

Toshifumi Noumi, JGRG 22(2012)111207

"Effective field theory approach to quasi-single inflation"



GENERAL RELATIVITY AND GRAVITATION

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Effective field theory approach to quasi-single field inflation

Toshifumi Noumi (The University of Tokyo, Komaba)

based on arXiv: 1211.1624 with Masahide Yamaguchi and Daisuke Yokoyama (Tokyo Institute of Technology)

JGRG22 - RESCEU, 12th November 2012

classification of inflation models w.r.t. relevant dof

	relevant dof
single field	adiabatic mode ζ (massless)
multiple field	adiabatic + isocurvatures
	••••••

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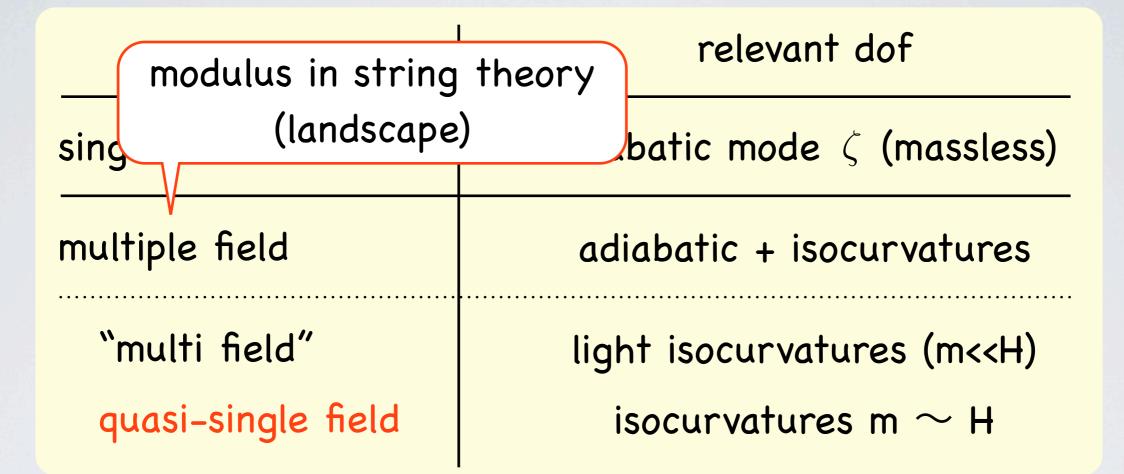
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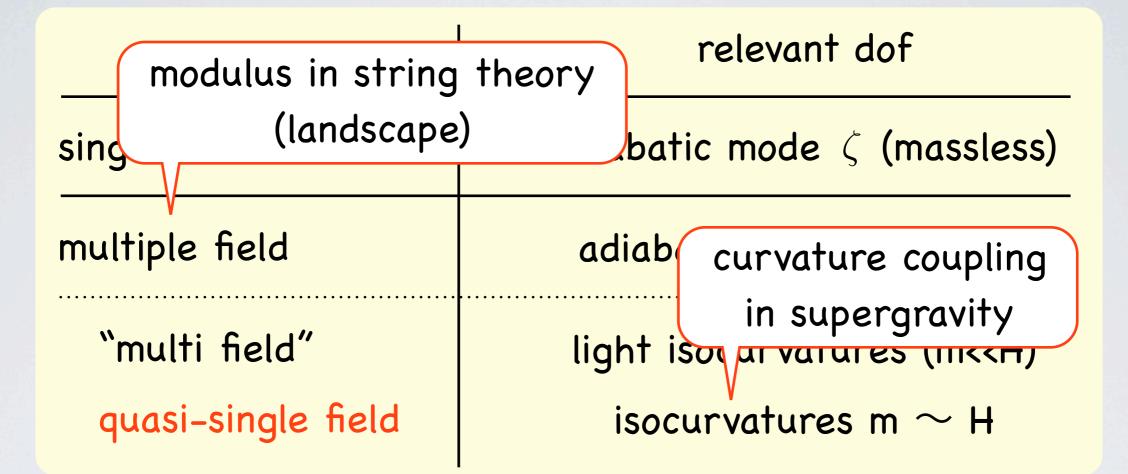
well motivated by model building
 (string inspired, supergravity based, ...)

