

Emergence of life in an inflationary universe

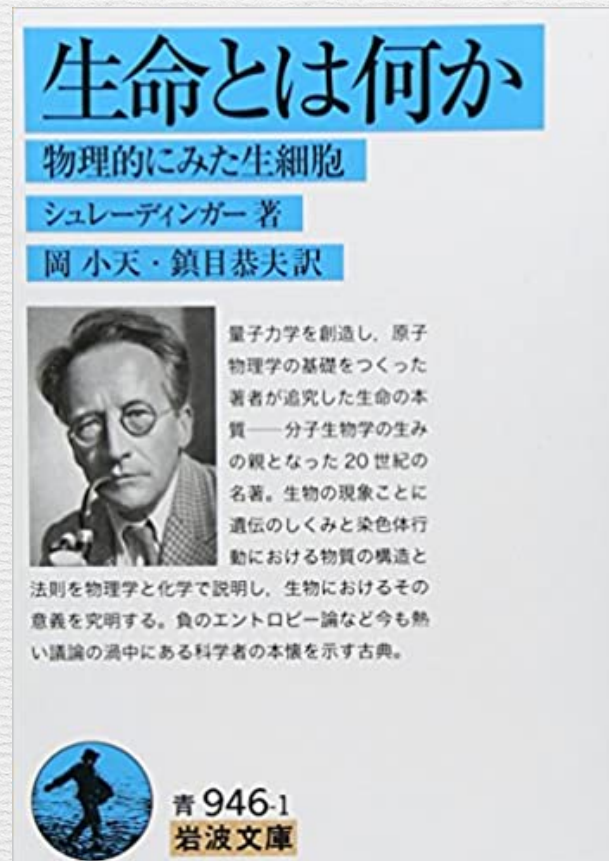
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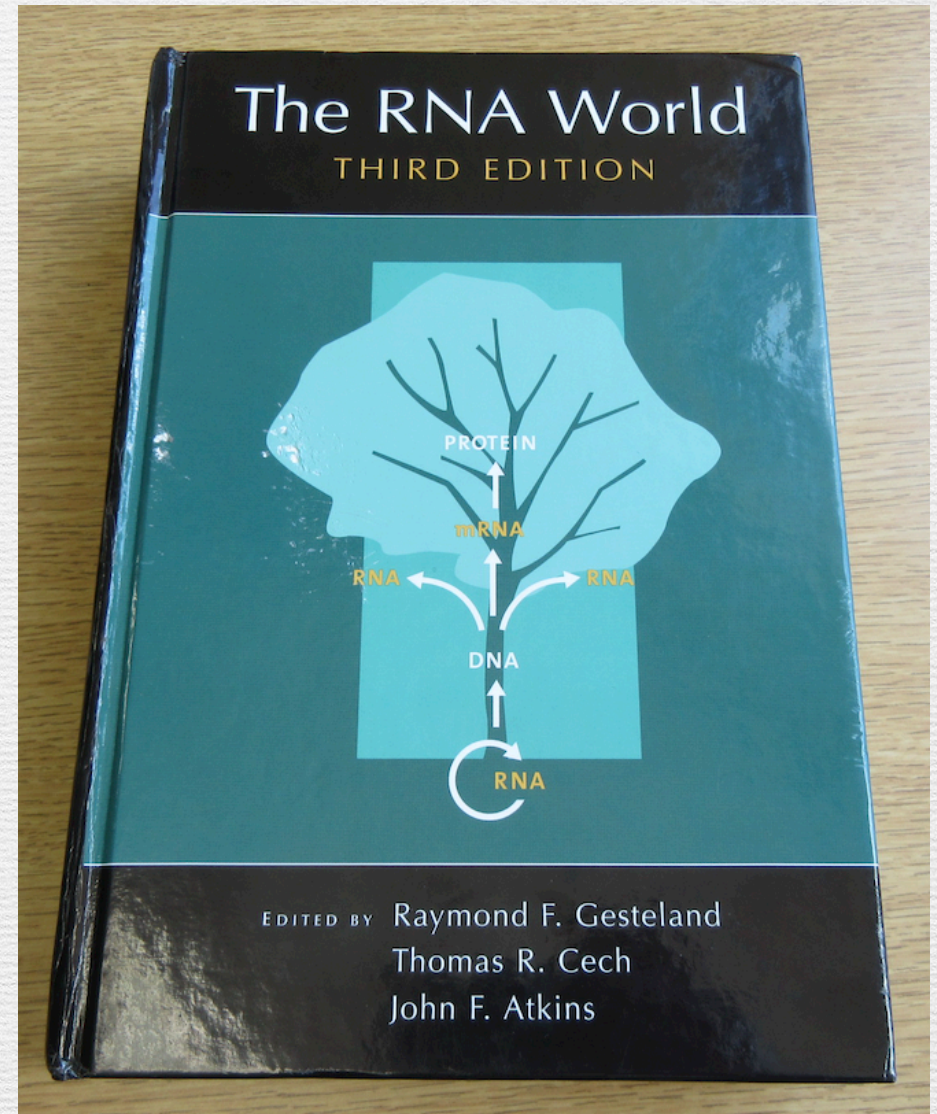
How an astrophysicist started to think about the origin of life?

- Some people say...
 - Sometimes a senior scientist starts a crazy study...
 - When a physicist starts to study about life, it is the end of his/her carrier...
- Some famous people in astronomy/physics started to think about life
 - E. Schrodinger
 - F. Hoyle
 - ... and more?



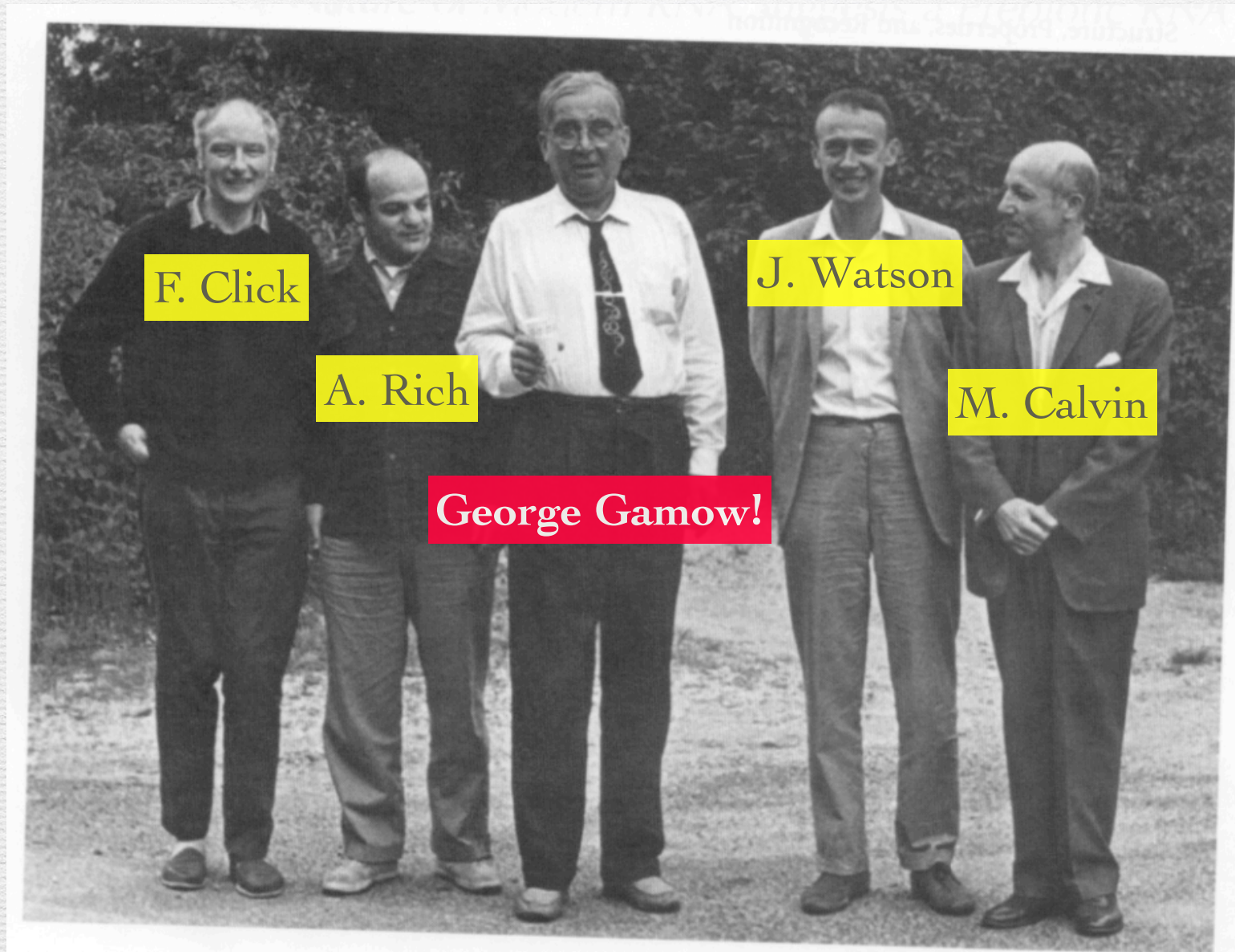
An example of a cosmologist turned into a biologist

- One day, I wanted to read a chapter of a book “the RNA world (3rd edition, 2006)
- Only in the library of School of Medicine in my university!
 - For the first time I visited the medicine library...



