"Ghost in multimetric gravity"

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# Ghost in <br> Multimetric Gravity 

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## 1. Introduction

Recently, the ghost problem in non-linear massive gravity has been solved.

Massive gravity : a challenge to give mass to graviton in general relativity
linear level : Pauli-Fiertz mass term • . . no problem non linear level : a serious difficulty emerges
graviton has spin 2 - . . mssless : 2 degrees of freedom massive : 5 degrees of freedom

In linear level, there are certainly 5 degrees of freedom.
However, in non-linear level, extra $6^{\text {th }}$ degree of freedom emerges and turns out to be a ghost particle.
ghost particle : a particle having negative kinetic energy
For the consistency of the theory, ghost particles must be excluded.

## Extension to bimetric gravity

In non-linear massive gravity, metric is always decomposed into backgroundand and fluctuation.
The background metric can be promoted to a dynamical variable.
a theory of two dynamical metrics
It has been shown that no $B D$ ghost is contained.

## Question

## Can we construct ghost-free theories with more than two metrics?

(1) : Trimetric case as a naïve extension of bimetric gravity is proposed. (2012, N. Khosravi et al)

The problem of BD ghost is unsolved.
(2) : Ghost-free multimetric gravity with vielbeins is proposed. (2012, K. Hinterbichler and R. A. Rosen)

## We solve the ghost problem in (1), and further develop our analysis to more general multimetric gravity.

## 2. Bimetric gravity $g, f$ : metric

$$
\begin{array}{rl|l}
S_{b i}= & M_{g}^{2} \int d^{4} x \sqrt{-\operatorname{det} g} R[g]+M_{f}^{2} \int d^{4} x \sqrt{-\operatorname{det} f} R[f] & \begin{array}{c}
\text { Einstein Hilbert term } \\
\\
\\
\\
\\
\end{array}+a \int d^{4} x \sqrt{-\operatorname{det} g} \Phi\left(\sqrt{g^{-1} f}\right)
\end{array}
$$

$M_{g}, M_{f}$ : Planck mass, $\quad R[g], R[f]$ : curvature , $a$ : coupling constant $\Phi(A)$ : linear combination of $\operatorname{Tr}(A),(\operatorname{Tr} A)^{2}-\operatorname{Tr}\left(A^{2}\right), \ldots$.
(up to fourth order in A)
No BD-ghost is contained, and there exist one massless and one massive graviton.

## 3. Trimetric gravity $g, f, h$ : metric

$$
\begin{aligned}
S_{r i i}= & M_{g}^{2} \int d^{4} x \sqrt{-\operatorname{det} g} R[g]+M_{f}^{2} \int d^{4} x \sqrt{-\operatorname{det} f} R[f]+M_{h}^{2} \int d^{4} x \sqrt{-\operatorname{det} h} R[h] \\
& +a_{1} \int \sqrt{-\operatorname{det} g} \Phi_{1}\left(\sqrt{g^{-1} f}\right)+a_{2} \int \sqrt{-\operatorname{det} f} \Phi_{2}\left(\sqrt{f^{-1} h}\right)+a_{3} \int \sqrt{-\operatorname{det} h} \Phi_{3}\left(\sqrt{h^{-1} g}\right)
\end{aligned}
$$

a nive extension of bimetric case
In this case, we verified the exsistence of a BD-ghost.

cutting one of interactions
Ghost-freedom is already shown in the vielbein formalism.


We perform the Hamiltonian analysis in homogeneous model.

$$
\begin{aligned}
& g_{\mu \nu} d x^{\mu} d x^{\nu}=-N(t)^{2} d t^{2}+\gamma(t)_{i j} d x^{i} d x^{j} \\
& f_{\mu \nu} d x^{\mu} d x^{\nu}=-L(t)^{2} d t^{2}+\omega(t)_{i j} d x^{i} d x^{j} \\
& h_{\mu \nu} d x^{\mu} d x^{\nu}=-Q(t)^{2} d t^{2}+\rho(t)_{i j} d x^{i} d x^{j}
\end{aligned}
$$

One of the metrics can be diagonalized through spacial rotations.