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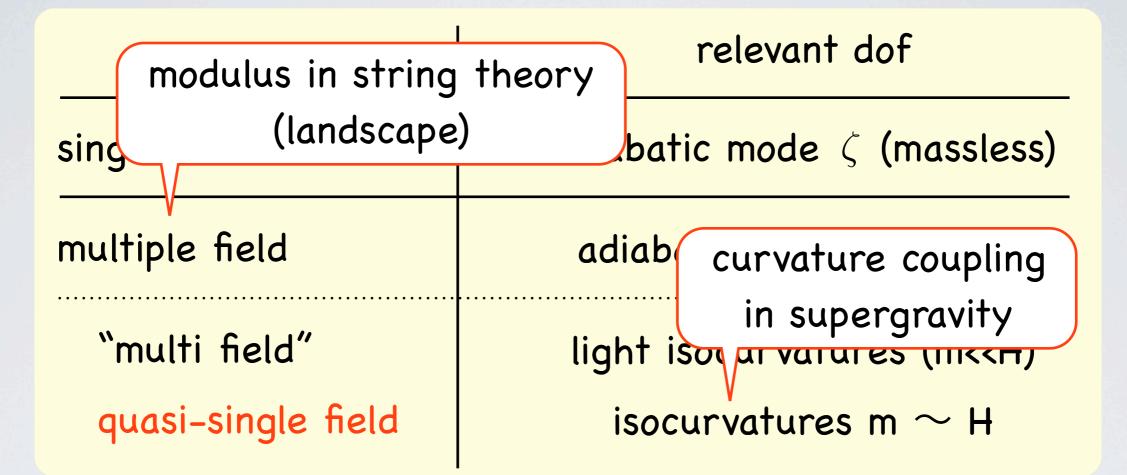
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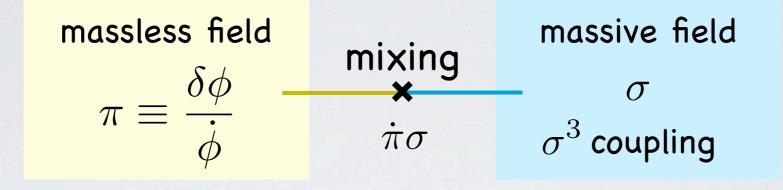
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well motivated by model building
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phenomenologically interesting
 (characteristic signatures in primordial perturbations)

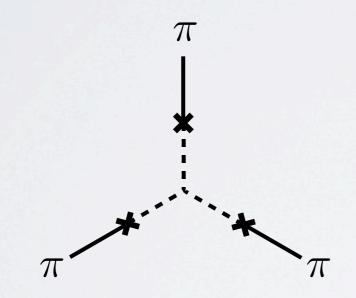
quasi-single field inflation [Chen-Wang '09]



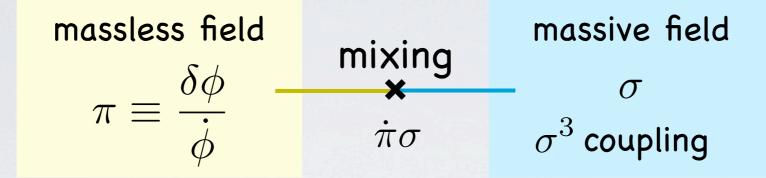
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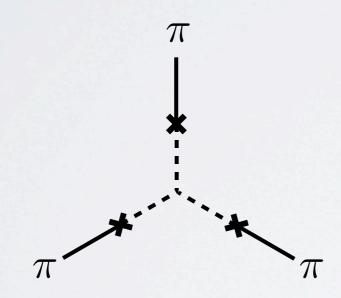
- can potentially give large non-Gaussianities



