

Cosmological Behaviour of a Parity and Charge-Parity Violating Varying Alpha Theory.

Debaprasad Maity



LeCosPA, National Taiwan University

COSMO/CosPA 2010 September 27 - October 1 2010, The University of Tokyo, Japan

D. Maity and P. Chen, arXiv:1005.5104

Plan of my talk



- Introduction to Varying alpha theory
- Parity violating extension and motivation
- Effect of this violation in different cosmic phenomena
- Conclusions

Introduction



- For the last several years, this idea of varying α constant has got much attention, mainly because
 - String theory
 - Studies on absorption spectrum of quasars, M. Murphy etal, Mon.
 Not. R. astr. Soc., 327, 1208 (2001), PRL, 99, 239001,(2007); J.K. Webb etal, PRL 82, 884 (1999) and PRL 87, 091301 (2001).

• α : smaller in the past, at z = 1 - 3.5; $\frac{\delta \alpha}{\alpha} = (-0.570.10)10^{-5}$.

- Motivated by this observation, Sandvik-Barrow-Magueijo extensively studied the varying alpha theory in cosmology (Bekenstein-Sandvik-Barrow-Magueijo (BSBM) theory) H. B. Sandvik, J. D. Barrow and J. Magueijo, PRL, 88 (2002); PRD 65, 063504 (2002); PRD 65, 123501 (2002); PRD 66, 043515 (2002); PLB 541, 201 (2002)
- BSBM construction is based on a model of varying alpha proposed by Bekenstein. J.D. Bekenstein, PRD 25, 1527 (1982).

Varying α Theory

- The simplest way to introduce the variation of α : Variation of electric charge as $e = e_0 e^{\phi(x)}$,
- $\phi(x)$: A dimensionless scalar field.
- Guiding principles: Shift symmetry and gauge invariance

$$\phi \to \phi + c \quad ; \quad e^{\phi} A_{\mu} \to e^{\phi} A_{\mu} + \chi_{,\mu}.$$

• The gauge invariant and shift symmetric action:

$$S_{em} = -\frac{1}{4} \int d^4x \sqrt{-g} e^{-2\phi} F_{\mu\nu} F^{\mu\nu} - \frac{\omega}{2} \int d^4x \sqrt{-g} \phi_{,\mu} \phi^{,\mu},$$

Where $F_{\mu\nu} = (e^{\phi}A_{\nu})_{,\mu} - (e^{\phi}A_{\mu})_{,\nu}$.

 $\omega = \hbar c/l^2$, where *l*: Characteristic scale above which the electric field around a point charge is exactly Coulombic.

PCP Violating Extension



- Based on this varying alpha theory, we extend the theory to incorporate parity violation.
 Motivations of this simple extension are mainly
- Recent interest in the parity violating effect on present high precession experimental observations to look for new physics
- To unify the different cosmic phenomena

PCP violating model



One of the assumptions of the above theory is time-reversal invariance. We will relax this assumption and try to analyse its implications.

Obvious term that we can add

$$S_{PV} = \frac{\beta}{8} \int d^4x \sqrt{-g} e^{-2\phi} F_{\mu\nu} \tilde{F}^{\mu\nu}$$

 β : A free dimensionless parameter

- So, total Lagrangian violates both P and CP.
- So, our total Lagrangian looks like

$$\mathcal{L} = M_p^2 R - \frac{\omega}{2} \partial_\mu \phi \partial^\mu \phi - \frac{1}{4} e^{-2\phi} F_{\mu\nu} F^{\mu\nu} + \frac{\beta}{8} e^{-2\phi} F_{\mu\nu} \tilde{F}^{\mu\nu} + \mathcal{L}_m,$$



Various cosmic phenomena

- Cosmic Birefringence Phenomena
- Non-vanishing effect on CMB polarization power spectrum because of parity violation
- Parity violating effect on Cosmic variation of fine structure constant

Cosmological Birefringence



- Cosmological birefringence (CB) is a wavelength-independent rotation of photon polarization vector after traversing a long cosmic distance.
- The equations for the polarization states $b_{\pm}(\eta) = \mathbf{B}_{0x}(\eta) \pm i\mathbf{B}_{0y}(\eta)$

$$\ddot{b}_{\pm} + 2\dot{\phi}\dot{b}_{\pm} + \left(\mathbf{k}^2 \mp 2\mathbf{k}\beta\dot{\phi}\right)b_{\pm} = 0,$$

Equation of motion for the scalar field in FRW background

$$\ddot{\phi} + 2\frac{\dot{a}}{a}\dot{\phi} = \frac{e^{-2\phi}}{\omega a^2} [-(\mathbf{E}^2 - \mathbf{B}^2) + 2\beta \mathbf{B} \cdot \mathbf{E}].$$

• We use the WKB approximation in large ω and long wavelength limit.

