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Teleparallel Conformal Invariant Models Induced by Kaluza-Klein Reduction

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- 1 Teleparallel Gravity
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- 5 Weak Field Approximation





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Standard Gravity Theory

- General Relativity
 - Einstein equation

$$G_{\mu\nu} = 8\pi G T_{\mu\nu}$$
 with $G_{\mu\nu} := R_{\mu\nu} - \frac{1}{2} R g_{\mu\nu}$

(Einstein, Nov. 25, 1915)





Hilbert action

$$\frac{-1}{2\kappa}\int d^4x\sqrt{-g}R+S_{\rm m}$$

(Hilbert, Nov. 20, 1915)

• "Spacetime tells matter how to move; matter tells spacetime how to curve." — John Wheeler.

Absolute Parallelism

Introducing the orthonormal frame $e_i = e_i{}^{\mu}\partial_{\mu}$ in Weitzenböck geometry T_4 :

$$g(e_i,e_j)=\eta_{ij} \quad \text{with} \quad \eta_{ij}=\text{diag}(+1,-1,-1,-1)$$

 $\Longrightarrow \partial_{\mu} \to e_{i} = e_{i}{}^{\mu}\partial_{\mu} \quad \text{and} \quad \Gamma^{\rho}{}_{\mu\nu} \to \omega^{i}{}_{j\nu} = e^{i}{}_{\rho}\Gamma^{\rho}{}_{\mu\nu}e_{j}{}^{\mu} + e^{i}{}_{\sigma}\partial_{\nu}e_{j}{}^{\sigma} \,.$

■ Parallel vectors (absolute parallelism) (Cartan, 1922/Eisenhart, 1925)

$$\nabla e_i = dx^{\nu} (\partial_{\nu} e_i{}^{\rho} + e_i{}^{\mu} \overset{\mathsf{w}}{\Gamma}{}^{\rho}{}_{\mu\nu}) \partial_{\rho} := dx^{\nu} (\nabla_{\!\nu} e_i{}^{\rho}) \partial_{\rho} = 0 \,.$$

Teleparallel Equivalent to GR in T_4

- Weitzenböck connection $\tilde{\Gamma}^{\rho}{}_{\mu\nu} = \{ {}^{\rho}{}_{\mu\nu} \} + K^{\rho}{}_{\mu\nu}$.
- Teleparallel Equivalent to GR (GR $_{\parallel}$ or TEGR):

$$R(\Gamma) = \tilde{R}(e) + T - 2\,\tilde{\nabla}_{\mu} T^{\mu} = 0 \quad \Longrightarrow \quad \left[-\tilde{R}(e) = T - 2\,\tilde{\nabla}_{\mu} T^{\mu} \right] \quad (T_{\mu} := T^{\nu}{}_{\nu\mu})$$

Torsion Scalar (Einstein, 1929)

$$T \equiv K^{\nu}{}_{\mu\nu}K^{\mu\sigma}{}_{\sigma} - K^{\rho}{}_{\mu\nu}K^{\mu\nu}{}_{\rho}$$
$$= \frac{1}{4}T^{\rho}{}_{\mu\nu}T^{\rho}{}_{\mu\nu} + \frac{1}{2}T^{\rho}{}_{\mu\nu}T^{\nu\mu}{}_{\rho} - 1T^{\nu}{}_{\mu\nu}T^{\sigma\mu}{}_{\sigma} = \frac{1}{2}T^{i}{}_{\mu\nu}S_{i}{}^{\mu\nu}$$

 $S_\rho{}^{\mu\nu} \equiv K^{\mu\nu}{}_\rho + \delta^\mu_\rho\,T^{\sigma\nu}{}_\sigma - \delta^\nu_\rho\,T^{\sigma\mu}{}_\sigma = -S_\rho{}^{\nu\mu} \text{ is superpotential }.$

TEGR action

$$S_{\text{TEGR}} = \frac{1}{2\kappa} \int d^4x \, e \, T \qquad (e = \sqrt{-g}) \, .$$

• New General Relativity (NGR): $\left(\frac{1}{4}, \frac{1}{2}, -1\right) \longrightarrow (a, b, c)$

(Hayashi & Shirafuji, 1979)

<u>Motivation</u>

- Fundamental fields in GR:
 - Metric tensor $g_{\mu\nu}$

$$\Longrightarrow$$
 Levi-Civita connection $\{ {}^{
ho}_{\mu
u} \} = rac{1}{2} g^{
ho\sigma} \Big(\partial_{\mu}g_{\sigma
u} + \partial_{
u}g_{\mu\sigma} - \partial_{\sigma}g_{\mu
u} \Big)$

- Fundamental field in Teleparallelism:
 - \blacksquare Veierbein fields $e^i{}_\mu$
 - \implies Weitzenböck connection $\Gamma^{\rho}{}_{\mu\nu} = e_i{}^{\rho}\partial_{\nu}e^i{}_{\mu}.$

What will be arised from the extra dimensions?