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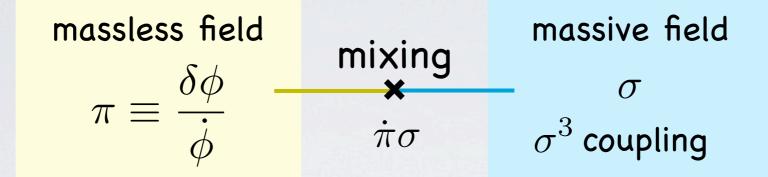


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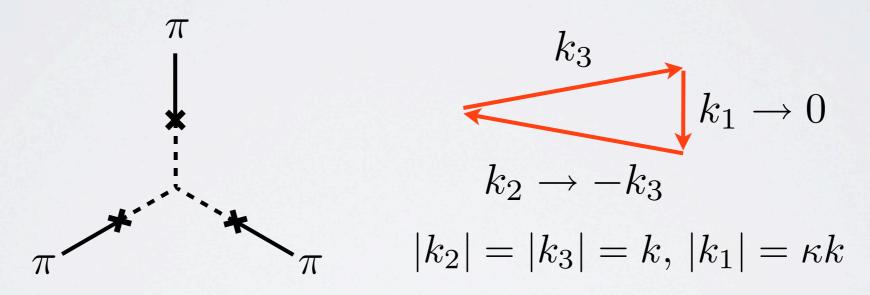


- intermediate shape between local and equilateral types

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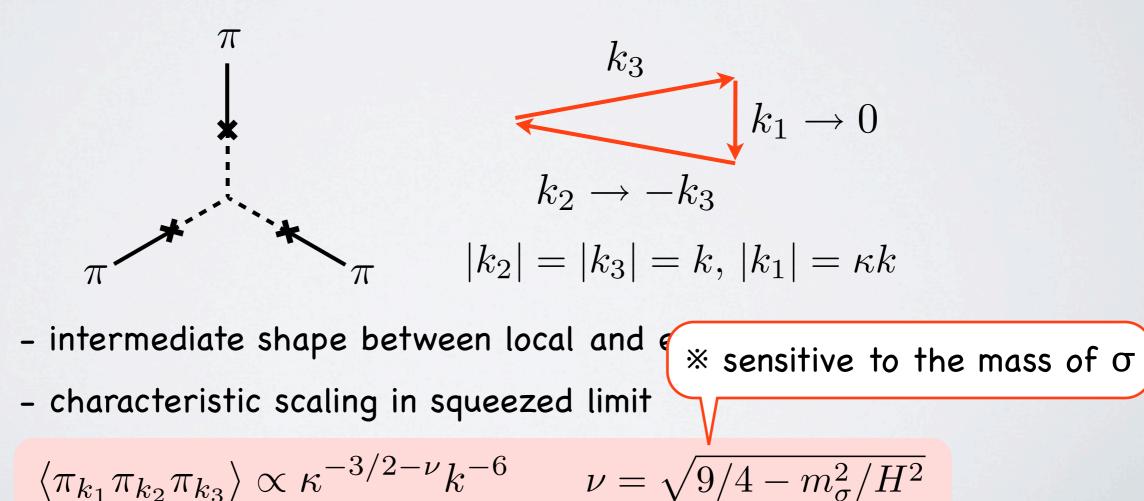
- intermediate shape between local and equilateral types
- characteristic scaling in squeezed limit

$$\langle \pi_{k_1} \pi_{k_2} \pi_{k_3} \rangle \propto \kappa^{-3/2 - \nu} k^{-6} \qquad \nu = \sqrt{9/4 - m_\sigma^2/H^2}$$

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Motivation

quasi-single field inflation:

- naturally realized in supergravity
- characteristic signatures in non-Gaussianities
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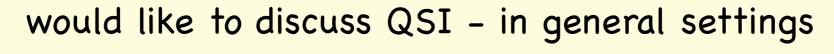
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would like to discuss QSI – in general settings – in a systematic way

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- in a systematic way

effective field theory approach!

