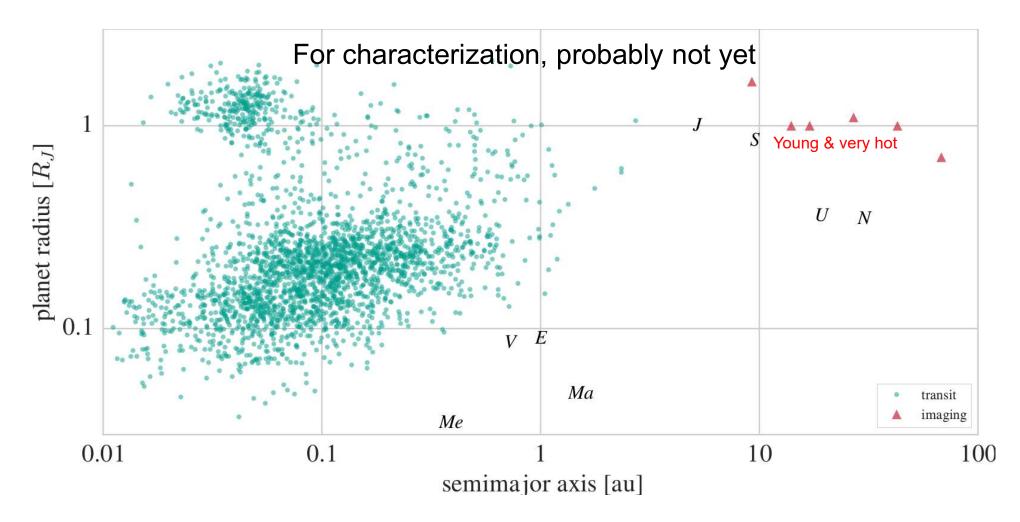
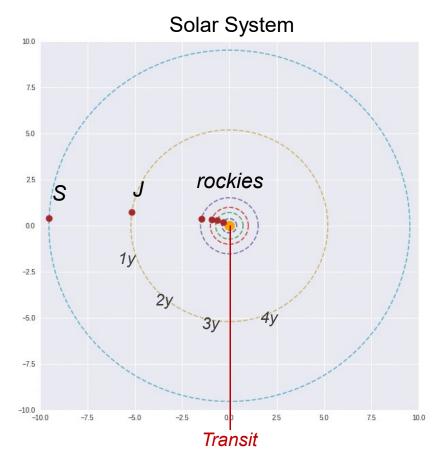


Have we already detected solar planet analogs?

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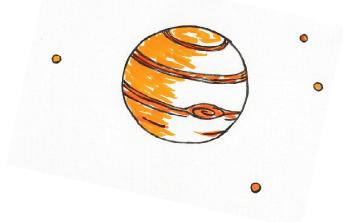


Hunting for single transiting events in Kepler archival data



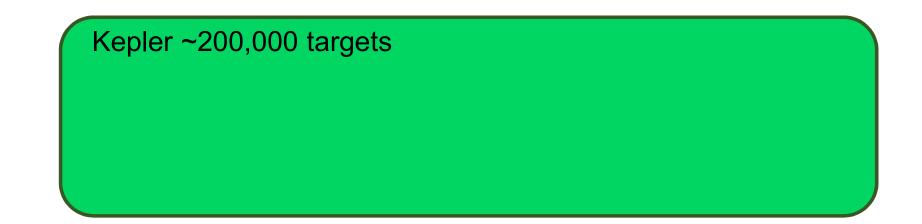
Standard pipelines normally require >~3 transiting events for efficient detection

Jupiter will exhibit a single or no transit during the lifetime of *Kepler spacecraft* (4yr)



- Visual Inspection by many citizen scientists (Planet Hunters; Wang+15)
- Visual Inspection by a student (Uehara, H.K., Masuda+16)
- Machine Learning (Foreman-Mackey+16)

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Kepler ~200,000 targets		
	Sun-like stars ~70,000 targets	KOI