An example of a cosmologist turned into a biologist

- On the first page of the book, there is a photo
- Titled “the RNA Tie Club Members”, taken in 1963



An example of a cosmologist turned into a biologist

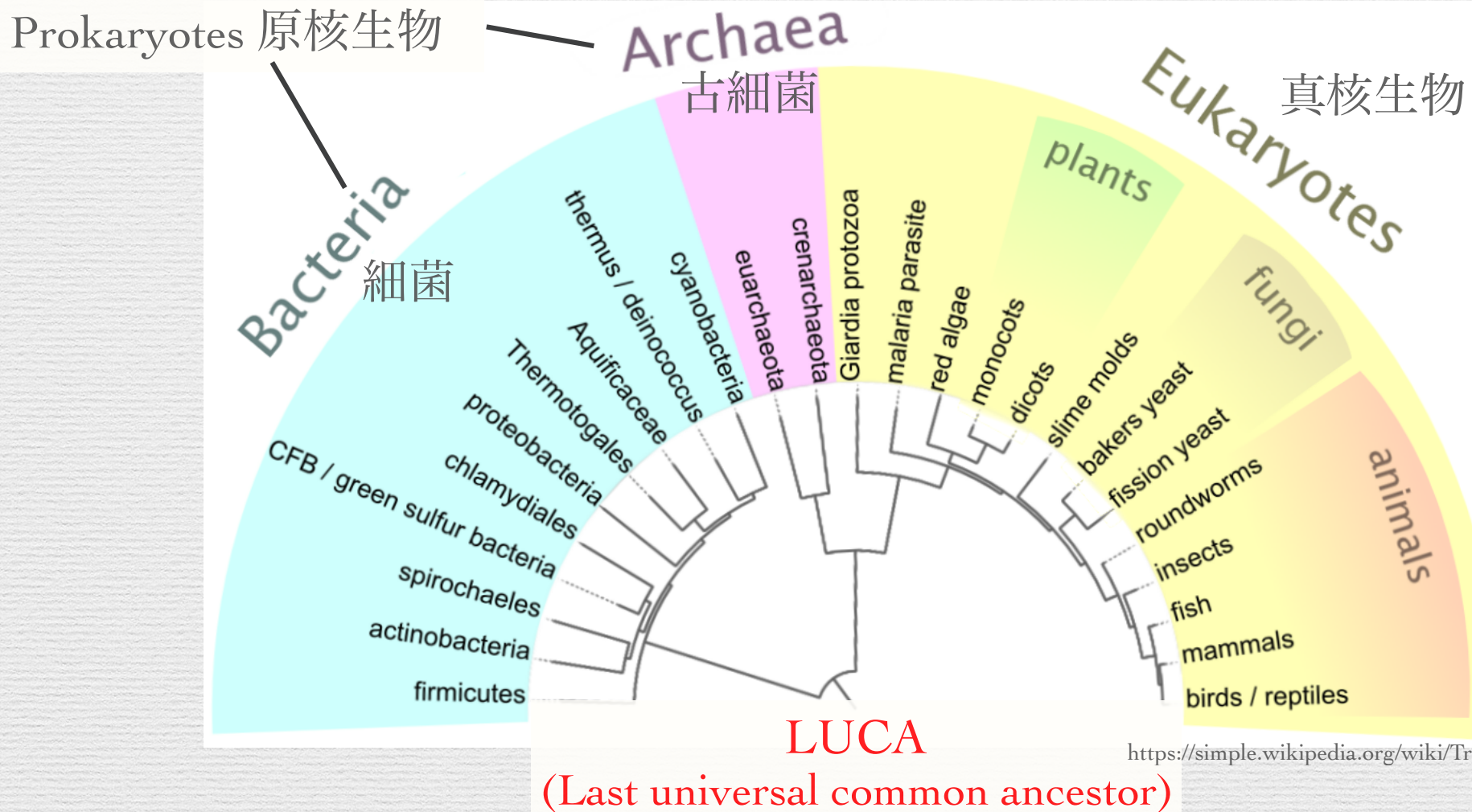
- According to Wikipedia
 - Gamow & Watson co-founded the RNA Tie Club (1954)
 - Crick said that Gamow's suggestions helped him in his own thinking about the protein formation by DNA information
 - 3 nucleotides with 4 different base types ($4^3=64$ combinations) code 20 different amino acids
 - Gamow pointed out: if the order of 3 nucleotides is irrelevant, the different combinations reduce to $(4+3-1)!/3! = 20 =$ types of amino acids!
 - This was incorrect, but contributed to giving rise to models of “biological degeneracy”

The standard genetic code
(codon)

		Second base of codon								
		U	C	A	G					
U	UUU	Phenylalanine phe	UCU	Serine ser	UAU	Tyrosine tyr	UGU	Cysteine cys	U	
	UUC		UCC		UAC		UGC		C	
	UUA	Leucine leu	UCA		UAA	STOP codon	UGA	STOP codon	A	
	UUG		UCG		UAG		UGG	Tryptophan trp	G	
C	CUU	Leucine leu	CCU	Proline pro	CAU	Histidine his	CGU	Arginine arg	U	
	CUC		CCC		CAC		CGC			C
	CUA		CCA		CAA	Glutamine gin	CGA			A
	CUG		CCG		CAG		CGG			G
A	AUU	Isoleucine ile	ACU	Threonine thr	AAU	Asparagine asn	AGU	Serine ser	U	
	AUC		ACC		AAC		AGC		C	
	AUA		ACA		AAA	Lysine lys	AGA	Arginine arg	A	
	AUG	Methionine met (start codon)	ACG		AAG		AGG		G	
G	GUU	Valine val	GCU	Alanine ala	GAU	Aspartic acid asp	GGU	Glycine gly	U	
	GUC		GCC		GAC		GGC			C
	GUA		GCA		GAA	Glutamic acid glu	GGA			A
	GUG		GCG		GAG		GGG			G

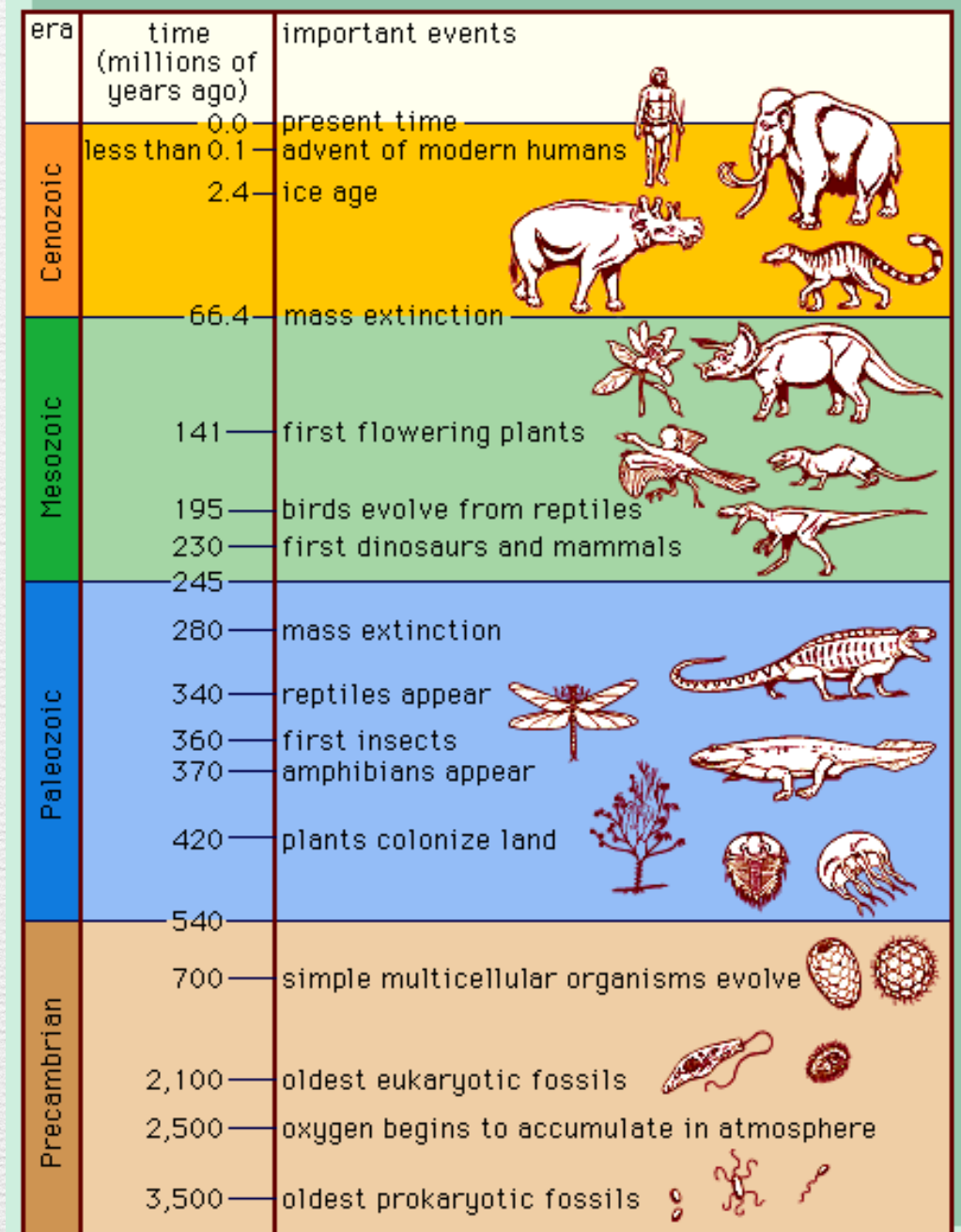
Evolutionary history of life on Earth

- All cellular life on Earth has evolved from the single last universal common ancestor (LUCA)
- No evidence for another independent abiogenesis event
 - “abiogenesis” = emergence of life from abiotic (=non-biotic) processes



