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"Ghost in multimetric gravity"



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# Ghost in Multimetric Gravity

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## 1. Introduction

Recently, the ghost problem in non-linear massive gravity has been solved.

Massive gravity : a challenge to give mass to graviton in general relativity

linear level : Pauli-Fiertz mass term • • • no problem non linear level : a serious difficulty emerges

> graviton has spin 2 • • • mssless : 2 degrees of freedom massive : 5 degrees of freedom

In linear level, there are certainly 5 degrees of freedom. However, in non-linear level, extra 6<sup>th</sup> degree of freedom emerges and turns out to be a ghost particle. BD ghost

ghost particle : a particle having negative kinetic energy

For the consistency of the theory, ghost particles must be excluded.

### **Extension to bimetric gravity**

In non-linear massive gravity, metric is always decomposed into backgroundand and fluctuation.

The background metric can be promoted to a dynamical variable.

a theory of two dynamical metrics

It has been shown that no BD ghost is contained.

## Question

## Can we construct ghost-free theories with more than two metrics?

① : Trimetric case as a naïve extension of bimetric gravity is proposed. (2012, N. Khosravi et al)

②: Ghost-free multimetric gravity with vielbeins is proposed. (2012, K. Hinterbichler and R. A. Rosen)

2 dose not contain 1.

The problem of BD ghost is unsolved.

We solve the ghost problem in (1), and further develop our analysis to more general multimetric gravity.

2. Bimetric gravity 
$$g, f$$
: metric  

$$S_{bi} = M_g^2 \int d^4x \sqrt{-\det g} R[g] + M_f^2 \int d^4x \sqrt{-\det f} R[f]$$
Einstein Hilbert term
$$+ a \int d^4x \sqrt{-\det g} \Phi\left(\sqrt{g^{-1}f}\right)$$
Interaction term
$$M = M \quad : \text{ Planck mass} \quad R[g] R[f]: \text{ curvature} \quad g: \text{ coupling constant}$$

 $M_g, M_f$ : Planck mass, R[g], R[f]: curvature, a: coupling constant  $\Phi(A)$ : linear combination of Tr(A),  $(TrA)^2 - Tr(A^2)$ ,.... (up to fourth order in A)

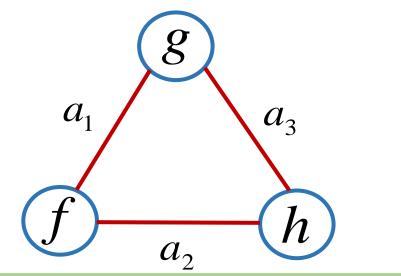
No BD-ghost is contained, and there exist one massless and one massive graviton. degrees of freedom : 2 + 5 = 7

## **3. Trimetric gravity** *g*, *f*, *h*: metric

$$S_{tri} = M_g^2 \int d^4 x \sqrt{-\det g} R[g] + M_f^2 \int d^4 x \sqrt{-\det f} R[f] + M_h^2 \int d^4 x \sqrt{-\det h} R[h] + a_1 \int \sqrt{-\det g} \Phi_1\left(\sqrt{g^{-1}f}\right) + a_2 \int \sqrt{-\det f} \Phi_2\left(\sqrt{f^{-1}h}\right) + a_3 \int \sqrt{-\det h} \Phi_3\left(\sqrt{h^{-1}g}\right)$$

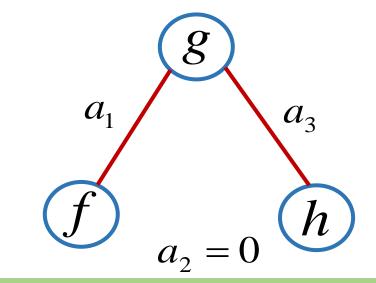
#### a nive extension of bimetric case

In this case, we verified the exsistence of a BD-ghost.



#### cutting one of interactions

Ghost-freedom is already shown in the vielbein formalism.



We perform the Hamiltonian analysis in homogeneous model.

$$g_{\mu\nu}dx^{\mu}dx^{\nu} = -N(t)^{2}dt^{2} + \gamma(t)_{ij}dx^{i}dx^{j}$$
  

$$f_{\mu\nu}dx^{\mu}dx^{\nu} = -L(t)^{2}dt^{2} + \omega(t)_{ij}dx^{i}dx^{j}$$
  

$$h_{\mu\nu}dx^{\mu}dx^{\nu} = -Q(t)^{2}dt^{2} + \rho(t)_{ij}dx^{i}dx^{j}$$

One of the metrics can be diagonalized through spacial rotations.