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effective field theory approach [Cheung et al '07]

time dependent background evolution during inflation (time-dependent vev: R(t), $\phi_0(t)$)

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time dependent background evolution during inflation $f \xrightarrow{\text{spatial slicing}} (\text{time-dependent vev: } R(t), \phi_0(t))$ $\xrightarrow{} x^i$

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>> spontaneous breakdown of time diffeo

% unbroken time-dep. spatial diffeo

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advantages:

- systematic expansions in fluctuations and derivatives
- simplification in the dynamics of Goldstone boson $\boldsymbol{\pi}$
- relations between physics and non-Gaussianities are clear!

general action of quasi-single field inflation

relevant dof = three physical modes of graviton + additional massive scalar field σ

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schematically written as $S = S_{\text{grav}} + S_{\sigma} + S_{\text{mix}}$

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ex.
$$S_{\text{mix}} = \int d^4x \sqrt{-g} \left[\beta_1(t) \delta g^{00} \sigma + \beta_2(t) \delta g^{00} \partial^0 \sigma + \beta_3(t) \partial^0 \sigma - (\dot{\beta}_3(t) + 3H\beta_3(t)) \sigma \right]$$

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Stuckelberg method

$$S_{\text{mix}} = \int d^4x \, a^3 \left[-2\beta_1 \dot{\pi}\sigma + (2\beta_2 - \beta_3) \dot{\pi}\dot{\sigma} + \beta_3 \frac{\partial_i \pi \partial_i \sigma}{a^2} \right]$$
$$-\beta_1 \left(\dot{\pi}^2 - \frac{(\partial_i \pi)^2}{a^2} \right) \sigma + 3\beta_2 \dot{\pi}^2 \dot{\sigma} - 2\beta_2 \frac{\partial_i \pi \partial_i \sigma}{a^2} \dot{\pi} - \beta_2 \frac{(\partial_i \pi)^2}{a^2} \dot{\sigma} + \dots \right]$$

general action of quasi-single field inflation

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large non-Gaussianitis from mixings β_1 and β_2

 $\zeta \sim -H\pi$

Power spectrum

power spectrum

$$S_{\rm mix} = \int d^4x \, a^3 \left[\frac{-2\beta_1 \dot{\pi} \sigma + (2\beta_2 - \beta_3) \dot{\pi} \dot{\sigma} + \beta_3 \frac{\partial_i \pi \partial_i \sigma}{a^2}}{\tilde{\beta}_3} \right]$$

Power spectrum

power spectrum

$$\begin{cases} S_{\text{mix}} = \int d^4x \, a^3 \left[\underbrace{-2\beta_1 \dot{\pi}\sigma}_{\tilde{\beta}_1} + \underbrace{(2\beta_2 - \beta_3)}_{\tilde{\beta}_2} \dot{\pi} \dot{\sigma} + \underbrace{\beta_3}_{\tilde{\beta}_3} \frac{\partial_i \pi \partial_i \sigma}{a^2} \right] \\ \langle \zeta_{\mathbf{k}} \, \zeta_{\mathbf{k}'} \rangle = H^2 \, \langle \pi_{\mathbf{k}} \, \pi_{\mathbf{k}'} \rangle \\ & \stackrel{\sim}{} \underbrace{\mathbf{x}}_{\pi} \quad \mathbf{x}}_{\pi} \quad \mathbf{x}}}_{\pi} \quad \mathbf{x}}}}_{\pi} \quad \mathbf{x}}}_{\pi} \quad \mathbf{x}}}_{\pi} \quad \mathbf{x}}}_{\pi} \quad \mathbf{x}}}_{\pi} \quad \mathbf{x}}}}} x}}}_{\pi}}}_{\pi}} x}}_{\pi}} x}}}_{\pi}} x}}_{\pi}} x}}}_{\pi}}}}_{\pi}} x}}}_{\pi}}}x}}_{\pi}} x}}}_{\pi}} x}}_{\pi}} x}}_{\pi}} x}}}_{$$