Cosmological Birefringence



• Assuming the form of solution for b_{\pm} to be

$$b_{\pm} = e^{ikS_{\pm}(\eta)}$$
; $S_{\pm}(\eta) = S_{\pm}^{0} + \frac{1}{k}S_{\pm}^{1} + \dots$

• Solution based on the above ansatz is

$$S^0_{\pm} = \eta \; ; \; S^1_{\pm} = -\frac{1}{2}(-2i \pm 2\beta) \int \dot{\phi} d\eta.$$

 So the expression for the optical rotation of the plane of polarization is

$$\Delta = 2\beta \int_{\eta_i}^{\eta_f} \dot{\phi} d\eta = 2\beta |\phi(\eta_f) - \phi(\eta_i)|,$$

where η_i and η_f are the initial and final time.

CMB Power Spectrum and Parity violation



- CMB power spectrum is the various correlation functions among the temperature fluctuation and polarization tensors.
- The angular distribution of CMB temperature anisotropy can be expanded as:

$$\frac{\Delta T}{T}(\mathbf{n}) = \sum_{l,m} a_{lm}^T Y_{lm}^T(\mathbf{n}) .$$

• The angular distribution of polarization tensor can also be expressed in terms of the matrix spherical harmonics

$$\mathcal{P}_{ab}^{E}(\mathbf{n}) = \sum a_{lm}^{E} Y_{lm,ab}^{E}(\mathbf{n})$$
$$\mathcal{P}_{ab}^{B}(\mathbf{n}) = \sum a_{lm}^{B} Y_{lm,ab}^{B}(\mathbf{n}) .$$

Birefringence on CMB anisotropy

• CMB power spectrum is defined as,

$$C_l^{XX'} \equiv \langle a_{lm}^X \; a_{lm}^{X'} \rangle \; ,$$

where X, X' = T, E, B

Now from the standard cosmological model

$$\langle a_{lm}^T \; a_{lm}^B \rangle = \langle a_{lm}^E \; a_{lm}^B \rangle = 0$$

• Due to explicit parity violating interaction, correlations such as C_l^{TB} and C_l^{EB} appear through the birefringence

$$C_l^{TB} = C_l^{TE} \sin 2\Delta$$
$$C_l^{EB} = \frac{1}{2} (C_l^{EE} - C_l^{BB}) \sin 4\Delta$$



Varying α Cosmology

• Assuming FRW metric ansatz,

$$ds^{2} = -dt^{2} + a(t)^{2}(dx^{2} + dy^{2} + dz^{2}),$$

• Equation of motions are

$$\left(\frac{\dot{a}}{a}\right)^2 = \frac{1}{3M_p^2} \left[\rho_m \left\{1 + e^{-2\phi}\zeta_m\right\} + e^{-2\phi}\rho_r + \rho_\phi\right] + \frac{\Lambda}{3}$$

where Λ : Cosmological constant; $\rho_{\phi} = \frac{1}{2}[\dot{\phi}^2 + V(\phi)].$

$$\ddot{\phi} + 3H\dot{\phi} = \frac{e^{-2\phi}}{\omega} \left[-2\zeta_m\rho_m + \frac{2}{a^3}\beta\langle \mathbf{E}\cdot\mathbf{B}\rangle\right],$$

where $H \equiv \dot{a}/a$, ζ_m is the fraction of matter carrying electric or magnetic charge.

Varying α Cosmology



• With the suitable boundary condition,

 $a\langle \mathbf{E}\cdot\mathbf{B}\rangle = \langle \mathbf{B}\cdot\mathbf{B}\rangle\beta,$

• The evolution equation for ϕ now becomes

$$\frac{d}{dt}(\dot{\phi}a^3) = \frac{2\beta^2 \langle \mathbf{B} \cdot \mathbf{B} \rangle}{a\omega} e^{-2\phi}.$$

• By adopting self consistent approximation, to the leading order $t \to \infty$,

$$\alpha \sim \mathcal{A} \ln(t) \exp\left[-\frac{2}{\ln(t)}\right].$$

 α increases logarithmically with time and controlled by the average energy density of the of the radiation and coupling, β .

Conclusions



- For the last many years considerable interest on parity violating extension of standard models leads us to consider parity violating extension of Varying alpha theory.
- Interesting points which is one of our motivations was that the model unifies different cosmic phenomena in a single framework.
- Cosmic birefringence, new non-vanishing multi-pole correlation in CMB and cosmic time variation of α, all these phenomena are unified in a single framework
- In our present study we did not mention anything about the observational constraints on our model parameter. This is our present interest of study.