_{\sigma} \sigma_*, Q_{\sigma\sigma} \sigma_*^2, Q_{\sigma\sigma\sigma} \sigma_*^3) = \left( -\frac{1}{3}, \frac{1}{3}, -\frac{2}{3} \right)$$



# § § Tensor and nonlinearity



# § § Tensor and nonlinearity



# § Summary

- *“Inflaton-modulated-curvaton mixed”*
- Interaction between  $\phi$  and  $\sigma$  is a subject to be studied.
- An interesting possibility: simultaneous large  $f_{\text{NL}}$  and  $r_{\text{T}}$