(i) Any new interaction different from GR?(ii) New classes of gravity theory in Teleparallelism?

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5D Teleparallelism

- Embedding: $T_4 \longrightarrow T_5$.
- 5D metric is given by

$$\bar{g}_{MN} = \begin{pmatrix} g_{\mu\nu} - \kappa^2 \phi^4 A_\mu A_\nu & \kappa \phi^2 A_\mu \\ \kappa \phi^2 A_\nu & -\phi^2 \end{pmatrix} \,.$$

 \blacksquare The coframes in T_5 are

$$\begin{cases} \theta^i = e^i{}_\mu \, dx^\mu \,, \\ \theta^5 = e^5{}_\mu \, dx^\mu + e^5{}_5 \, dx^5 = -\kappa \, \phi \, A_\mu \, dx^\mu + \phi \, dy \,. \end{cases} \implies \quad dy = \kappa \, A_\mu \, e_i{}^\mu \theta^i + \frac{1}{\phi} \, \theta^5 \end{cases}$$

The torsion components in T_5

$$\begin{split} \bar{T}^{i}{}_{jk} &= T^{i}{}_{jk} + \kappa \, A_{\mu}(\partial_{5}e^{i}{}_{\nu})(e_{j}{}^{\mu}e_{k}{}^{\nu} - e_{k}{}^{\mu}e_{j}{}^{\nu})\,,\\ \bar{T}^{i}{}_{5j} &= \frac{1}{\phi}(\partial_{5}e^{i}{}_{\mu})e_{j}{}^{\mu}\,,\\ \bar{T}^{5}{}_{ij} &= -\frac{\kappa}{2}\,\phi\,e_{i}{}^{\mu}e_{j}{}^{\nu}\,F_{\mu\nu} + \kappa^{2}\,\phi\,A_{\mu}(\partial_{5}A_{\nu})(e_{i}{}^{\nu}e_{j}{}^{\mu} - e_{j}{}^{\nu}e_{i}{}^{\mu}),\\ \bar{T}^{5}{}_{i5} &= \frac{1}{\phi}e_{i}{}^{\mu}(\partial_{\mu}\phi) + \frac{1}{\phi}\kappa\,A_{\mu}e_{i}{}^{\mu}(\partial_{5}\phi) + \kappa\,e_{i}{}^{\mu}(\partial_{5}A_{\mu})\,. \end{split}$$

Extending to 5D

■ NGR torsion scalar in 4D is given by

$$T_{\text{NGR}} = a \, T_{ijk} \, T^{ijk} + b \, T_{ijk} \, T^{kji} + c \, T^{j}{}_{ji} \, T^{k}{}_{k}{}^{i} := \frac{1}{2} \, T^{i}{}_{jk} \, \Sigma_{i}{}^{jk} \,,$$

where

$$\Sigma_i{}^{jk} = 2a T_i{}^{jk} + b \left(T^{kj}{}_i - T^{jk}{}_i \right) + c \left(\delta_i^k T^{lj}{}_l - \delta_i^j T^{lk}{}_l \right).$$

■ Extending to a 5D torsion scalar theory

Extended torsion scalar

$${}^{(5)}T^{(\text{ext})} = a\,\bar{T}_{LMN}\bar{T}^{LMN} + b\,\bar{T}_{LMN}\bar{T}^{NML} + c\,\bar{T}^{L}{}_{LM}\bar{T}^{N}{}_{N}{}^{M}$$

• Decomposition
$$(\eta_{\hat{5}\hat{5}} = \eta^{\hat{5}\hat{5}} = -1)$$

$$\begin{split} ^{(5)}T^{(\text{ext})} = \bar{T}_{\text{NGR}} + 2a\,\bar{T}_{i\hat{5}j}\bar{T}^{i\hat{5}j} + a\,\bar{T}_{\hat{5}ij}\bar{T}^{\hat{5}ij} + b\,\bar{T}_{i\hat{5}j}\bar{T}^{j\hat{5}i} + 2b\,\bar{T}_{\hat{5}ij}\bar{T}^{j\hat{5}} \\ &+ (2a+b+c)\,\bar{T}_{\hat{5}i\hat{5}}\bar{T}^{\hat{5}i\hat{5}} + 2c\,\bar{T}^{j}{}_{j}{}^{i}\bar{T}^{\hat{5}}{}_{\hat{5}}{}^{i} + c\,\bar{T}^{i}{}_{i\hat{5}}\bar{T}^{j}{}_{j}{}^{\hat{5}} \,. \end{split}$$