Power spectrum

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power spectrum

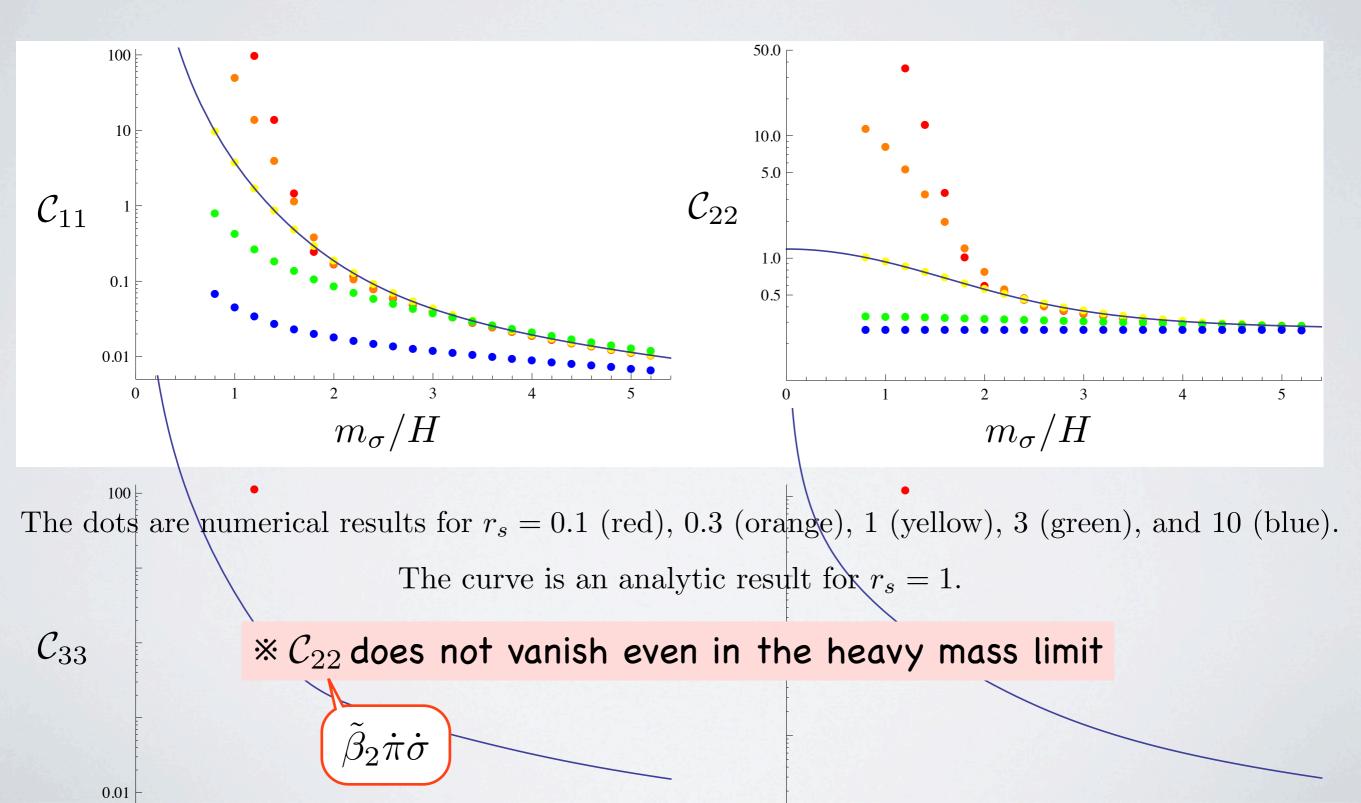
$$\begin{cases} S_{\text{mix}} = \int d^4x \, a^3 \left[\underbrace{-2\beta_1 \dot{\pi}\sigma}_{\tilde{\beta}_1} + \underbrace{(2\beta_2 - \beta_3)}_{\tilde{\beta}_2} \dot{\pi} \dot{\sigma} + \underbrace{\beta_3}_{\tilde{\beta}_3} \frac{\partial_i \pi \partial_i \sigma}{a^2} \right] \\ \langle \zeta_{\mathbf{k}} \zeta_{\mathbf{k}'} \rangle = H^2 \langle \pi_{\mathbf{k}} \pi_{\mathbf{k}'} \rangle \\ & \overset{\sim}{\mathbf{x}}_{\pi} & \overset{\times}{\pi} & \overset{\times}{$$

