# Patterns in Multiplanet Systems as Fossils of Planet Formation 

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\begin{aligned}
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& 0.99 \\
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Exoplanets Beyond the Snowline - University of Tokyo

## Questions we'd like to answer

- What properties are common in multiplanet systems?
- What do common properties teach us about planet formation?

What properties are common in multiplanet systems?

California-Kepler Survey
New exoplanet science from Keck/HIRES Spectra of 1305 Kepler Planet-hosting Stars


## The California Kepler Survey yielded precise parameters for

909 transiting planets in 355 multiplanet systems


Do you see any patterns?

| K02220 | $0-0-0$ |  |  | 1.03 |
| :---: | :---: | :---: | :---: | :---: |
| K00939 | $\bigcirc \bigcirc$ |  |  | 1.03 |
| K01860 | $\bigcirc 000$ |  |  | 1.01 |
| K02433 |  | - |  | 1.0 |
| SOL |  | - - . | - | 1.0 |
| K01432 |  |  |  | 0.99 |
| K00157 | 00000 | $\bigcirc$ |  | 0.99 |
| K00232 | - 000 |  |  | 0.98 |
| K00408 | - 000 |  |  | 0.98 |
| K00116 | - 0 - 0 |  |  | 0.96 |
| K00935 | - | $\bigcirc$ |  | 0.96 |
| K01052 | $\bigcirc 000$ |  |  | 0.96 |
| K02169 | $\bigcirc$ |  |  | 0.94 |
| K00435 | $\bigcirc \bigcirc \bigcirc \bigcirc$ |  |  | 0.94 |
| K00070 | $\bigcirc \circ 0-0$ |  |  | 0.94 |
| K00880 | $\bigcirc \bigcirc \bigcirc$ |  |  | 0.93 |
| K04032 | - ○ oo |  |  | 0.92 |
| K00841 | $\bigcirc \bigcirc$ | - |  | 0.89 |
| K00907 | - 0 | - |  | 0.88 |
| K00720 | - 0 O |  |  | 0.87 |
| K00623 | -0-0 |  |  | 0.86 |
| K00869 | $\bigcirc 000$ |  |  | 0.86 |
| K00520 | - 000 |  |  | 0.86 |
| K00505 | - 000 | $\bigcirc$ |  | 0.85 |
| K01278 | - 0 - |  |  | 0.85 |
| K01364 | $\bigcirc \bigcirc \bigcirc 00$ |  |  | 0.85 |
| K01563 | $\bigcirc \bigcirc \bigcirc \bigcirc$ |  |  | 0.84 |
| K01151 | $\bigcirc \circ \bigcirc 00$ |  |  | 0.84 |

Do you see any

Planets in the same system have similar sizes

| K02220 | 000 |  |  | 1.03 |
| :---: | :---: | :---: | :---: | :---: |
| K00939 | 0000 |  |  | 1.03 |
| K01860 | $\bigcirc 000$ |  |  | 1.01 |
| K02433 | - 00 | $\bigcirc$ |  | 1.0 |
| SOL |  | - - | - | 1.0 |
| K01432 | $\bigcirc 00$ |  |  | 0.99 |
| K00157 | 00000 | $\bigcirc$ |  | 0.99 |
| K00232 | - 0 - 0 |  |  | 0.98 |
| K00408 | - 000 |  |  | 0.98 |
| K00116 | - 0 - 0 |  |  | 0.96 |
| K00935 | - | $\bigcirc$ |  | 0.96 |
| K01052 | $\bigcirc \bigcirc \bigcirc 0$ |  |  | 0.96 |
| K02169 | 0001 |  |  | 0.94 |
| K00435 | $\bigcirc \bigcirc \bigcirc \bigcirc$ |  |  | 0.94 |
| K00070 | $\bigcirc \circ 0$ - |  |  | 0.94 |
| K00880 | $\bigcirc \bigcirc \bigcirc$ |  |  | 0.93 |
| K04032 | 0000 |  |  | 0.92 |
| K00841 | $\bigcirc \bigcirc$ | - |  | 0.89 |
| K00907 | - 0 | - |  | 0.88 |
| K00720 | - 000 |  |  | 0.87 |
| K00623 | 000 |  |  | 0.86 |
| K00869 | $\bigcirc 000$ |  |  | 0.86 |
| K00520 | - 000 |  |  | 0.86 |
| K00505 | $\bigcirc 0$ | $\bigcirc$ |  | 0.85 |
| K01278 | $0-0$ |  |  | 0.85 |
| K01364 | 0001 |  |  | 0.85 |
| K01563 | $\bigcirc \bigcirc \bigcirc \bigcirc$ |  |  | 0.84 |
| K01151 | - 0 O 0 |  |  | 0.84 |



The sizes of pairs of planets in the same system are correlated.


## Test Null Hypothesis with Bootstrap Trials

Observed system:


Possible bootstrap system:


Star, number of planets, orbital periods are preserved Planet size is drawn at random
Only detectable planets are counted

One example bootstrap trial: no correlation between planet sizes


The sizes of pairs of planets in the same system are correlated.


1000 bootstrap trials: the planet size correlation is not reproduced with a null hypothesis + detection biases




The orbital period ratios of planets in the same system are correlated.


One example bootstrap trial: no correlation between planet spacings


1000 bootstrap trials: the period ratio correlation is not reproduced with a null hypothesis + detection biases



## The spacing and size of a pair of planets are correlated

Is size-spacing pattern related to mutual Hill radii?


Spacing of Pair

## Estimated mutual Hill radii for all planet pairs




Masses estimated with

- Weiss \& Marcy 2014
- Weiss et al. 2013

Weiss+ (in review)

## Do mutual Hill radii affect stability?

10-50 mutual Hill radii apart

10-50 mutual Hill radii apart
>20 mutual Hill radii apart

Hill instability does not affect the smallest planets


Weiss+ (in review)

## How close can two planets be?



Weiss+ (in review)

What do these patterns teach us about planet formation?

## Theories of oligarchic growth

Lissauer \& Stewart (1993):
The self-limiting nature of runaway growth strongly implies that massive protoplanets form at regular intervals in semimajor axis.

Kokuba \& Ida (1998):
We have shown the oligarchic growth of protoplanets in the post-runaway stage. Protoplanets with the same order masses with the orbital separation larger than about $5 r_{\mathrm{H}}$ is the inevitable outcome of planetary accretion in the post-


Kepler multis = aged oligarchs?

Planets in the same system have similar sizes and masses


## Questions we have addressed

- What properties are common in multi-planet systems?

- What do common properties teach us about planet formation?

Similar masses \& separations were predicted in oligarchic growth