$$\begin{split} \bar{T}_{\text{NGR}} = & T_{\text{NGR}} + 4a\kappa T_{l}{}^{\rho\sigma}A_{\rho}(\partial_{5}e^{l}{}_{\sigma}) - 2a\kappa^{2}(g^{\mu\rho}A_{\mu}A_{\rho})(g^{\nu\sigma}\eta_{il}(\partial_{5}e^{i}{}_{\nu})(\partial_{5}e^{l}{}_{\sigma})) \\ & - 2a\kappa^{2}\eta_{il}(g^{\nu\rho}A_{\rho}(\partial_{5}e^{i}{}_{\nu}))(g^{\mu\sigma}A_{\mu}(\partial_{5}e^{l}{}_{\sigma})) + 2b\kappa T^{\sigma\rho}{}_{k}A_{\rho}(\partial_{5}e^{k}{}_{\sigma}) \\ & - 2b\kappa T^{\rho\sigma}{}_{k}A_{\rho}(\partial_{5}e^{k}{}_{\sigma}) + b\kappa^{2}(g^{\mu\rho}A_{\mu}A_{\rho})(\partial_{5}e^{i}{}_{\nu})(\partial_{5}e^{k}{}_{\sigma})e_{k}{}^{\nu}e_{i}{}^{\sigma} \\ & - 2b\kappa^{2}(g^{\mu\sigma}A_{\mu}(\partial_{5}e^{k}{}_{\sigma}))(A_{\rho}e_{i}{}^{\rho})(\partial_{5}e^{i}{}_{\nu})e_{k}{}^{\nu} \\ & + b\kappa^{2}(A_{\mu}e_{k}{}^{\mu})(A_{\rho}e_{i}{}^{\rho})(g^{\nu\sigma}(\partial_{5}e^{i}{}_{\nu})(\partial_{5}e^{k}{}_{\sigma})) \\ & + 2c\kappa T^{j}{}_{j}{}^{\sigma}(A_{\rho}e_{k}{}^{\rho})(g^{\nu\sigma}(\partial_{5}e^{j}{}_{\nu})(\partial_{5}e^{k}{}_{\sigma})) \\ & + c\kappa^{2}(A_{\mu}e_{j}{}^{\mu})(A_{\rho}e_{k}{}^{\rho})(g^{\nu\sigma}(\partial_{5}e^{j}{}_{\nu}))(\partial_{5}e^{k}{}_{\sigma})) \\ & - 2c\kappa^{2}(A_{\mu}e_{j}{}^{\mu})(\partial_{5}e^{k}{}_{\sigma})e_{k}{}^{\sigma}(g^{\nu\rho}A_{\rho}(\partial_{5}e^{j}{}_{\nu})) \\ & + c\kappa^{2}(g^{\mu\rho}A_{\mu}A_{\rho})(\partial_{5}e^{j}{}_{\nu})e_{j}{}^{\nu}(\partial_{5}e^{k}{}_{\sigma})e_{k}{}^{\sigma}, \end{split}$$

$$\bar{T}_{i\hat{5}j}\bar{T}^{i\hat{5}j} = \frac{1}{\phi^2}\,\eta^{\hat{5}\hat{5}}\eta_{ik}g^{\mu\nu}(\partial_5 e^i{}_{\mu})(\partial_5 e^k{}_{\nu})\,,$$

$$\bar{T}_{\hat{5}ij}\bar{T}^{\hat{5}ij} = \frac{\kappa^2}{4} \phi^2 \eta_{\hat{5}\hat{5}} g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} + 2\kappa^3 \phi^2 \eta_{\hat{5}\hat{5}} g^{\mu\sigma} g^{\nu\rho} A_{\mu} (\partial_5 A_{\nu}) F_{\sigma\rho} + 2\kappa^4 \phi^2 \eta_{\hat{5}\hat{5}} (g^{\mu\rho} A_{\mu} A_{\rho}) (g^{\nu\sigma} (\partial_5 A_{\nu}) (\partial_5 A_{\sigma})) - 2\kappa^4 \phi^2 \eta_{\hat{5}\hat{5}} (g^{\mu\sigma} A_{\mu} (\partial_5 A_{\sigma})) (g^{\nu\rho} (\partial_5 A_{\nu}) A_{\rho}) ,$$

$$\begin{split} \bar{T}_{i\hat{5}j}\bar{T}^{j\hat{5}i} &= \frac{1}{\phi^2} \eta^{\hat{5}\hat{5}} (\partial_5 e^i{}_{\mu}) (\partial_5 e^j{}_{\nu}) e_j{}^{\mu} e_i{}^{\nu} , \\ \bar{T}_{\hat{5}ij}\bar{T}^{j\hat{1}\hat{5}} &= \frac{\kappa}{2} e_j{}^{\nu} g^{\mu\rho} F_{\mu\nu} (\partial_5 e^j{}_{\rho}) - \kappa^2 (A_{\mu} e_j{}^{\mu}) (g^{\nu\rho} (\partial_5 A_{\nu}) (\partial_5 e^j{}_{\rho})) \\ &+ \kappa^2 (A_{\mu} g^{\mu\rho}) (\partial_5 A_{\nu}) (\partial_5 e^j{}_{\rho}) e_j{}^{\nu} , \end{split}$$