$$\begin{split} \langle \zeta_{\mathbf{k}} \zeta_{\mathbf{k}'} \rangle &= (2\pi)^3 \delta^{(3)} (\mathbf{k} + \mathbf{k}') \frac{2\pi^2}{k^3} \mathcal{P}_{\zeta}(k) \quad \text{deviation from single field} \\ \mathcal{P}_{\zeta}(k) &= \frac{H^2}{8\pi^2 M_{\text{Pl}}^2 \epsilon \, c_\pi} \left[1 + \frac{\pi}{4\alpha_\sigma^2} \frac{c_\pi^2}{M_{\text{Pl}}^2 (-\dot{H})} \left(\frac{\tilde{\beta}_1^2}{H^2} \, \mathcal{C}_{11} + \tilde{\beta}_2^2 \, \mathcal{C}_{22} + \tilde{\beta}_3^2 \, \mathcal{C}_{33} \right. \\ &+ \frac{\tilde{\beta}_1}{H} \, \tilde{\beta}_2 \, \mathcal{C}_{12} + \frac{\tilde{\beta}_1}{H} \, \tilde{\beta}_3 \, \mathcal{C}_{13} + \tilde{\beta}_2 \, \tilde{\beta}_3 \, \mathcal{C}_{23} \right) \right] \end{split}$$

Power spectrum

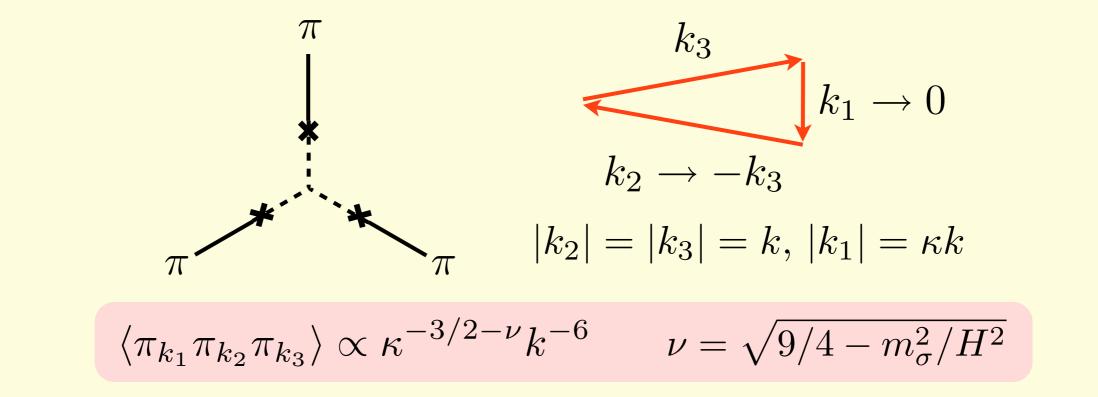
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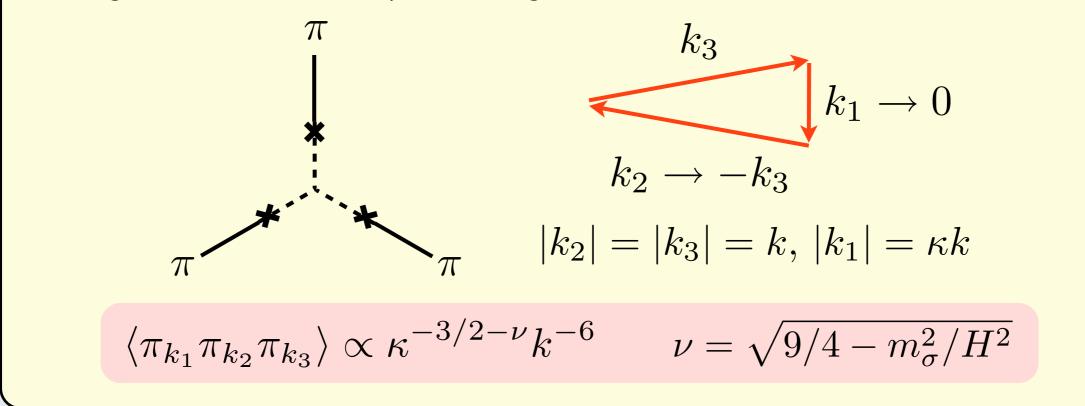
three point functions in the squeezed limit