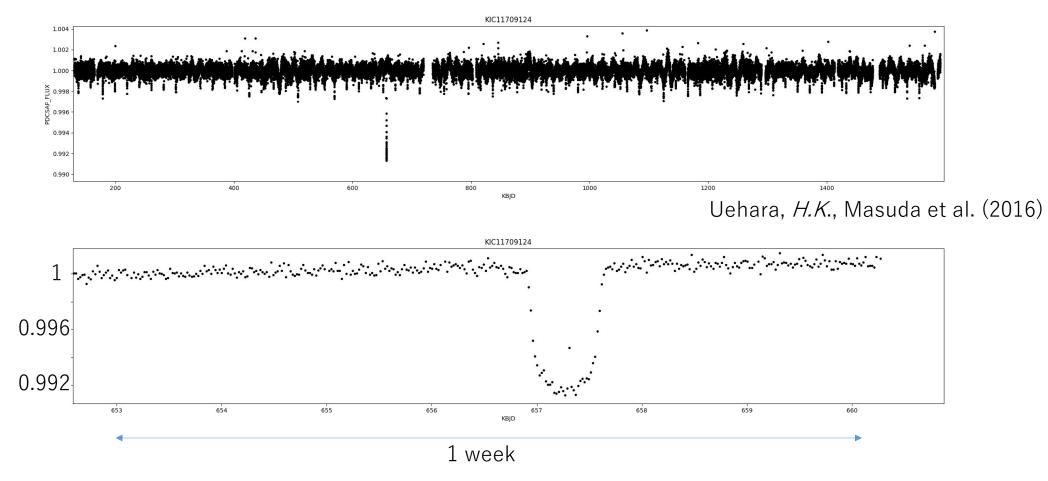
• Visual Inspection by many citizen scientists (Planet Hunters; Wang+15) Visual Inspection by a student (Uehara, *H.K.,* Masuda+16)

• Machine Learning (Foreman-Mackey+16, Osborn+16,)

Visual Inspection by Sho Uehara

C.C.

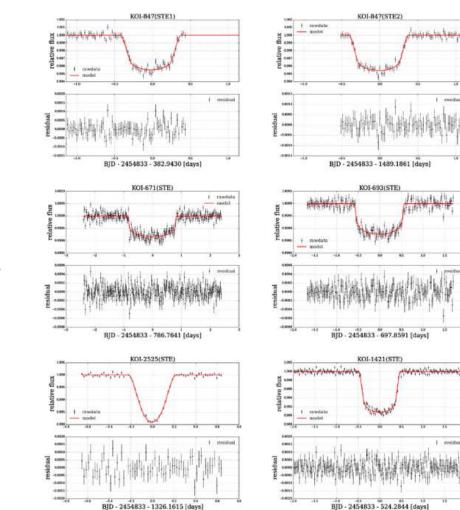
Long Period Transiting Jupiters in Kepler

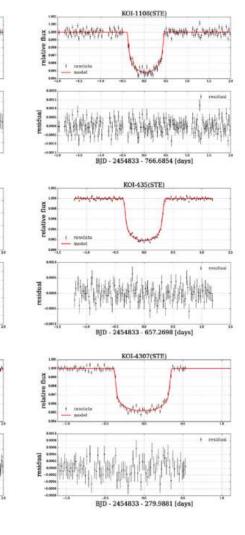


We searched for such events in ~8,000 KOI targets

14 new systems,7 are consistentwith exoplanets

others: EB or high-e planets





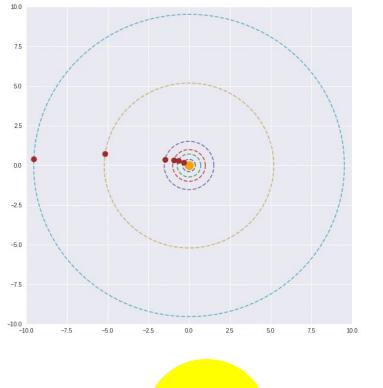
Inner transit events give us

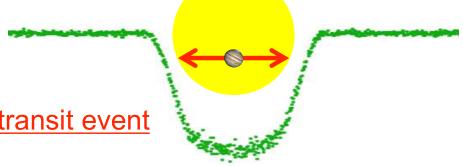
Duration
$$T_{\text{tot}} = \frac{2\sqrt{1-b^2}R_{\star}}{V}$$
,
Kepler-3rd law $V = \left(\frac{2\pi GM_{\star}}{P}\right)^{\frac{1}{3}}$
 $\therefore T_{\text{tot}}^3 = \frac{3(1-b^2)^{3/2}}{\pi^2}\frac{P}{\rho_{\star}}$