$$\begin{split} \bar{T}_{5i\bar{5}}\bar{T}^{5i\bar{5}} &= \frac{1}{\phi^2}\eta_{55}\eta^{5\bar{5}}(g^{\mu\nu}(\partial_{\mu}\phi)(\partial_{\nu}\phi)) + \frac{2\kappa}{\phi^2}\eta_{55}\eta^{5\bar{5}}(\partial_5\phi)(g^{\mu\nu}(\partial_{\mu}\phi)A_{\nu}) \\ &+ \frac{2\kappa}{\phi}\eta_{55}\eta^{5\bar{5}}(g^{\mu\nu}(\partial_{\mu}\phi)(\partial_5A_{\nu})) + \frac{\kappa^2}{\phi^2}\eta_{55}\eta^{5\bar{5}}(\partial_5\phi)^2(g^{\mu\nu}A_{\mu}A_{\nu}) \\ &+ \frac{2\kappa^2}{\phi}\eta_{55}\eta^{5\bar{5}}(\partial_5\phi)(g^{\mu\nu}A_{\mu}(\partial_5A_{\nu})) + \kappa^2\eta_{5\bar{5}}\eta^{5\bar{5}}(g^{\mu\nu}(\partial_5A_{\mu})(\partial_5A_{\nu})) , \\ \bar{T}^{j}{}_{ji}\bar{T}^{\bar{5}}{}_{\bar{5}}{}^{i} &= \frac{-\frac{1}{\phi}T^{\rho}(\partial_{\rho}\phi)}{-\frac{1}{\phi}T^{\rho}A_{\rho}(\partial_5\phi) - \kappa T^{\rho}(\partial_5A_{\rho})} \\ &- \frac{\kappa}{\phi}(A_{\mu}e_{j}{}^{\mu})(g^{\nu\rho}(\partial_5e^{j}{}_{\nu})(\partial_{\rho}\phi)) + \frac{\kappa^2}{\phi^2}(\partial_5\phi)(\partial_5e^{j}{}_{\nu})e_{j}{}^{\nu}(g^{\mu\rho}A_{\mu}A_{\rho}) \\ &- \kappa^2(A_{\mu}e_{j}{}^{\mu})(g^{\nu\rho}(\partial_5e^{j}{}_{\nu})(\partial_5A_{\rho})) + \kappa^2(\partial_5e^{j}{}_{\nu})e_{j}{}^{\nu}(g^{\mu\rho}A_{\mu}(\partial_5A_{\rho})) , \\ \bar{T}^{i}{}_{i\bar{5}}\bar{T}^{j}{}_{j}{}^{\bar{5}} &= \frac{1}{\phi^2}\eta^{5\bar{5}}(\partial_5e^{i}{}_{\mu})e_{i}{}^{\mu}(\partial_5e^{j}{}_{\nu})e_{j}{}^{\nu} . \end{split}$$

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Kaluza-Klein (KK) Theory

KK ansatz:

- Cylindrical condition (no y dependency)
- Compactification: $G = S^1$ and only consider zero KK mode



- The manifold is $M_4 \times S^1$ $(y = r \theta)$
- Harmonic expansions

$$\begin{split} e^{i}{}_{\mu}(x,y) &= \sum_{n} e^{i}{}_{\mu}^{(n)}(x) \, e^{iny/2r} & \Longrightarrow \quad g_{\mu\nu}(x,y) = \sum_{n} g^{(n)}_{\mu\nu}(x) \, e^{iny/r} \,, \\ A_{\mu}(x,y) &= \sum_{n} A^{(n)}_{\mu}(x) \, e^{iny/r} \,, \quad \phi(x,y) = \sum_{n} \phi^{(n)}(x) \, e^{iny/r} \,. \end{split}$$

KK Zero mode

 \blacksquare n = 0 mode

$$\begin{cases} \bar{T}_{\rm NGR} = T_{\rm NGR}^{(0)} \,, \\ \bar{T}_{\hat{5}ij} \bar{T}^{\hat{5}ij} = \frac{\kappa^2}{4} \phi^{(0)\,2} \,\eta_{\hat{5}\hat{5}} g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu}^{(0)} F_{\rho\sigma}^{(0)} \,, \\ \bar{T}_{\hat{5}i\hat{5}} \bar{T}^{\hat{5}i\hat{5}} = \frac{1}{\phi^{(0)\,2}} \eta_{\hat{5}\hat{5}} \eta^{\hat{5}\hat{5}} (g^{\mu\nu} (\partial_{\mu}\phi^{(0)}) (\partial_{\nu}\phi^{(0)})) \,, \\ \bar{T}^j{}_j{}^i \bar{T}^{\hat{5}}{}_{\hat{5}}{}^i = -\frac{1}{\phi^{(0)}} \, T^{(0)\,\rho} (\partial_{\rho}\phi^{(0)}) \,. \end{cases}$$

$$\bullet \ \bar{T}_{i\hat{5}j}\bar{T}^{i\hat{5}j} = \bar{T}_{i\hat{5}j}\bar{T}^{j\hat{5}i} = \bar{T}_{\hat{5}ij}\bar{T}^{ji\hat{5}} = \bar{T}^{i}{}_{i\hat{5}}\bar{T}^{j}{}_{j}{}^{\hat{5}} = 0$$