Hamiltonian :  $H = NC_N + LC_L + QC_Q$ 

N, L, Q: Laglange multipliers Constraints :  $C_N = 0, C_L = 0, C_Q = 0$ Degrees of freedom : 3 + 6 + 6 = 15 Constraints must be preserved in the time evolution.

$$\dot{C}_{N} = \{C_{N}, H\} = \{C_{N}, C_{L}\}L + \{C_{N}, C_{Q}\}Q \approx 0$$
  
$$\dot{C}_{L} = \{C_{L}, H\} = \{C_{L}, C_{N}\}N + \{C_{L}, C_{Q}\}Q \approx 0$$
  
$$\dot{C}_{Q} = \{C_{Q}, H\} = \{C_{Q}, C_{N}\}N + \{C_{Q}, C_{L}\}L \approx 0$$
  
$$\{C_{N}, C_{L}\} \propto a_{1}, \quad \{C_{L}, C_{Q}\} \propto a_{2}, \quad \{C_{N}, C_{Q}\} \propto a_{2}$$

$$\{C_N, C_L\} \propto a_1, \quad \{C_L, C_Q\} \propto a_2, \quad \{C_N, C_Q\} \propto a_3$$

$$\begin{pmatrix} 0 & a_1 & a_3 \\ -a_1 & 0 & a_2 \\ -a_3 & -a_2 & 0 \end{pmatrix} \begin{pmatrix} N \\ L \\ Q \end{pmatrix} = 0$$

We must determine the Lagrange multipliers to satisfy this condition.

Case 1: 
$$\begin{pmatrix} g \\ a_1 \\ f \\ a_2 = 0 \end{pmatrix} \begin{pmatrix} a_3 \\ a_3 \\ a_2 = 0 \end{pmatrix} \begin{pmatrix} 0 & a_1 & a_3 \\ -a_1 & 0 & 0 \\ -a_3 & 0 & 0 \end{pmatrix} \begin{pmatrix} N \\ L \\ Q \end{pmatrix} = 0$$

Secondary Constraints 
$$a_1 \propto \{C_N, C_L\} \approx 0, \quad a_3 \propto \{C_N, C_Q\} \approx 0$$

consistency in the time evolution

$$\begin{split} &d\{C_N, C_L\}/dt = NC_{NL,N} + LC_{NL,L} + QC_{NL,Q} \approx 0\\ &d\{C_N, C_Q\}/dt = NC_{NQ,N} + LC_{NQ,L} + QC_{NQ,Q} \approx 0 \end{split}$$

One of the Lagrange multipliers is left undetermined. (one gauge freedom)
The total number of constraints is five.

Masless graviton : 1

Massive graviton : 2

Total degrees of freedom : 
$$(15 \times 2 - 5 - 1)/2 = 12$$
 -

Case 2:  

$$a_1$$
 $a_3$ 
 $\begin{pmatrix} 0 & a_1 & a_3 \\ -a_1 & 0 & a_2 \\ -a_3 & -a_2 & 0 \end{pmatrix} \begin{pmatrix} N \\ L \\ Q \end{pmatrix} = 0$ 

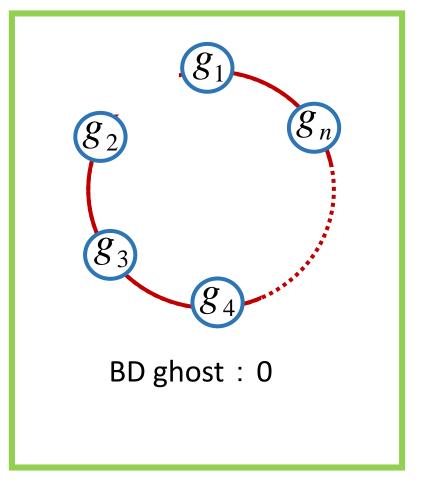
#### No seconday constraint

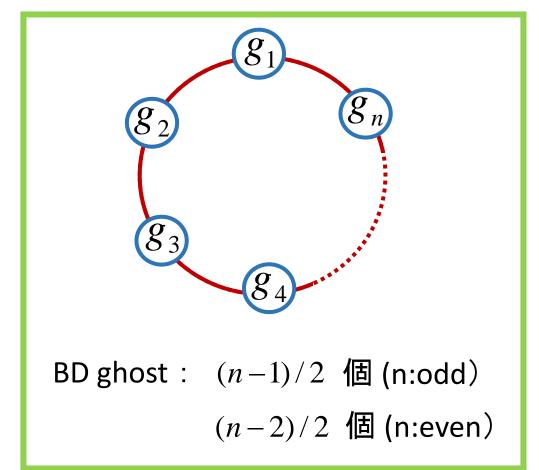
$$N = \frac{a_2}{a_1}Q, \quad L = -\frac{a_3}{a_1}Q, \quad Q: \text{ arbitrary}$$
Constraints : 3
Gauge freedom : 1
Sof freedom :  $(15 \times 2 - 3 - 1)/2 = 13$ 
Massive graviton : 1
Massive graviton : 2
BD ghost : 1

Total degrees of freedom : 
$$(15 \times 2 - 3 - 1)/2 = 13$$

**4. More general Multimetric gravity**  $g_1, g_2, ..., g_n$ : metrics

$$S_{n} = \sum_{k=1}^{n} M_{k}^{2} \int d^{4}x \sqrt{-\det g_{k}} R[g_{k}] + \sum_{k=1}^{n} a_{k} \int d^{4}x \sqrt{-\det g_{k}} \Phi_{k} \left( \sqrt{g_{k}^{-1}g_{k+1}} \right)$$





## 5. summary

• We showed that there exists a BD-ghost in the trimetric gravity as a naive extension of the recently proposed ghost-free bimetric gravity.

• We also studied more general multi metric gravity, and showed there always exist BD-ghosts if loop type interactions are contained.