● *period* + duration ⇒ stellar density

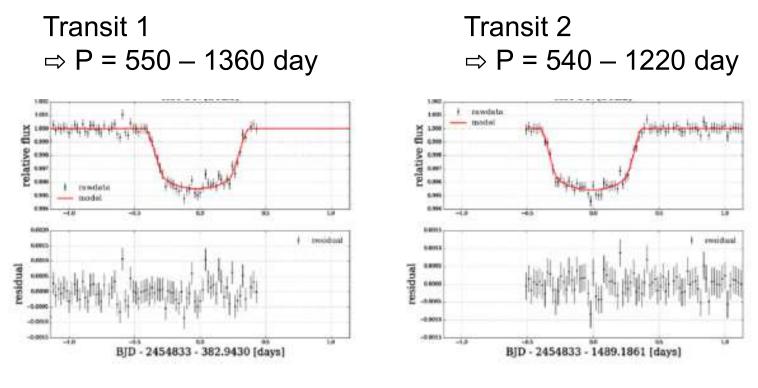
Known inner planets, known its period

 \rightarrow we know the stellar density \rightarrow we can estimate the period even for a single transit event



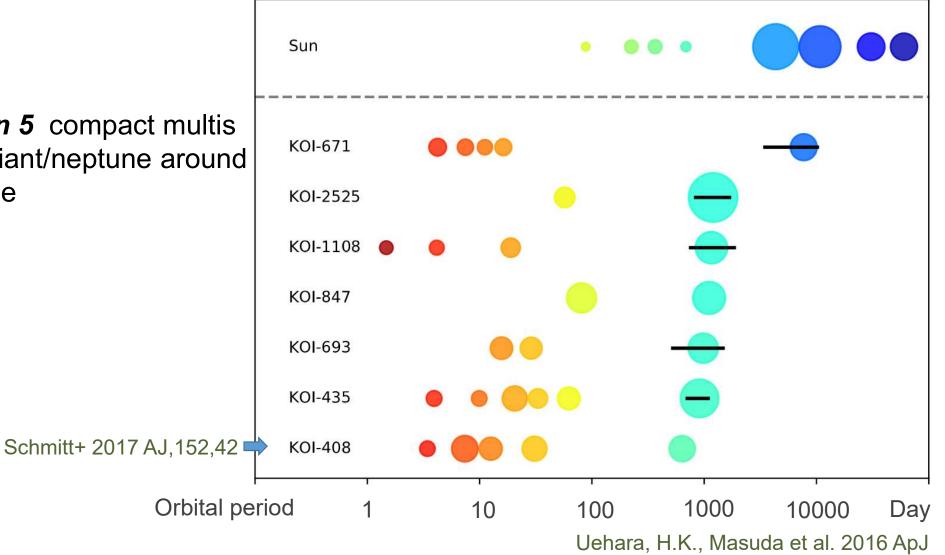


1 in **7** systems is a double transiting events: Test for the period estimate



Interval = 1106 day

At least, 1 in 5 compact multis has a gas giant/neptune around the snow line

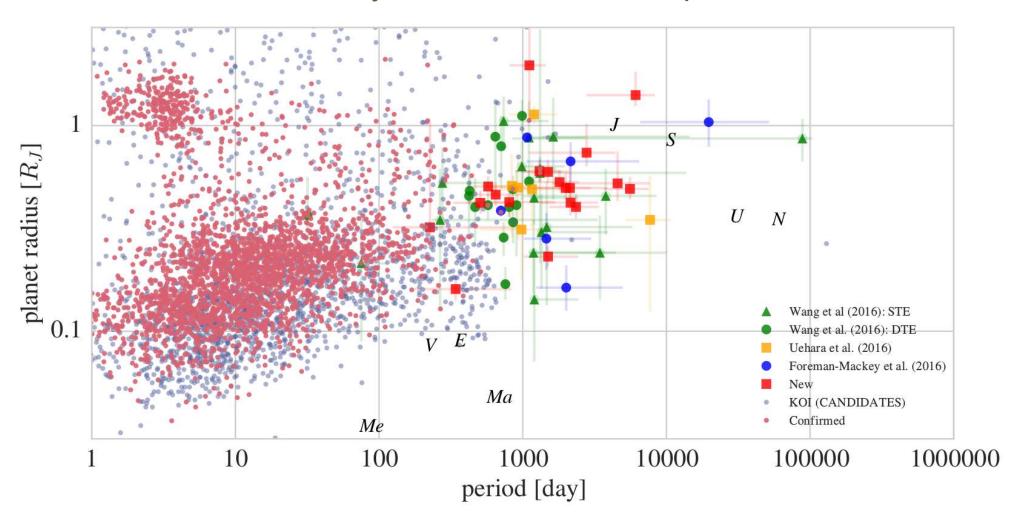


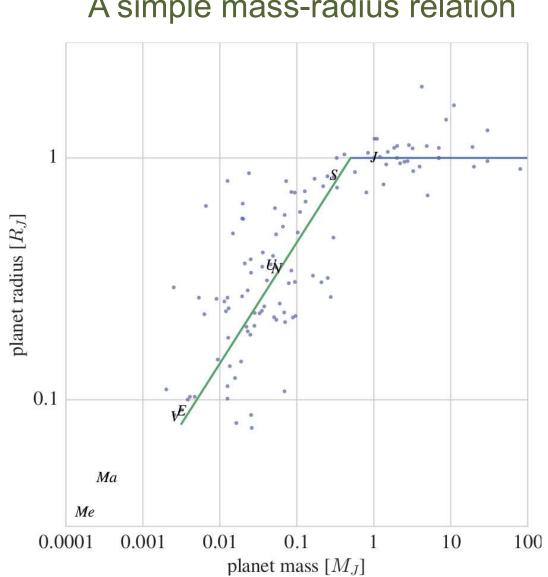