Zero KK mode in 4D effective extended gravity in teleparallism

$$\begin{split} S_{\rm eff} &= \int d^4 x \, e \bigg(\frac{1}{2\kappa_4} \phi \, T_{\rm NGR} - \frac{a}{8} \phi^3 \, g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} \\ &+ \frac{2a+b+c}{2\kappa_4} \frac{1}{\phi} g^{\mu\nu} \partial_\mu \phi \partial_\nu \phi - \frac{c}{\kappa_4} \, T^\mu \, \partial_\mu \phi \bigg) \,. \end{split}$$

Weyl

By considering 5D matter Lagrangian (Universal Extra Dimensions, UED) ${}^{(5)}\mathcal{L}_{m} = {}^{(5)}e L_{m}(e^{I}{}_{M}, \Psi, \mathcal{D}_{M}\Psi)$

• The *n*-mode harmonic expansion of Ψ is given by

$$\Psi(x^{\mu}, y) = \sum_{n} \Psi(x^{\mu}) e^{iny/r} ,$$

For the massless zero mode, the matter field $\Psi^{(0)}$ is assumed to be localized on the T_4 hypersurface, the resulting effective matter Lagrangian is

$$\mathcal{L}_{\mathrm{m, eff}} = e\lambda\phi L_{\mathrm{m}}(e^{i}{}_{\mu}, A_{\mu}, \phi, \Psi, \mathcal{D}_{\mu}\Psi)\,.$$

Note:

For 4D localized matter, the Lagrangian can be identified as

$$^{(4)}\mathcal{L}_{\mathsf{m}} = \mathcal{L}_{\mathsf{m,eff}} \Big|_{\lambda=1,\,A_{\mu}=0,\,\phi=1} = eL_{\mathsf{m}}(e^{i}{}_{\mu},\Psi,\mathcal{D}_{\mu}\Psi)\,.$$

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Conformal Transformations

 \blacksquare Considering the conformal transformations Ω^λ with weight λ

$$\widetilde{e}^{i}{}_{\mu} = \Omega \, e^{i}{}_{\mu} \quad (\lambda_{e} = 1) \quad \Longrightarrow \quad \widetilde{g}_{\mu\nu} = \Omega^{2}(x) \, g_{\mu\nu} \quad (\lambda_{g} = 2) \,,$$

Transformation of the torsion tensor and vector

$$\begin{split} \widetilde{T}^{i}{}_{\mu\nu} &= \Omega \, T^{i}{}_{\mu\nu} + \left(\partial_{\mu}\Omega\right) e^{i}{}_{\nu} - e^{i}{}_{\mu} \left(\partial_{\nu}\Omega\right), \\ \widetilde{T}_{\mu} &= T_{\mu} - 3 \, \Omega^{-1} \, \partial_{\mu}\Omega \,, \end{split}$$

■ Torsion scalar of NGR,

$$\begin{split} T_{\mathrm{NGR}} &= \Omega^2 \, \widetilde{T}_{\mathrm{NGR}} + \left(4 \, a + 2 \, b + 6 \, c \right) \widetilde{g}^{\mu\nu} \widetilde{T}_{\mu}(\Omega \, \partial_{\nu} \Omega) \\ &+ \left(6 \, a + 3 \, b + 9 \, c \right) \widetilde{g}^{\mu\nu} \partial_{\mu} \Omega \, \partial_{\nu} \Omega \, . \end{split}$$

• Scalar
$$\phi$$
: $\tilde{\phi} = \Omega^{\lambda_{\phi}} \phi$ with $\lambda_{\phi} = -2$
• Vector A_{μ} : $\tilde{A}_{\mu} = A_{\mu}$ and $\tilde{F}_{\mu\nu} = F_{\mu\nu}$

Transformed effective Lagrangian density

$$\begin{aligned} \mathcal{L}_{g} &= \widetilde{e} \frac{1}{2\kappa_{4}} \left\{ \widetilde{\phi} \, \widetilde{T}_{\mathsf{NGR}} + \left(4\,a + 2\,b + 2\,c \right) \widetilde{\phi} \, \widetilde{g}^{\mu\nu} \, \widetilde{T}_{\mu} \, \partial_{\nu} \omega \right. \\ &+ \left(14\,a + 7\,b + c \right) \widetilde{\phi} \, \widetilde{g}^{\mu\nu} \, \partial_{\mu} \omega \, \partial_{\nu} \omega \\ &- e^{6\omega} \, \frac{a\kappa^{2}}{4} \widetilde{\phi}^{3} \, \widetilde{g}^{\mu\rho} \widetilde{g}^{\nu\sigma} \, \widetilde{F}_{\mu\nu} \widetilde{F}_{\rho\sigma} + \left(2a + b + c \right) \frac{1}{\widetilde{\phi}} \, \widetilde{g}^{\mu\nu} \partial_{\mu} \widetilde{\phi} \, \partial_{\nu} \widetilde{\phi} \\ &- 2\,c \, \widetilde{g}^{\mu\nu} \, \widetilde{T}_{\mu} \, \partial_{\nu} \widetilde{\phi} + \left(8\,a + 4\,b - 2\,c \right) \widetilde{g}^{\mu\nu} \, \partial_{\mu} \omega \, \partial_{\nu} \widetilde{\phi} \right\}, \quad (*) \end{aligned}$$

with the conformal scalar $\omega := \ln \Omega$.