- original models of quasi-single field inflation



three point functions in the squeezed limit

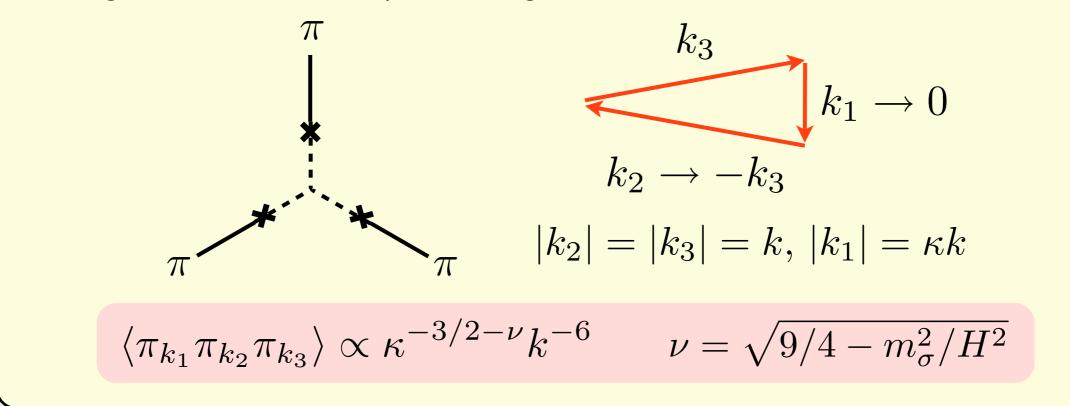
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- for general mixing and cubic couplings

three point functions in the squeezed limit

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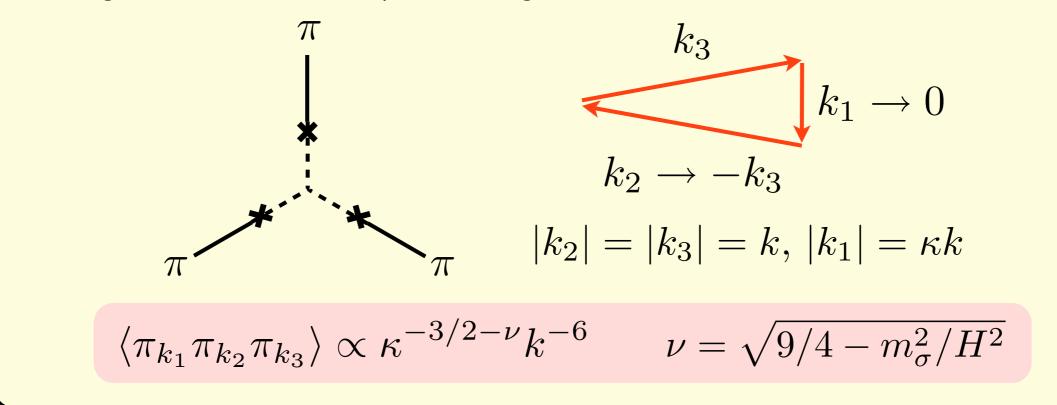