Compiling everything

- Visual Inspection by many citizen scientists (Planet Hunters; Wang+15)
- Visual Inspection by a student (Uehara, *H.K.,* Masuda+16)
- Machine Learning (Foreman-Mackey+16)
- + new STEs in KIC 200,000 targets (but less confident, not complete yet)

=> Compiling those findings, we have a tentative list of transiting planets around the snowline

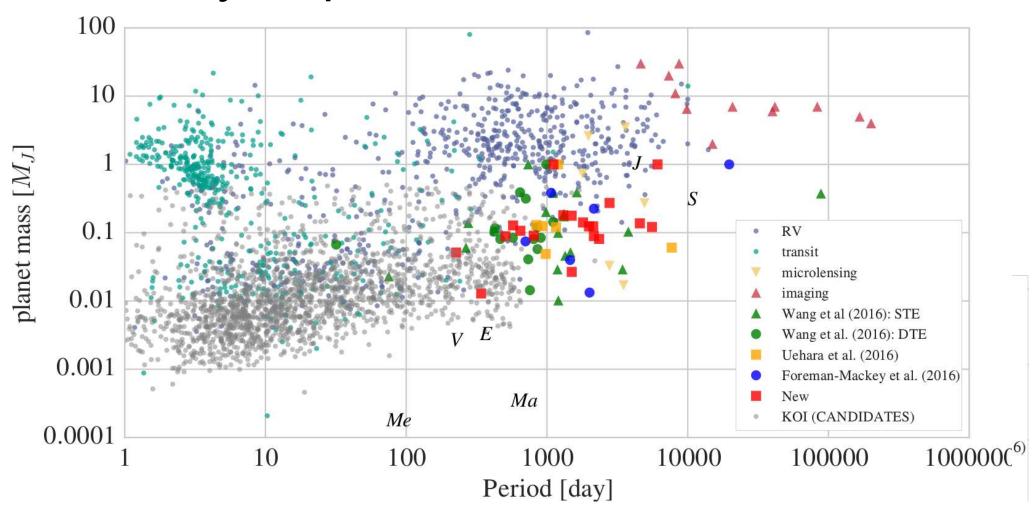
Preliminary results for all the Kepler stars





A simple mass-radius relation

Many Neptunes around snowline?