Conformal 1

<u>The Existence Einstein-Frame</u>

- In general, the Einstein-frame does not exist for the non-minimal torsion scalar $\phi T_{\rm NGR}$.
- The non-minimal coupling $\tilde{g}^{\mu\nu}\tilde{T}_{\mu}(\Omega \partial_{\nu}\Omega)$ will be always generated.
- The Einstein-frame is obtained by eliminating the term $\tilde{g}^{\mu\nu}\tilde{T}_{\mu}\partial_{\nu}\omega$ in the transformed Lagrangian.

Necessary condition: 2a + b + c = 0.

The Lagrangian is reduced to

$$\mathcal{L}_g = e \frac{1}{2\kappa_4} \left(\phi T_{\mathsf{NGR}} - 2 c g^{\mu\nu} T_\mu \partial_\nu \phi - \frac{a\kappa^2}{4} \phi^3 g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} \right).$$

The corresponding Einstein-frame

$$\mathcal{L}_{g}^{(\mathsf{E})} = \widetilde{e} \left(\frac{1}{2\kappa_{4}} \widetilde{T}_{\mathsf{NGR}} - \frac{c}{2} \, \widetilde{g}^{\mu\nu} \partial_{\mu} \varphi \, \partial_{\nu} \varphi \right. \\ \left. - \frac{a\kappa^{2}}{8\kappa_{4}} \, e^{6\omega} \, \widetilde{g}^{\mu\rho} \widetilde{g}^{\nu\sigma} \widetilde{F}_{\mu\nu} \widetilde{F}_{\rho\sigma} \right),$$

where where $\varphi := \sqrt{6/\kappa_4} \, \omega$ and the ghost-free condition is $c \leq 0$.

Einstein-frame condition

$$2a+b+c=0 \quad \text{and} \quad c \le 0.$$

■ For a simple choice of *c* = −1, one gets the minimal coupled one-parameter family model with 2*a* + *b* = 1 in teleparallelism.

Conformal Invariant Gravity

• Only keeping the terms of $\widetilde{\phi}^{-1} \widetilde{g}^{\mu\nu} \partial_{\mu} \widetilde{\phi} \partial_{\nu} \widetilde{\phi}$ and $\widetilde{g}^{\mu\nu} \widetilde{T}_{\mu} \partial_{\nu} \widetilde{\phi}$ · Lagrangian

$$\begin{split} & 4a + 2b + 2c = 0 \,, \\ & 14a + 7b + c = 0 \,, \\ & 8a + 4b - 2c = 0 \,. \end{split}$$

Conformal Invariant condition2a + b = 0and

 Corresponding to the simple one-parameter conformal invariant gravity in teleparallelism

$$S_g^{(\mathsf{c})} = \int d^4x \left\{ e \frac{a}{2\kappa_4} \phi \left(T_{ijk} T^{ijk} - 2 T_{ijk} T^{kji} \right) \right\}.$$

The gravitational equation of motion

$$\phi \left\{ \frac{1}{2} e_i^{\mu} T^j{}_{\rho\nu} \left(T_j^{\rho\nu} - 2 T^{\nu\rho}{}_j \right) - 4 e_i^{\rho} T^j{}_{\rho\nu} K^{\mu\nu}{}_j \right\}$$

+
$$\frac{2\phi}{e} \partial_{\nu} \left\{ e \left(T_i^{\mu\nu} - 2 T^{\nu\mu}{}_i \right) \right\} + 2 \left(T_i^{\mu\nu} - 2 T^{\nu\mu}{}_i \right) \partial_{\nu} \phi = 0.$$

 \blacksquare The equation of motion of ϕ

$$e \frac{a}{2\kappa_4} T_{ijk} (T^{ijk} - 2 T^{kji}) = 0.$$

- $\begin{array}{l} \bullet \ T_{ijk} = 0 \implies \text{no gravity} \implies \overbrace{\text{forbidden}!!} \\ \bullet \ T^{ijk} = 2 \ T^{kji} \implies \text{it implies } T^{iik} = 2T^{kii} = 0, \ \text{no torsion vector}!! \\ \implies NO \ \text{new interaction}!! \end{array}$
- The gravitational equation is reduced to

$$e_i^{\rho} T^j{}_{\rho\nu} T^{\nu\mu}{}_j = 0 \,.$$

Weyl Gauge Invariance

- The torsion vector T_{μ} is identified as the gauge field.
- Rewriting the effective Lagrangian

$$\begin{split} \mathcal{L}_g &= \, e \frac{1}{2\kappa_4} \bigg\{ \phi \left(T_{\mathrm{NGR}} - kc \, g^{\mu\nu} \, T_\mu T_\nu \right) \\ &+ \frac{c}{k\phi} \left(g^{\mu\nu} (\partial_\mu - kT_\mu) \phi (\partial_\nu - kT_\nu) \phi \right) \bigg\} \,, \end{split}$$
 where $k = \frac{c}{2a + b + c}$ is a fixed ratio.