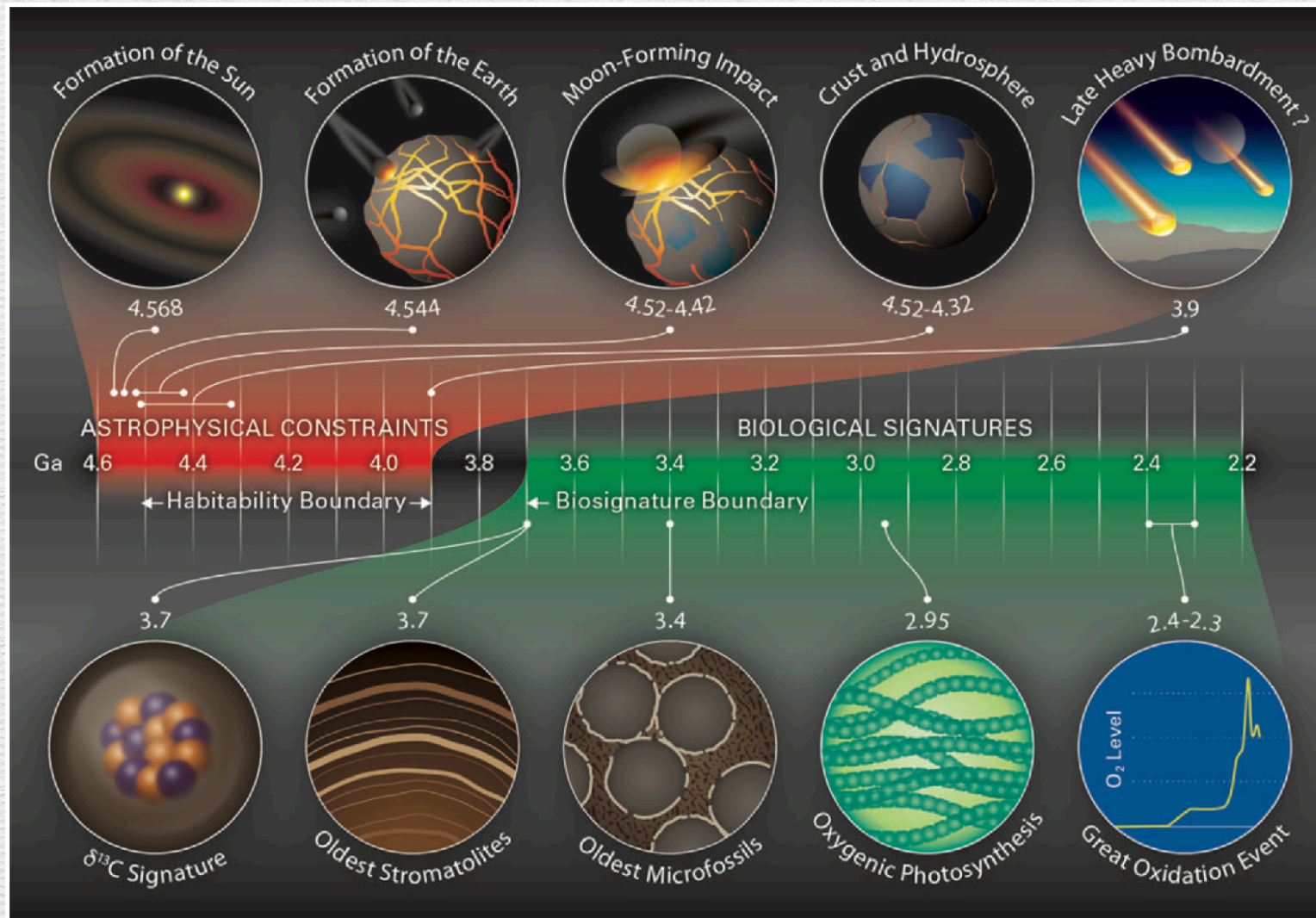
Evolutionary history of life on Earth

- >3.5 Ga (Gyr ago)
 - The first life
- 2.1 Ga
 - The oldest eukaryote fossils
- 0.7 Ga
 - Appearance of multicellular organisms
- 0.5 Ga
 - Increase of oxygen level
 - Plants colonize land
 - Diverse multicellular organisms



When the first life appeared on Earth?

- Older than 3.7 Ga from Fossil records / carbon isotope
- Appearance of the earliest crust & hydrosphere: 4.5 Ga (soon after the Earth formation)
- Life may be possible only after late heavy bombardment (3.9 Ga)
- The time taken for life to appear: 0.2-0.8 Ga (4-18% of the Earth current age)

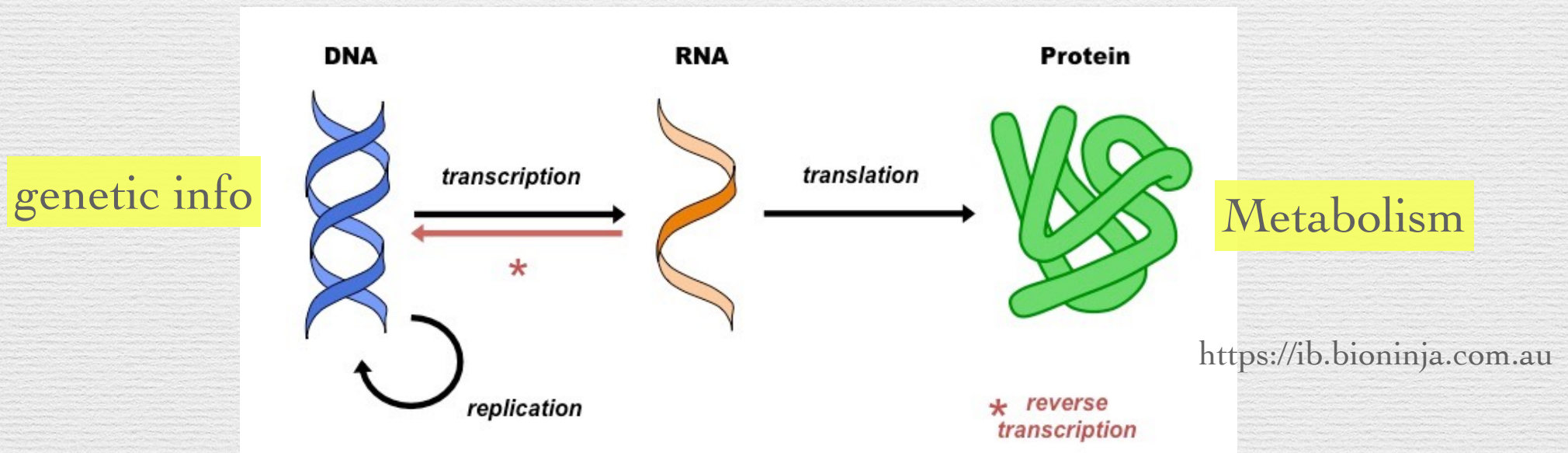


Abiogenesis probability on a planet?

- The early emergence of life in the Earth history implies a high probability?
 - (the only one) argument that favors a high probability
 - An immediate counter argument: no evidence for multiple abiogenesis events
- No strong conclusions can be derived for the moment (Lineweaver+02; Spiegel+12)
 - Early appearance may be just by chance (4-18% chance probability)
 - Early Earth may have been more favorable environment for abiogenesis
 - e.g. higher geological activities, more hydrothermal vents, ...
 - Anthropropic effect: an intelligent life must appear before the end of the Earth habitability
 - Increase of solar luminosity may cause the end of habitability in 1 Gyr?
- Arbitrarily small abiogenesis probability cannot be excluded
 - We do not know any other life of a different origin in Earth or in the universe
 - We will find ourselves on a planet where abiogenesis happened

What is life actually?

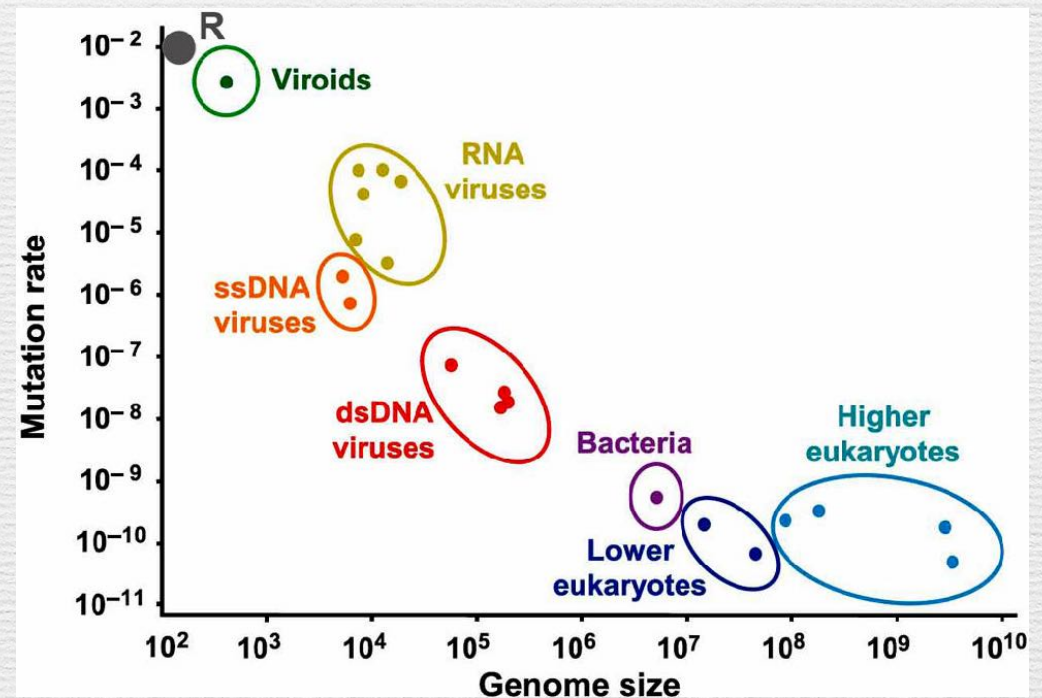
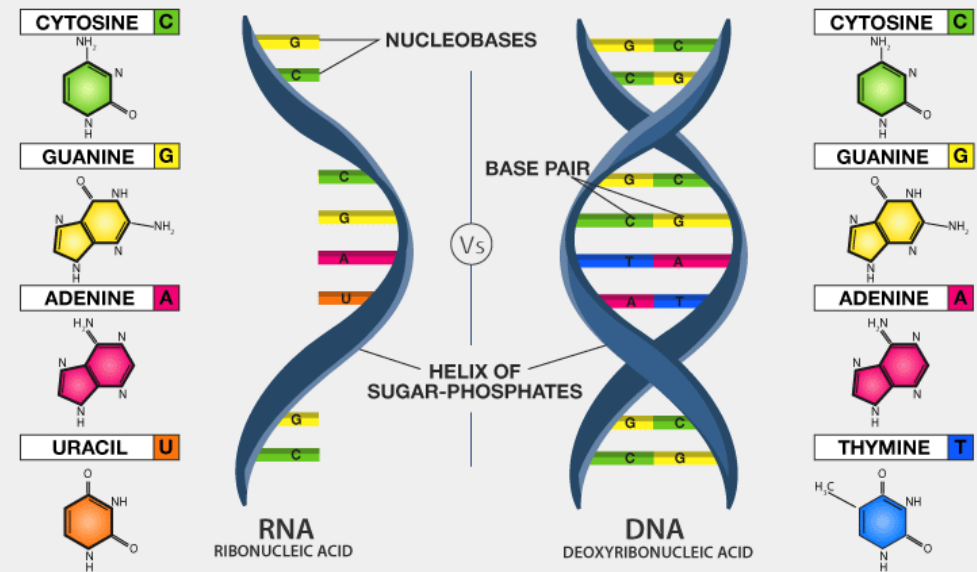
- Definition of life?
 - The NASA Astrobiology Institute Definition of Life: “self-sustaining chemical system capable of Darwinian evolution.”
- Essential aspects of life
 - Self-replication with genetic information
 - Metabolism (take energy from outside for necessary chemical reactions)
 - Compartmentalized by lipid membrane from outside as a cell
- The central dogma of the life on Earth



Genome in Life

- Genome size (in DNA length [base-pair])
 - Polychaos dubium (a sort of amoeba) 6.7×10^{11} (the biggest size in all organisms)
 - Human 3.0×10^9
 - Bacteria/Archaea $1.1 \times 10^5 - 1.6 \times 10^7$
 - Virus $1759 - 2.5 \times 10^6$
 - Viroid 246
 - RNA polymerase ribozyme in in-vitro experiments ~ 100

DIFFERENCE BETWEEN DNA AND RNA



Gago+'09; Ruiz-Mirazo+'14

How did the first life appear?