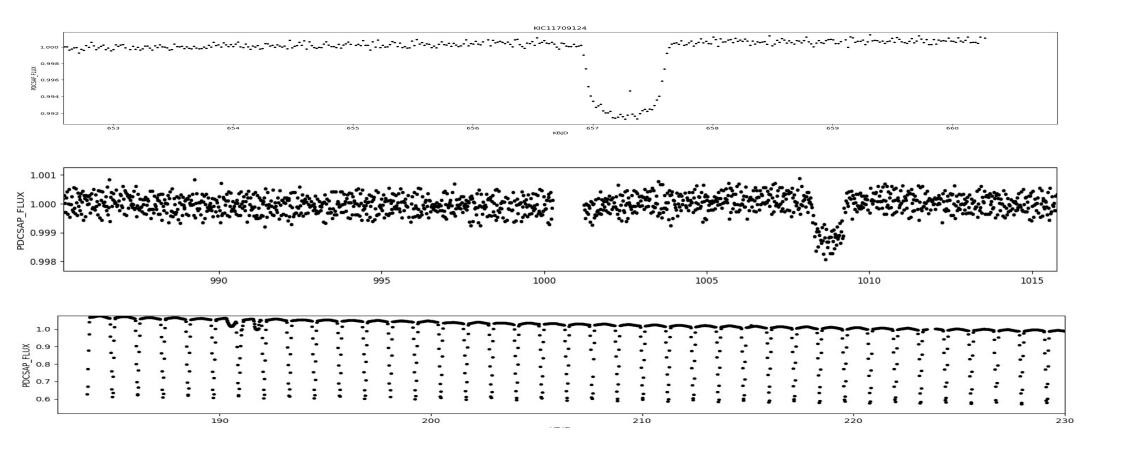


Long-period transiting objects more

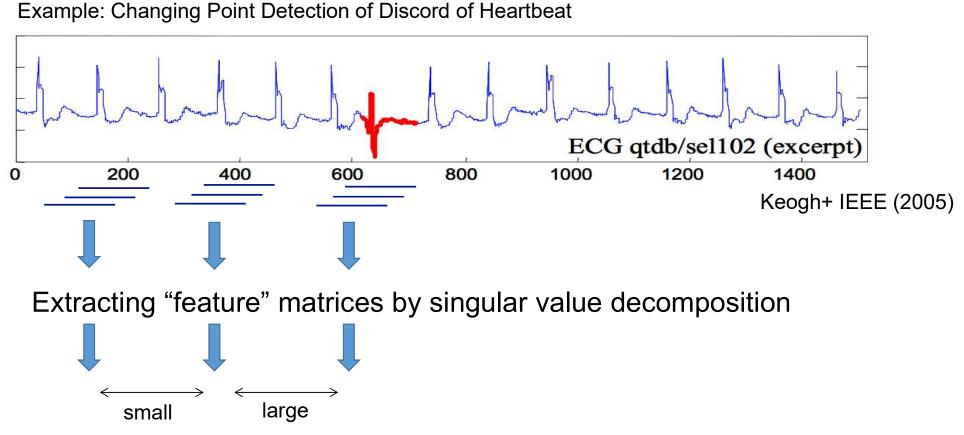
- Visual Inspection by a student (Uehara, H.K., Masuda+16)
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- Machine Learning (Foreman-Mackey+16, Osborn+16,)

Anomaly Detection (Tajiri, *H.K.*, Masuda in prep)

• Pair-wise Signal Search (H.K. Masuda, MacLeod to be submitted)