- We need *ghost-free* 2a + b + c > 0 and $c \neq 0$
- Conformal transformation

$$g_{\mu\nu} \longrightarrow e^{2\omega}g_{\mu\nu}, \qquad T_{\mu} \longrightarrow T_{\mu} - 3\,\partial_{\mu}\omega, \qquad \phi \longrightarrow e^{-2\omega}\,\phi.$$

Under the conformal transformation, the effective Lagrangian becomes as

$$\begin{split} \mathcal{L}_{g} &= \underbrace{\tilde{e}}_{e \, e^{4\omega}} \frac{1}{2\kappa_{4}} \left\{ \widetilde{\phi} \, \widetilde{T}_{\mathsf{NGR}} - kc \, \widetilde{\phi} \, \widetilde{g}^{\mu\nu} \, \widetilde{T}_{\mu} \widetilde{T}_{\nu} + \left(\frac{1}{k} + 2 - 3k \right) 2c \, \widetilde{\phi} \, \widetilde{g}^{\mu\nu} \, \widetilde{T}_{\mu} \, \partial_{\nu} \omega \\ &+ \left(\frac{1}{k} + 2 - 3k \right) 3c \, \widetilde{\phi} \, \widetilde{g}^{\mu\nu} \, \partial_{\mu} \omega \, \partial_{\nu} \omega \right. - e^{6\omega} \frac{a\kappa^{2}}{4} \widetilde{\phi}^{3} \, \widetilde{g}^{\mu\rho} \widetilde{g}^{\nu\sigma} \, \widetilde{F}_{\mu\nu} \, \widetilde{F}_{\rho\sigma} \\ &+ \underbrace{\frac{c}{k\widetilde{\phi}} \left[\widetilde{g}^{\mu\nu} \left(\partial_{\mu} - k\widetilde{T}_{\mu} + (2 - 3k) \partial_{\mu} \omega \right) \widetilde{\phi} \left(\partial_{\nu} - k\widetilde{T}_{\nu} + (2 - 3k) \partial_{\nu} \omega \right) \widetilde{\phi} \right] \right\} . \\ &\left. e^{-4\omega} \frac{c}{k\phi} \left(g^{\mu\nu} (\partial_{\mu} - kT_{\mu}) \phi \left(\partial_{\nu} - kT_{\nu} \right) \phi \right) \\ &= \frac{1}{k} + 2 - 3k = 0 \Longrightarrow \text{ conformal invariance}!! \end{split}$$

Conformal Invariant condition

$$\begin{cases} 2a+b+4c=0 & \text{for } k=-\frac{1}{3} \\ \text{or} \\ 2a+b=0 & \text{for } k=1 \end{cases}$$

with 2a + b + c > 0 and $c \neq 0$

• Weyl derivative for general field ψ :

$$^*\partial_{\mu}^{(\psi)} = \partial_{\mu} + \frac{\lambda_{\psi}k}{2}T_{\mu}$$

 \blacksquare Weyl derivative for $\phi,\,e^i{}_\nu$ and $g_{\nu\rho}$

$$\begin{cases} *\partial_{\mu}^{(\phi)}\phi = \left(\partial_{\mu} + \frac{\lambda_{\phi}k}{2}T_{\mu}\right)\phi & \text{with } \lambda_{\phi} = -2\\ *\partial_{\mu}^{(e)}e^{i}{}_{\nu} = \left(\partial_{\mu} + \frac{\lambda_{e}k}{2}T_{\mu}\right)e^{i}{}_{\nu} & \text{with } \lambda_{e} = 1\\ *\partial_{\mu}^{(g)}g_{\nu\rho} = \left(\partial_{\mu} + \frac{\lambda_{g}k}{2}T_{\mu}\right)g_{\nu\rho} & \text{with } \lambda_{g} = 2 \end{cases}$$