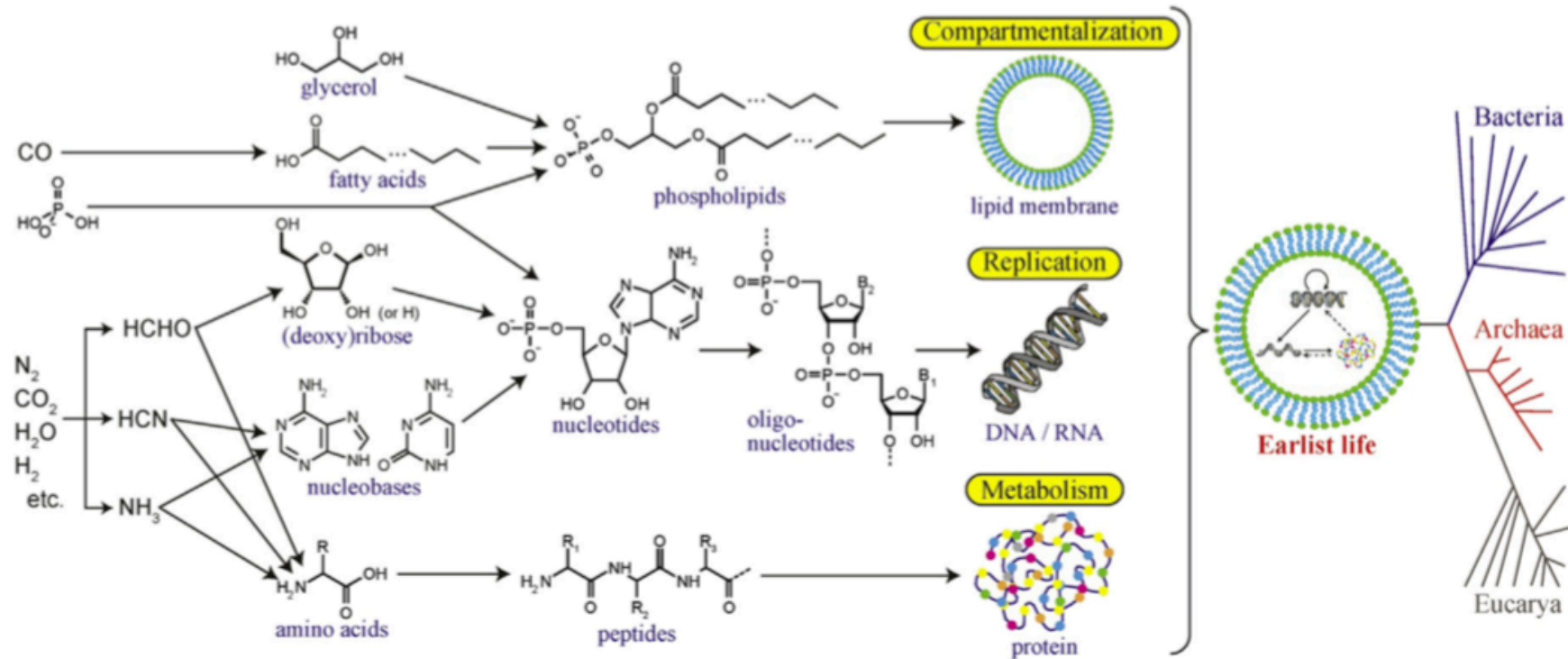
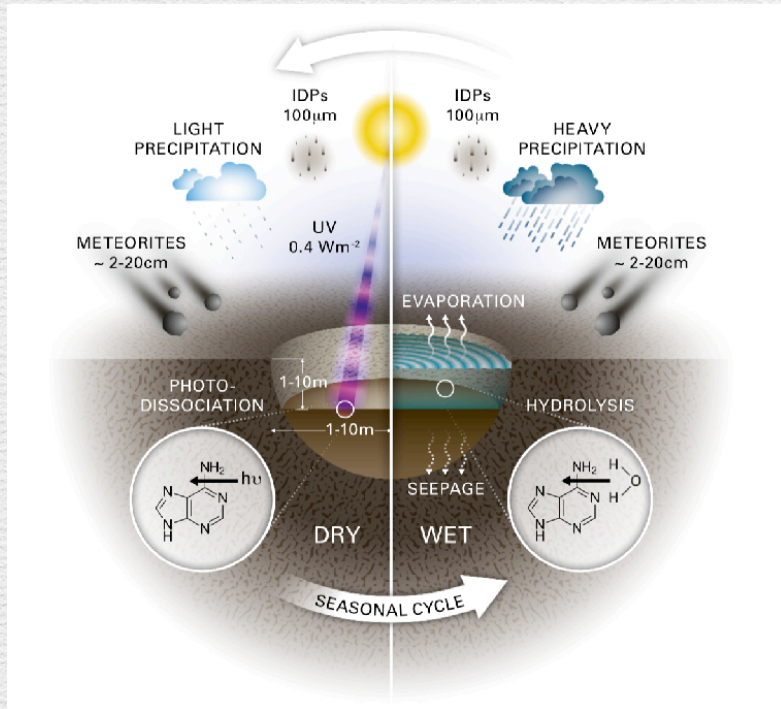


Figure 2. Structures, components, and abiotic synthetic pathways of bio-macromolecules operating three fundamental functions of life (replication, compartmentalization, and metabolism).

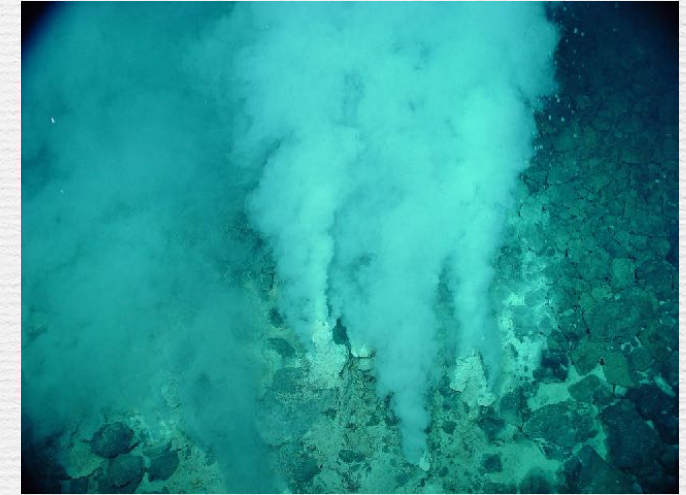
Possible sites of abiogenesis

Warm little ponds?

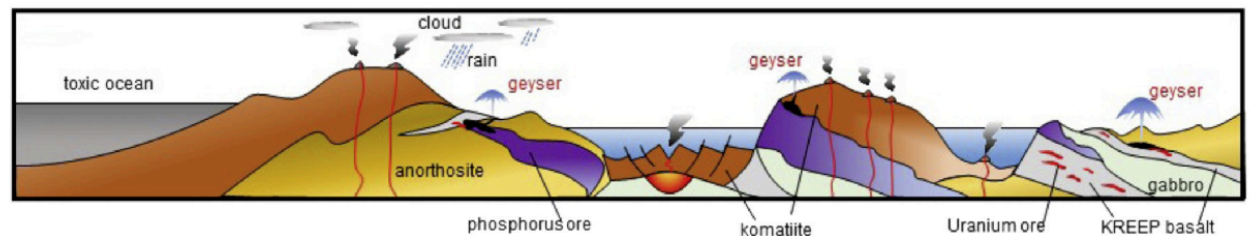
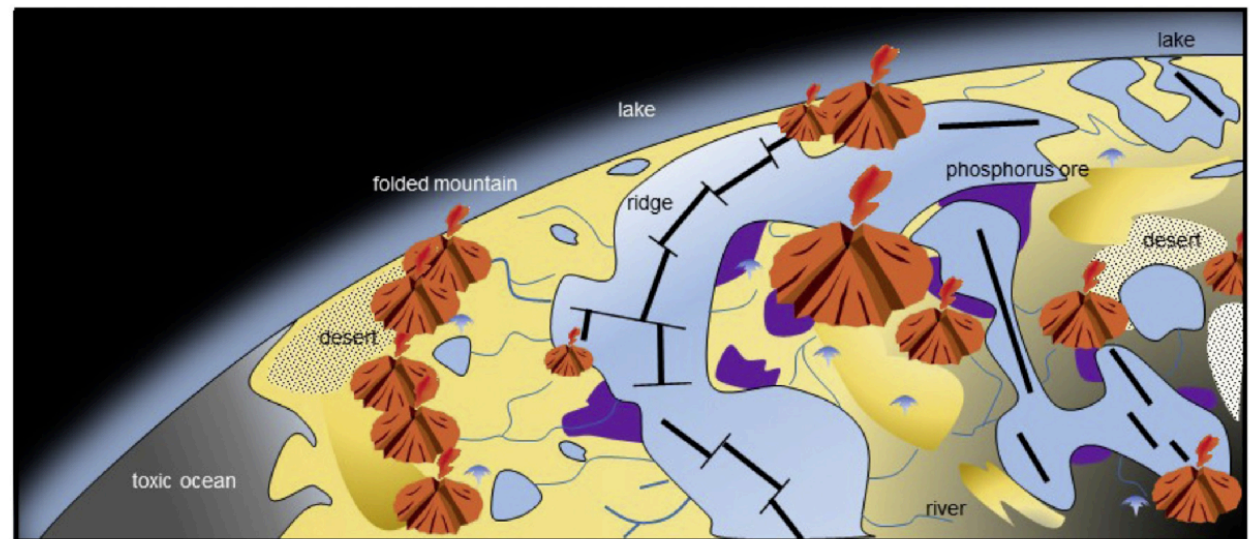


Pearce+'17

Hydrothermal vents?