- for general mixing and cubic couplings

① scaling does not depend on details of mixing

three point functions in the squeezed limit

original models of quasi-single field inflation



- for general mixing and cubic couplings

1 scaling does not depend on details of mixing

2 determined only from cubic interaction in the diagram

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three point vertices	momentum dependence
$\dot{\pi}^3, \dot{\pi} \frac{(\partial_i \pi)^2}{a^2}$	$\kappa^{-1}k^{-6}$
$\dot{\pi}^2\sigma, \dot{\pi}\dot{\sigma}, \dot{\pi}\sigma^2, \dot{\pi}\sigma\dot{\sigma}, \dot{\pi}\dot{\sigma}^2, \ddot{\pi}\sigma\dot{\sigma},$	$\int \kappa^{-3/2-\nu} k^{-6} \qquad \text{for} m_{\sigma} < \frac{3}{2}H$
$\sigma^3, \sigma^2 \dot{\sigma}, \sigma \dot{\sigma}^2, \sigma \frac{(\partial_i \sigma)^2}{a^2}, \dot{\sigma}^3, \dot{\sigma} \frac{(\partial_i \sigma)^2}{a^2}$	$\left\{ \begin{array}{ll} \kappa^{-3/2} k^{-6} \sin[i\nu \log \kappa + \delta_{\nu}] & \text{for} m_{\sigma} > \frac{3}{2}H \end{array} \right.$
$\dot{\pi} \frac{\partial_i \pi \partial_i \sigma}{a^2}$	$\kappa^{-2}k^{-6}$
$\frac{(\partial_i \pi)^2}{a^2}\sigma, \frac{(\partial_i \pi)^2}{a^2}\dot{\sigma}, \frac{\partial_i \pi \partial_i \sigma}{a^2}\sigma, \frac{\partial_i \pi \partial_i \sigma}{a^2}\dot{\sigma}$	$\begin{cases} \kappa^{-3/2-\nu}k^{-6} & \text{for} m_{\sigma} < \sqrt{2}H \\ \kappa^{-2}k^{-6} & \text{for} m_{\sigma} > \sqrt{2}H \end{cases}$
	$\int \kappa^{-2} k^{-6} \qquad \text{for} m_{\sigma} > \sqrt{2}H$
$\dot{\pi} \frac{(\partial_i \sigma)^2}{a^2}$	$\begin{cases} \kappa^{-1/2-\nu}k^{-6} & \text{for} m_{\sigma} < \sqrt{2}H \\ \kappa^{-1}k^{-6} & \text{for} m_{\sigma} > \sqrt{2}H \end{cases}$
	$\left \begin{array}{c} & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & \\ & $

non-trivial scaling in the squeezed limit when mixing is relevant!

 $\nu = \sqrt{9/4 - m_{\sigma}^2/H^2}$ $0 < \nu < 3/2$ or $\nu =$ pure imaginary

Summary and prospects

summary

applied EFT approach to QSI

- systematic expansions in fluctuations and derivatives
- simplification of action for π in decoupling regime

relation between physics & non-Gaussianities is clear
calculated power spectrum for constant mixing
discussed scaling of 3-pt functions in squeezed limit
sensitive to # of fields and their mass
also discussed effects of heavy particles, sharp turning

prospects

full non-Gaussianities, detectability, ...

EFT for sugra based inflation,

more on sharp turning, ...

THANK YOU!!