Anomaly detection: You don't really know what you want to find



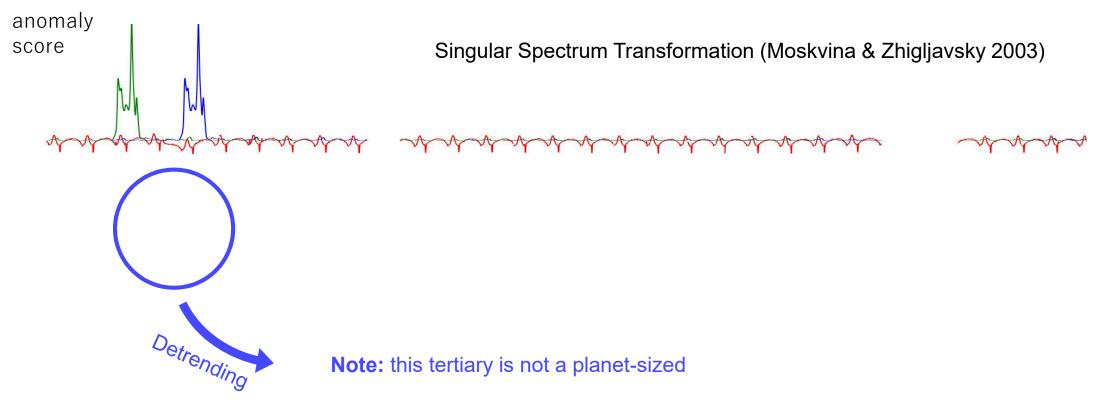
Measuring "distance" between the feature matrices (anomaly score)

Irregular heartbeat in a heartbeat star

Preliminary results in a survey in ~2,000 Kepler EB stars

(Tajiri, *H.K.*, Masuda in prep)

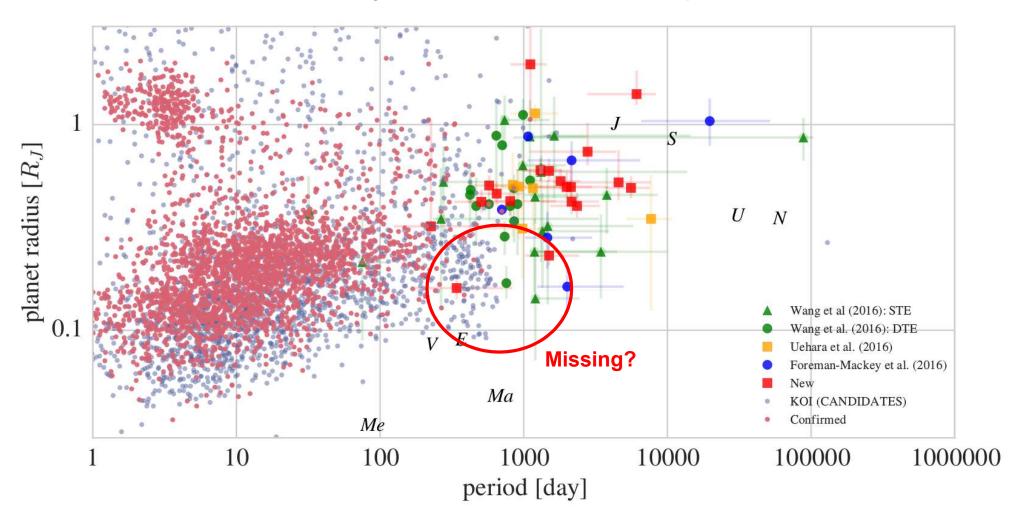
Irregular heartbeat in a heartbeat star



(Tajiri, *H.K.*, Masuda in prep)

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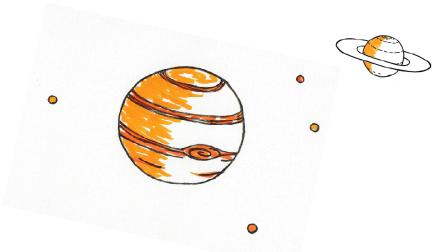
Preliminary results for all the Kepler stars



Problems

- Probably, many FPs in the tentative compiled list!
- -> follow-up observation using Subaru or other facilities
- Difficult to explore those planets beyond a period radius plane
- -> "Saturn mass at 2-3 au is detectable" Lauren's talk!
- We need a long-life <u>continuous</u> monitoring of nearby stars TESS, PLATO, or a new dedicated (micro)satellite

Summary



1. Single and double events reveal exoplanets around the snowline *Many Neptunes/sub-Saturns around the snowline?*

2. Those events are also a useful probe of circumbinary stars/planets and compact object binaries via self-gravitational lensing

Need long-life continuous monitoring of nearby stars!