https://en.wikipedia.org/wiki/Hydrothermal_vent

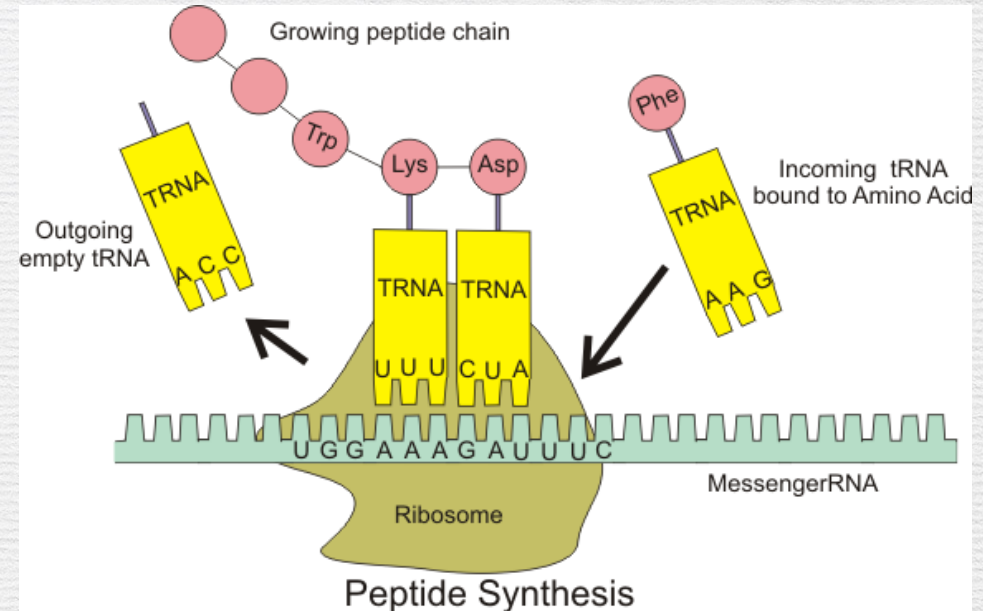


Kitadai+'18

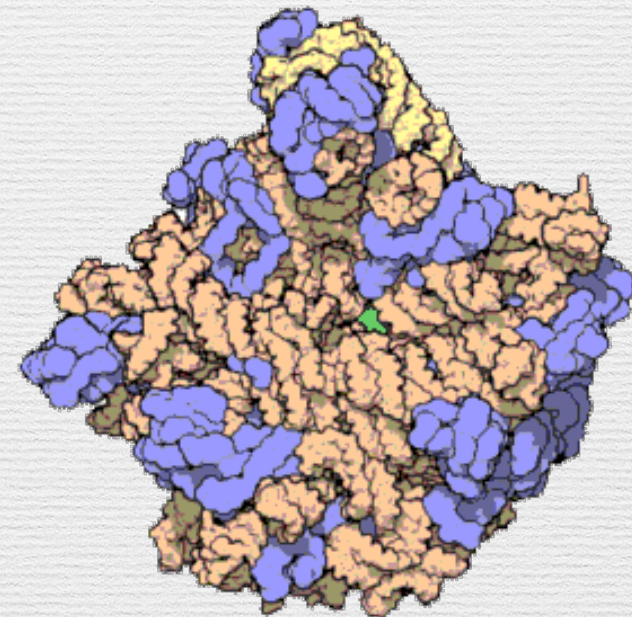
Figure 1. Schematic diagram of a Hadean surface environment proposed by Maruyama et al. (2013) and Santosh et al. (2017).

The RNA world hypothesis

- A chicken-and-egg problem:
 - Protein is assembled based on the genetic information coded in DNA
 - DNA needs protein for their self-replication
- DNA vs. proteins: which was the first?
 - A solution: RNA was the first!
- RNA can catalyze some biological reactions like protein enzymes — “ribozyme”
 - The central part of a ribosome (the machine producing protein from RNA) is made of RNA
- The RNA world hypothesis: self-replicating RNA polymerase ribozyme was the earliest form of life
 - It does not reject a pre-RNA world with different types of molecules
- The most popular hypothesis in the origin-of-life studies

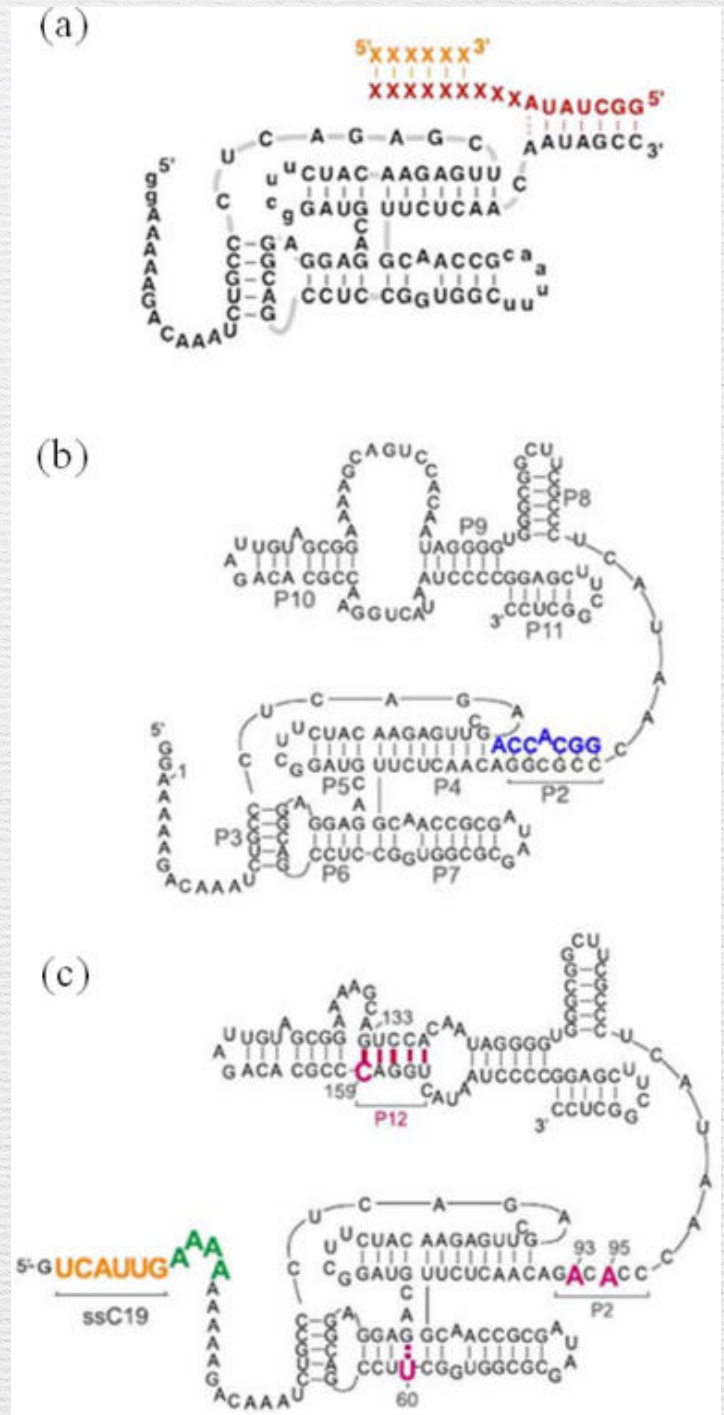


<https://en.wikipedia.org/wiki/Ribosome>



What did the first RNA polymer look like?

- The RNA world hypothesis does not give any explanation of the first self-replicating RNA polymer
- *In vitro* evolution experiments have produced ribozymes with template-dependent RNA polymerase activity
 - Typically 100-200 nt (nucleotides)
- How short can the first self-replicating RNA be? According to expert biologists (Szostak '93; Robertson+'12):
 - RNA shorter than 25-nt do not show a specified function
 - A reasonable hope of RNA replicase ribozyme with a length of 40-60 nt



The first key step to abiogenesis

- A polymer of RNA (or its precursor) longer than 40-100 nt needs to be produced abiotically, with a particular nucleotide sequence (genetic info) allowing a self-replicating activity
- Not just one, but a pair of identical RNA strand may be required
 - One acts as a ribozyme, and the other as a template
 - Then the required length would be twice longer
- The first big step for abiogenesis
 - Generation of biological information as an ordered sequence from zero
 - Statistically the biggest barrier (in my view)
- After that, the Darwinian evolution is triggered toward the first cell
 - Encapsulation by membrane
 - Incorporating protein and DNA

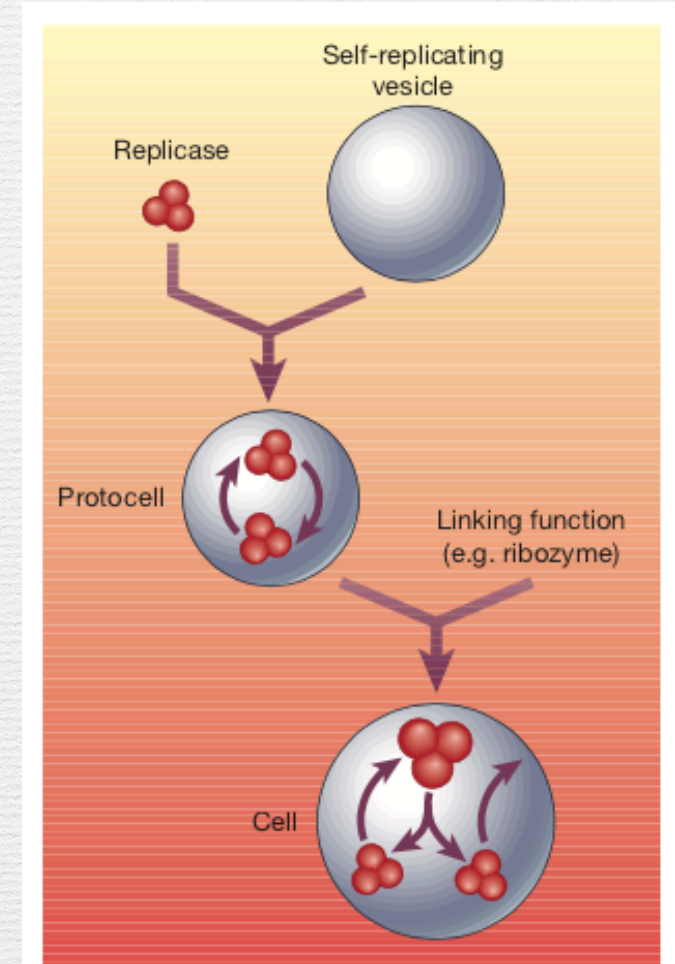


Figure 2 Outline of proposed pathway for synthesis of a cell. The first major synthetic intermediates are an RNA replicase and a self-replicating vesicle. These are combined into a protocell, enabling rapid evolutionary optimization of the replicase. Addition of an RNA-coded linking function, such as a lipid-synthesizing ribozyme, completes the cellular structure.

