

Large-Scale Suppression from Stochastic Inflation

Florian Kühnel



Excellence Cluster 'Universe' & Arnold Sommerfeld Center for Theoretical Physics, LMU Munich

Overview

In this work (based on [1-4]):

- We introduce a **novel technique** to properly calculate correlation functions and in turn the **power spectrum**,
- derive a **large-scale damping**, which
- solves the problem of infra-red divergencies of two-point correlation functions.

Application to CMB Data



• We provide a **better description of CMB data**.

Setup

One fruitful framework to study **large quantum fluctuations** is **stochastic inflation**:

- Split the quantum field Φ into a short- and longwavelength part, Φ = φ + φ, and treat φ as classical, evolving in a bath of the quantum modes φ, giving rise to a
- Langevin-type differential equation,

 $(\Box + m^2)\varphi(t, \boldsymbol{x}) = h(t, \boldsymbol{x}).$

Technicalities

Angular temperature correlation function $\mathcal{C}(\theta)$.

Large-Scale Damping



We would like to compute the noise-averaged

• connected **correlation functions**, and hence their **generating functional** $\ln(\mathcal{Z})$.

This is **difficult** to achieve directly. Hence,

• use the **replica trick**:

 $\overline{\ln(\mathcal{Z})} = \lim_{n \to 0} \ln\left(\left(\overline{\mathcal{Z}^n}\right)^{1/n}\right),\,$

• i.e., **compute** only $\overline{Z^n}$ (instead of the logarithm):

$$\overline{\mathcal{Z}^n} = \int \prod_{a=1}^n \mathcal{D}[\varphi_a] \exp\left(-\mathcal{S}_n[\varphi_1,\ldots,\varphi_n]\right).$$

• **Approximate** the action $S_n[\varphi]$ by

$$S_0[\varphi] := \frac{1}{2} \sum_{a,b=1}^n \int \mathrm{d}t \, \int_k \, \mathrm{G}^{-1}{}_{ab}(t,k) \, \varphi_a(t,k) \, \varphi_b(t,-k)$$

Power spectrum $\mathcal{P}(k)$ of a massless test field.

Formulæ

For a single scalar-field model with quartic self-coupling we derive:

- the **power spectrum** $\mathcal{P}(t,k)$ at late times:
- a, o-1
- Use the Feynman-Jensen inequality: $\ln(\mathcal{Z}) \ge \ln(\mathcal{Z}_0) + \langle \mathcal{S}_0 - \mathcal{S} \rangle_0.$
- Maximise the r.h.s. of this inequality, by writing $G^{-1}{}_{ab}(t,k) := G_0^{-1}(t,k) \,\delta_{ab} - \sigma_{ab}$, and varying with respect to the self-energy matrix (σ_{ab}) .

 $\mathcal{P}(t,k) \simeq \frac{H^2}{(2\pi)^2} \begin{cases} \left(\frac{k}{k_*(t)}\right)^3 & : k \ll k_*(t) ,\\ 1 & : k_*(t) \ll k \ll 1/|\tau(t)| , \end{cases}$

• the scale of deviation from scale invariance:

$$k_* \Big|_{\text{today}} \simeq \sqrt[3]{\frac{3}{2\pi^2}} \sqrt[3]{\lambda N} e^{-N} \left(\frac{H}{T_{\text{reh}}}\right) \left(\frac{T_0}{H_0}\right) H_0 \ .$$

[1] F. Kühnel and D. J. Schwarz, Phys. Rev. D 78, 103501 (2008).
[2] F. Kühnel and D. J. Schwarz, Phys. Rev. D 79, 44009 (2009).

[3] F. Kühnel, Ph.D. Thesis (2009).
[4] F. Kühnel and D. J. Schwarz, submitted to Phys. Rev. Lett. (2010).