$$
\text { Hamiltonian : } H=N C_{N}+L C_{L}+Q C_{Q}
$$

$N, L, Q$ : Laglange multipliers
Constraints: $C_{N}=0, C_{L}=0, C_{Q}=0$
Degrees of freedom : $3+6+6=15$

Constraints must be preserved in the time evolution.

$$
\begin{aligned}
& \dot{C}_{N}=\left\{C_{N}, H\right\}=\left\{C_{N}, C_{L}\right\} L+\left\{C_{N}, C_{Q}\right\} Q \approx 0 \\
& \dot{C}_{L}=\left\{C_{L}, H\right\}=\left\{C_{L}, C_{N}\right\} N+\left\{C_{L}, C_{Q}\right\} Q \approx 0 \\
& \dot{C}_{Q}=\left\{C_{Q}, H\right\}=\left\{C_{Q}, C_{N}\right\} N+\left\{C_{Q}, C_{L}\right\} L \approx 0
\end{aligned}
$$

$$
\left\{C_{N}, C_{L}\right\} \propto a_{1}, \quad\left\{C_{L}, C_{Q}\right\} \propto a_{2}, \quad\left\{C_{N}, C_{Q}\right\} \propto a_{3}
$$

$$
\left(\begin{array}{ccc}
0 & a_{1} & a_{3} \\
-a_{1} & 0 & a_{2} \\
-a_{3} & -a_{2} & 0
\end{array}\right)\left(\begin{array}{l}
N \\
L \\
Q
\end{array}\right)=0 \quad \begin{aligned}
& \text { We must determine the Lagrange multipliers } \\
& \text { to satisfy this condition. }
\end{aligned}
$$

Case 1:


Secondary Constraints $\quad a_{1} \propto\left\{C_{N}, C_{L}\right\} \approx 0, \quad a_{3} \propto\left\{C_{N}, C_{Q}\right\} \approx 0$
consistency in the time evolution $d\left\{C_{N}, C_{L}\right\} / d t=N C_{N L, N}+L C_{N L, L}+Q C_{N L, Q} \approx 0$ $d\left\{C_{N}, C_{Q}\right\} / d t=N C_{N Q, N}+L C_{N Q, L}+Q C_{N Q, Q} \approx 0$

- One of the Lagrange multipliers is left undetermined. (one gauge freedom)
- The total number of constraints is five.

Total degrees of freedom : $(15 \times 2-5-1) / 2=12\left\{\begin{array}{l}\text { Masless graviton : } 1 \\ \text { Massive graviton : } 2 \\ \text { BD ghost : } 0\end{array}\right.$

Case 2:

$$
\left(\begin{array}{ccc}
0 & a_{1} & a_{3} \\
-a_{1} & 0 & a_{2} \\
-a_{3} & -a_{2} & 0
\end{array}\right)\left(\begin{array}{l}
N \\
L \\
Q
\end{array}\right)=0
$$

No seconday constraint

$$
N=\frac{a_{2}}{a_{1}} Q, \quad L=-\frac{a_{3}}{a_{1}} Q, \quad Q: \text { arbitrary }
$$

Constraints : 3 Gauge freedom : 1

Total degrees of freedom : $(15 \times 2-3-1) / 2=13\left\{\begin{array}{l}\text { Masless graviton : } 1 \\ \text { Massive graviton : } 2 \\ \text { BD ghost : } 1\end{array}\right.$
4. More general Multimetric gravity $g_{1}, g_{2}, \ldots, g_{n}$ : metrics

$$
S_{n}=\sum_{k=1}^{n} M_{k}^{2} \int d^{4} x \sqrt{-\operatorname{det} g_{k}} R\left[g_{k}\right]+\sum_{k=1}^{n} a_{k} \int d^{4} x \sqrt{-\operatorname{det} g_{k}} \Phi_{k}\left(\sqrt{g_{k}^{-1} g_{k+1}}\right)
$$



## 5. summary

- We showed that there exists a BD-ghost in the trimetric gravity as a naive extension of the recently proposed ghost-free bimetric gravity.
- We also studied more general multi metric gravity, and showed there always exist BD-ghosts if loop type interactions are contained.