Szostak '01

How was the first active RNA built abiotically?

- RNA polymerization experiments
 - polymerization occurs by reactions of activated monomers
 - can be accelerated on catalyzer (e.g. clay mineral surface) (e.g. Ferris+'06)
 - consistent with statistical adding of monomers randomly
 - Yields rapidly decrease with polymer length, up to ~10 nt (Kawamura+'94, Cafferty+14)
- Some reports reporting >30-nt polymer production remain controversial, possibly misidentifying polymers and aggregates
 - Huang+03, Costanzo+09, Burcar+13

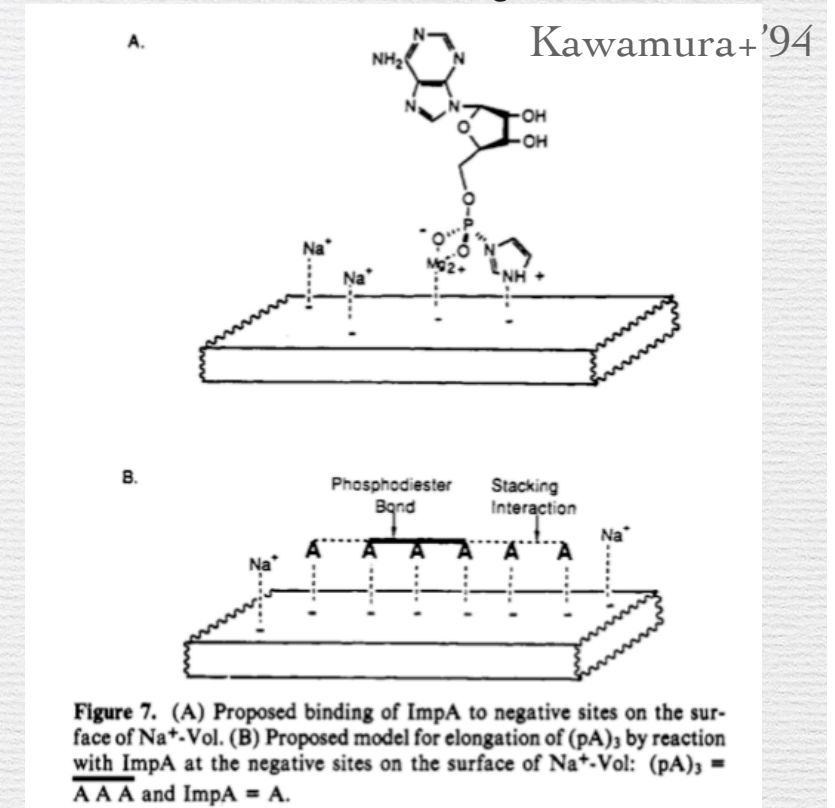
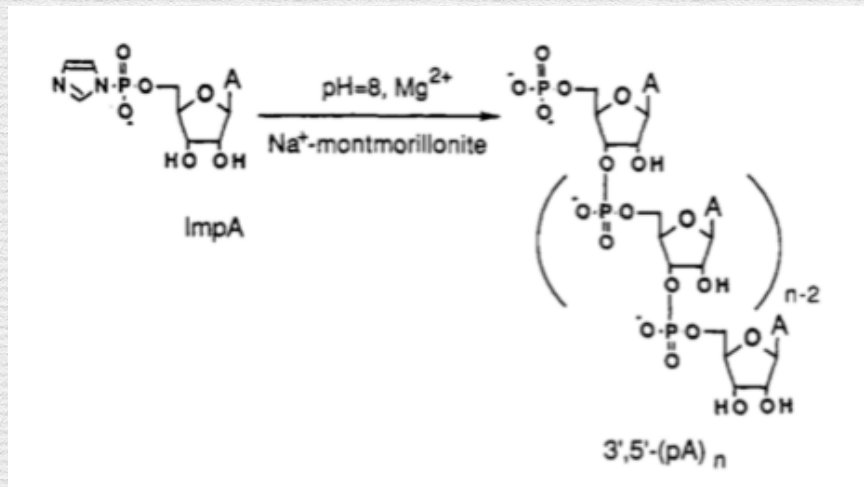


Figure 7. (A) Proposed binding of ImpA to negative sites on the surface of Na⁺-Vol. (B) Proposed model for elongation of (pA)₃ by reaction with ImpA at the negative sites on the surface of Na⁺-Vol: (pA)₃ = A A A and ImpA = A.

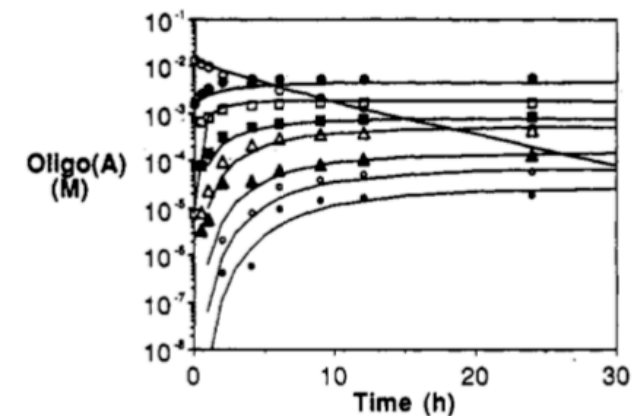
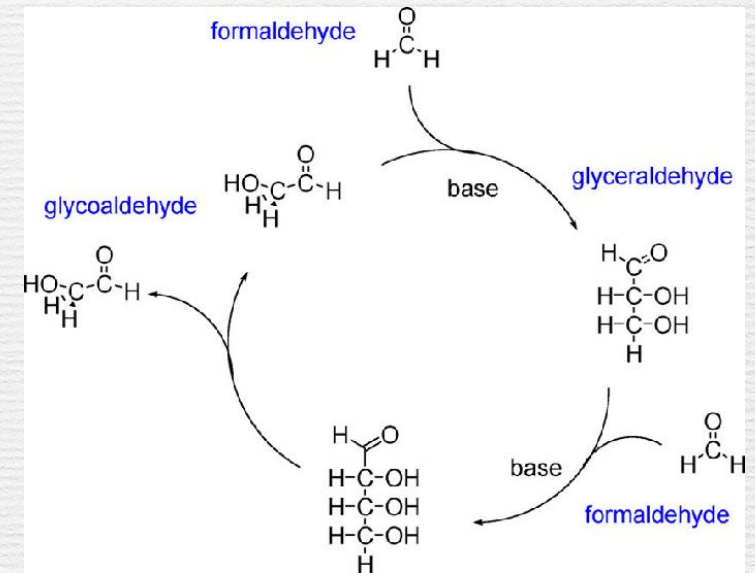


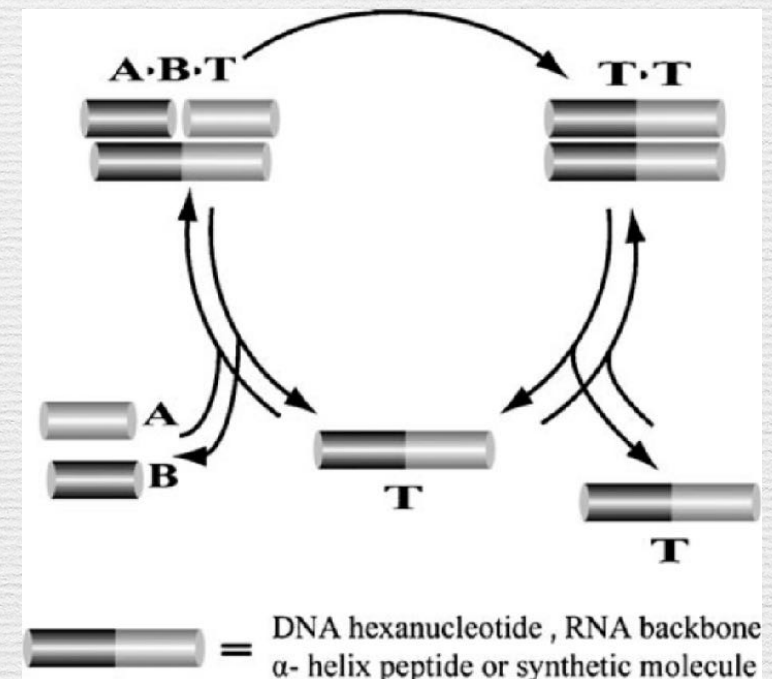
Figure 5. Reaction curve for the loss of 15 mM ImpA and the formation of oligo(A)s in the presence of Na⁺-Vol. Total concentrations of the oligomer in the aqueous phase and bound to the Na⁺-Vol are plotted. The lines drawn through the experimental points were fit by SIMFIT: ○, ImpA; ●, pA; □, 2-mer; ■, 3-mer; △, 4-mer; ▲, 5-mer; ○, 6-mer; ●, 7-mer.

More efficient processes to the first RNA?

- More efficient processes than randomly adding monomers?
 - Cyclic (auto- or cross-catalytic) reaction network has been considered in the context of abiogenesis (Coveney+'12; Hordijk+'18)
 - Mostly theoretical, but some experimental demonstrations for basic reactions
 - e.g. template-directed ligation of short oligomer fragments
 - No experimental proof for such reactions actually working at the abiogenesis
- Such a reaction network may be useful for self-replication of necessary molecules (building blocks), but unlikely useful for template-independent polymerization of the first long polymer at the starting point of abiogenesis
- Conservatively, the simplest polymerization by adding monomers is the only one path that certainly exists



Ruiz-Mirazo+'14



Dadon+'08

How small can the abiogenesis probability be?

- Abiogenesis by statistically adding monomers looks extremely unlikely
 - Can a monkey hitting a keyboard at random type a complete work of Shakespeare?
 - “The probability of abiogenesis is comparable to the chance that a tornado sweeping through a junkyard might assemble a Boeing 747” (F. Hoyle)



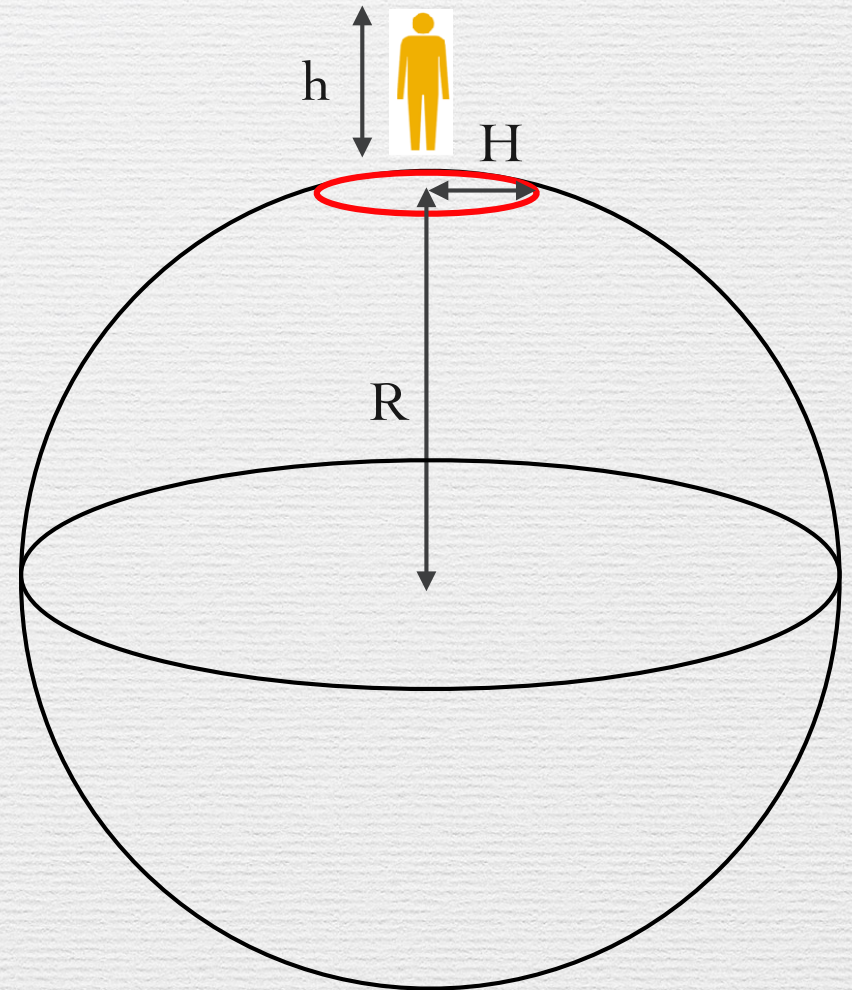
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- In literature, a small abiogenesis probability is often considered as that in the whole Galaxy, or in the observable universe (13.8 billion light year radius including 10^{22} Sun-like stars and 10^{11} galaxies)
- From a viewpoint of cosmology, considering abiogenesis probability in the observable universe is almost nonsense!
 - The size of the observable universe is determined by the light travel distance since the Big Bang
 - “the cosmic horizon” in which we can directly see

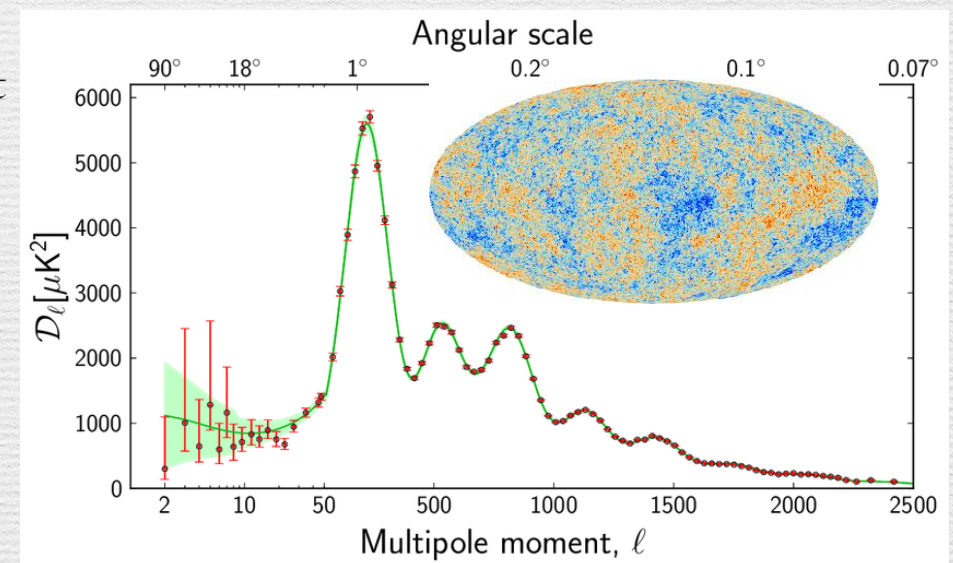
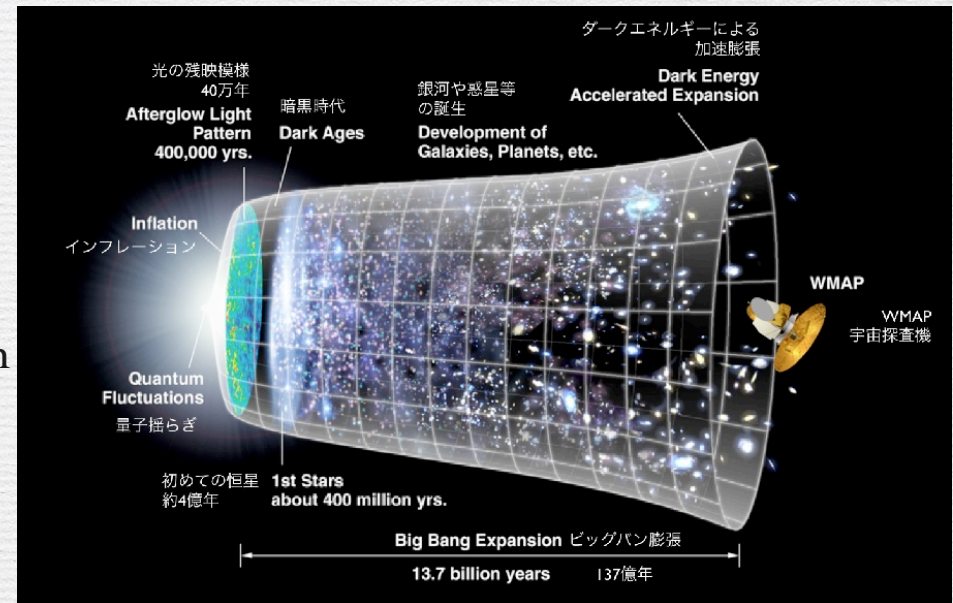
The horizon on the Earth surface

- Horizon length for a person (height h) is:
 - $H = [2 h (R + h)]^{1/2}$
 - R =Earth radius= 6371 km
 - $H \sim 5$ km for a $h=1.7$ m person
- Suppose you are considering on the abiogenesis probability on Earth. How do you feel if somebody considers about the probability within one's horizon
- “abiogenesis probability in the observable universe” is similar to this
- Then, how large is the real physical universe?



The inflationary universe

- The observable universe now is
 - extremely homogeneous
 - spatially flatfar beyond the horizon at the early universe
- The widely accepted solution: cosmic inflation
 - The universe experienced an exponential expansion in the early universe (typically $t < 10^{-37}$ s, $kT > 10^{16}$ GeV)
 - Such an expansion provides a flat and homogeneous universe beyond the horizon
 - The minimum expansion factor required:
 - $\exp(N_i)$ with $N_i > 60$
 - 3×10^{-27} cm \rightarrow ~ 1 cm (our observable universe at the end of inflation)
- Further observational support:
 - Scale-invariant matter density fluctuations in the universe (power spectrum $\propto 1/\lambda$)
 - Confirmed at a high precision by the cosmic microwave background radiation



How large is the inflationary universe?

- The minimum requirement to solve the homogeneity/flatness problems:
 - $N_i > N_{i,\min} \sim 60$
 - No observational upper bound on N_i
 - $N_i \propto$ inflation duration
- If you change some model parameters of inflation, the duration easily changes, say, by a factor of a few
- It would be a fine tuning if the actual N_i is very close to $N_{i,\min}$
 - Rather, we expect $(N_i - N_{i,\min}) \gg N_{i,\min}$
- If inflation duration is twice, three, and four times larger than $N_{i,\min} \sim 60$:
 - The physical size of the universe is $\exp(60)$, $\exp(120)$, and $\exp(180)$ times larger than the observable universe now
 - 10^{78} , 10^{156} , and 10^{234} times larger in volume
 - contains 10^{100} , 10^{178} , 10^{256} Sun-like stars
- If abiogenesis occurs many times in an inflationary universe, it is consistent with our existence even if the abiogenesis probability is negligibly small in the observable universe (10^{22} stars)

This work

T. Totani 2020, Scientific Reports 10, 1671

- Examine whether life can emerge by the most conservative process (random reactions of monomers) in the total volume of an inflationary universe
 - Derive a quantitative relation between:
 - The minimum length of RNA (or its precursor) required to trigger an abiogenesis event
 - The universe size necessary to expect the formation of such a long RNA by random adding of monomers
- Koonin (2008) considered abiogenesis in the universe assuming “the eternal inflation scenario”
 - In this scenario, multi-verses are produced by repeated inflation events
 - Predicts a infinite volume of the universe, thus abiogenesis is possible with an infinitely small probability
 - Quantitative discussion impossible
 - Eternal inflation is difficult to confirm observationally
- This work assumes the minimal inflation to explain observations:
 - A homogeneous region created by a single inflation event, including our observable universe
 - Thus a quantitative discussion becomes possible

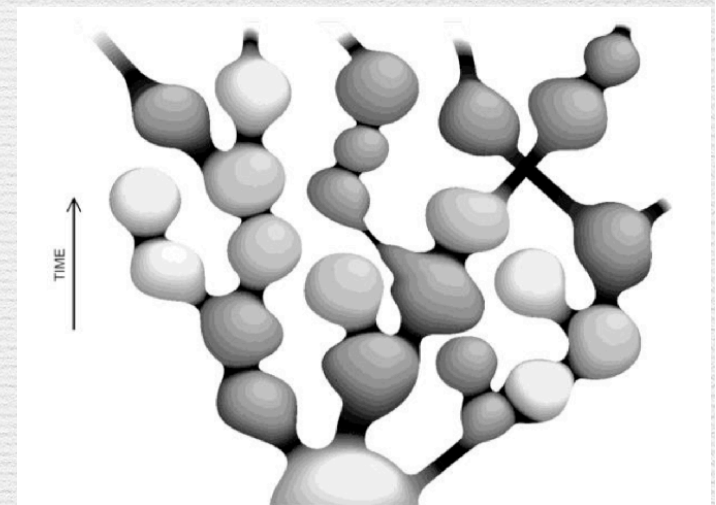


Fig. 4 - Global structure of the chaotic self-replicating multiverse. The different gray hues symbolize “mutations” in the physics laws in the domains, as compared to the parent island universes.
Adapted from Linde (2014; 2015)

Poissonian RNA polymerization

- Suppose a cycle of polymerization, where activated monomers are injected into a reaction site
- Polymerization by adding monomers
 - Basic equations for the abundance of l-length RNA, x_l $\dot{x}_{l+1} = \kappa x_l - \kappa x_{l+1}$
 - κ : the reaction probability per unit time
 - solving with a condition of $x_l \ll x_{l+1}$, abundance at a time t is:
 - $p_r = \kappa t$
 - can also be derived from the Poisson distribution
- We only consider the regime of $p_r = \kappa t < 1$
 - Beyond $t > 1/\kappa$, a significant amount of activated monomers would be lost, so the above simple formula does not hold
- $p_r \sim 1$ can be achieved if polymerization in a cycle continues up to $t \sim 1/\kappa$
 - Experiments on clay surface found $p_r \sim 1$ is achieved with $\kappa \sim (\text{hour})^{-1}$
 - The reaction rate $\kappa \propto (\text{monomer concentration})$
 - High monomer concentration necessary. Dry-wet cycle in warm little ponds?

Equation of abiogenesis probability

- The expected number of abiogenesis events in a region including N_* stars

$$N_{\text{life}} = N_* f_{\text{pl}} t_d r_p P_{\text{ac}} P_{\text{ev}}$$

- f_{pl} : mean number of habitable planets per star
- t_d : duration of polymerization cycles continue
- r_p : production rate of a l_{min} -length RNA polymer on a planet

- N_m : the number of activated monomers on a planet
- t_c : time interval between cycles

$$r_p = N_m \frac{p_r^{l_{\text{min}}}}{l_{\text{min}}!} t_c^{-1}$$

- P_{ac} : probability of active nucleobase sequences among all sequences
 - N_{ac} : number of active sequences, N_{nb} (=4): number of nucleobase types

$$P_{\text{ac}} = \frac{N_{\text{ac}}}{N_{\text{nb}}^{l_{\text{min}}}} = \frac{1}{N_{\text{nb}}^{l_{\text{min}} - \Delta l}}$$

$$N_{\text{ac}} \equiv N_{\text{nb}}^{\Delta l}$$

- N_{ac} is uncertain, but perhaps 10^4 sequences among all sequences of 40-mer ($4^{40} \sim 10^{24}$ sequences), i.e. $\Delta l = 6.6$ for $l_{\text{min}} = 40$ (Robertson & Joyce '12)
 - Assume $\Delta l = 0$ in the baseline model
- P_{ev} : probability that an active RNA polymer evolves to life (=1)
 - very uncertain, can be much smaller

lg N* vs. l_{min}

- Requiring N_{life} = 1, and taking a logarithm

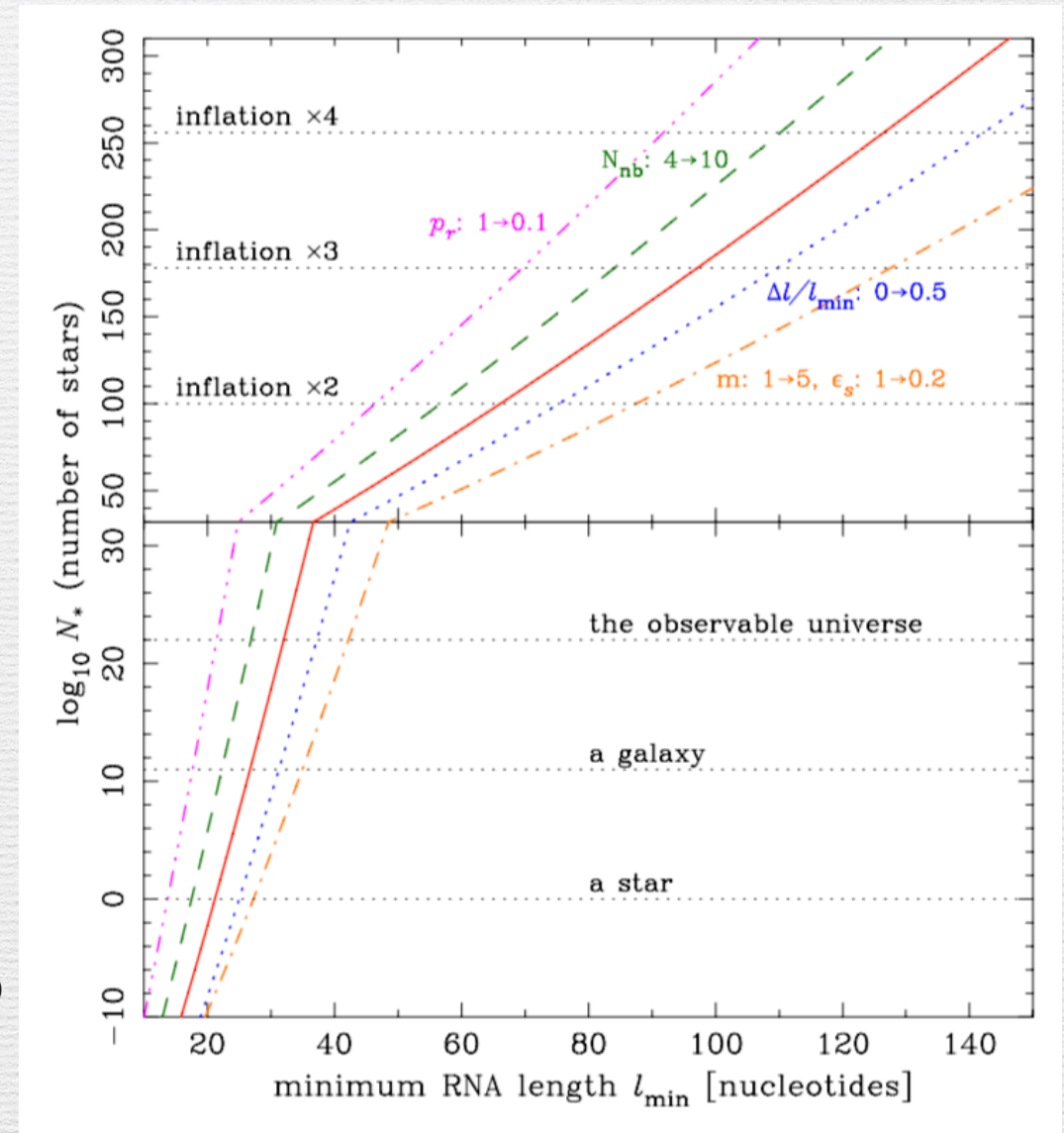
$$2.3 \lg N_* = \ln(l_{\min}!) - l_{\min} \ln p_r + (l_{\min} - \Delta l) \ln N_{nb} - \ln C$$

$$C \equiv f_{pl} N_m t_d t_c^{-1} P_{ev}$$

- Non-logarithmic main terms: lg N* ($\propto N_i$) and l_{min}
- Other parameters appear only logarithmically, though some of them are very uncertain
 - C is uncertain probably by more than 10 orders of magnitude
- Baseline estimate of C:
 - f_{pl} = 0.1 (from exoplanet observations)
 - t_d = 0.5 Gyr (time scale from Earth formation to abiogenesis)
 - t_c = 1 yr (assuming a seasonable cycle, 1 day may also be good for day-night cycle)
 - N_m = 10²⁵ (~ 1 kg) (very uncertain, from a model of nucleobase delivery by meteorites)
 - N_m ~ 10³⁸ for the total biomass on the present Earth
 - ln C = 75.3

Results

- l_{\min} vs $\lg N_*$
 - $l_{\min} = 21$ for $\lg N_* = 0$ (a star)
 - $l_{\min} = 27$ for $\lg N_* = 11$ (a galaxy)
 - $l_{\min} = 31$ for $\lg N_* = 22$ (the observable universe)
- c.f. $l_{\min} > 40-100$ required for a biological self-replicating activity
- $\lg N_* = 39$ for $l_{\min} = 40$
- Inflationary universe:
 - $l_{\min} = 66, 97, 127$ for $\lg N_* = 100, 178, 256$ ($\times 2, \times 3, \times 4$ longer inflation than the minimum)
- $\lg N_*$ changes by 10 when C is changed by $\times 10^{10}$
 - Not a large effect especially in the large limit of l_{\min} and $\lg N_*$



Discussion & Conclusions (1/2)

- A long (>40-100nt) and biologically active RNA can be produced in the total volume of an inflationary universe only by random reactions of adding monomers
 - A (currently the only) solution to the initial abiotic polymerization problem about the origin of life
 - Life can emerge in the framework of science, and we do not have to consider something beyond science (e.g. intelligent design)
- The expected number of abiogenesis events is small even in the observable universe, but it does not contradict with available observational facts
 - The observable universe is just a tiny part (likely less than $1/10^{78}$) of the real universe size
 - No strong reason to expect more than one abiogenesis events in the observable universe
 - Giving an explanation to the Fermi's paradox
- Implication for the homochirality of life:
 - Life uses only L-chiral (left-handed) amino acids and D-chiral (right-handed) sugars in DNA/RNA
 - The I_{\min} vs $\lg N^*$ statistics does not significantly change when N_{nb} (base types) increased by $\times 2$
 - If homochirality is a necessary condition for a biological activity, simply a bit larger volume of inflationary universe is required to expect an abiogenesis event

Discussion & Conclusions (2/2)

- Implications for extraterrestrial life search:
 - I regret that, but I bet on negative results in future searches in the Solar system or nearby exoplanets!
 - If discovered, I would first consider the possibility of interplanetary/interstellar traveling of microorganisms (i.e., the common origin with the life on Earth)
 - Caveat: a much more efficient process of RNA polymerization may change this conclusion, but in my view, no such convincing process is known yet
 - ... and we do not need such a hypothetical process to explain the origin of life
- Ultimate mystery of life?
 - I assumed that life can emerge if a sufficiently long RNA polymer is abiotically assembled
 - This may be trivial under the physical laws that we know, because we know the phenomenon of life (esp. ribozyme) on Earth
 - But does it easily occur when the physical laws are arbitrarily made?
 - This may be the ultimate mystery about life, which is not answered at all by this work