• Define an invariant connection ${}^*\Gamma^{\rho}_{\nu\mu} = e_i{}^{\rho}{}^*\partial^{(e)}_{\mu}e^i{}_{\nu} = \Gamma^{\rho}_{\nu\mu} + \frac{k}{2}\delta^{\rho}_{\nu}T_{\mu}.$

$$\implies {}^*\partial^{(g)}_{\mu}g_{\nu\rho} - \Gamma^{\sigma}_{\nu\mu}g_{\sigma\rho} - \Gamma^{\sigma}_{\rho\mu}g_{\nu\sigma} = \frac{\lambda_g k}{2}T_{\mu}g_{\nu\rho}$$
$$\implies {}^*R^{\rho}_{\sigma\mu\nu} = \frac{k}{2}\delta^{\rho}_{\sigma}(\partial_{\mu}T_{\nu} - \partial_{\nu}T_{\mu}) \neq 0$$
$$\implies \text{Weyl-Cartan geometry!!!}$$

 \blacksquare The modified covariant derivative for ψ given by

$$^{*}\nabla^{(\psi)} = ^{*}d^{(\psi)} + ^{*}\Gamma.$$

The nonmetricity vanishes

$${}^*\nabla^{(g)}_{\mu}g_{\nu\rho} = {}^*\partial^{(g)}_{\mu}g_{\nu\rho} - {}^*\Gamma^{\sigma}_{\nu\mu}g_{\sigma\rho} - {}^*\Gamma^{\sigma}_{\rho\mu}g_{\nu\sigma} = \nabla_{\mu}g_{\nu\rho} = 0.$$

 \blacksquare We define ${}^*\widetilde{T}{}^{\rho}{}_{\mu\nu}:=\widetilde{({}^*T)}{}^{\rho}{}_{\mu\nu},$ we have

$$\begin{cases} {}^{*}T^{\rho}{}_{\mu\nu} = T^{\rho}{}_{\mu\nu} + \frac{k}{2} (\delta^{\rho}{}_{\nu}T_{\mu} - \delta^{\rho}{}_{\mu}T_{\nu}) ,\\ {}^{*}T_{\mu} = (1 - \frac{3}{2}k) T_{\mu} , \end{cases} \quad \text{and} \quad \begin{cases} {}^{*}\widetilde{T}^{\rho}{}_{\mu\nu} = \widetilde{T}^{\rho}{}_{\mu\nu} + \frac{k}{2} (\delta^{\rho}{}_{\nu}\widetilde{T}_{\mu} - \delta^{\rho}{}_{\mu}\widetilde{T}_{\nu}) ,\\ {}^{*}\widetilde{T}_{\mu} = (1 - \frac{3}{2}k) \widetilde{T}_{\mu} . \end{cases}$$

•
$$^*\partial^{(\psi)}_{\mu}$$
 in terms of $^*T_{\mu} \Longrightarrow ^*\partial^{(\psi)}_{\mu} = \partial_{\mu} - \lambda_{\psi}k^{2*}T_{\mu}$.

• Due to
$${}^*\widetilde{T}^{\rho}{}_{\mu\nu} = {}^*T^{\rho}{}_{\mu\nu} - (1/2)(\delta^{\rho}_{\nu}\partial_{\mu}\omega - \delta^{\rho}_{\mu}\partial_{\nu}\omega)$$
, we have

$$e \phi \left(T_{\text{NGR}} - kc g^{\mu\nu} T_{\mu} T_{\nu} \right)$$
$$= e \phi \left({}^{*}T_{\text{NGR}} - kc g^{\mu\nu} {}^{*}T_{\mu} {}^{*}T_{\nu} \right)$$
$$= \widetilde{e} \widetilde{\phi} \left({}^{*}\widetilde{T}_{\text{NGR}} - kc \widetilde{g}^{\mu\nu} {}^{*}\widetilde{T}_{\mu} {}^{*}\widetilde{T}_{\nu} \right)$$

and

Weyl gauge invariance

$$\mathcal{L}_{g} = \tilde{e} \frac{1}{2\kappa_{4}} \left\{ \tilde{\phi} \left({}^{*} \widetilde{T}_{\mathsf{NGR}} - kc \, \widetilde{g}^{\mu\nu} \, {}^{*} \widetilde{T}_{\mu} \, {}^{*} \widetilde{T}_{\nu} \right) + \frac{c}{k \widetilde{\phi}} \, \widetilde{g}^{\mu\nu} \, {}^{*} \widetilde{\nabla}_{\mu}^{(\phi)} \widetilde{\phi} \, {}^{*} \widetilde{\nabla}_{\nu}^{(\phi)} \widetilde{\phi} \right\},$$

<u>Outline</u>

1 Teleparallel Gravity

- 2 Five-Dimensional Geometry
- 3 Kaluza-Klein Theory
- 4 Specific Models
- 5 Weak Field Approximation

6 Summary

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Weak Field Approximation

• Define a canonical field $\Phi := \sqrt{\phi/\kappa_4}$,

$$\begin{split} S &= S_g + S_{\mathsf{m}} \\ &= \int d^4 x \left\{ e \bigg(\frac{1}{2} \, \Phi^2 \, T_{\mathsf{NGR}} - \frac{a \kappa^2 \kappa_4^2}{8} \, \Phi^6 \, g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} \right. \\ &\left. + \left(4a + 2b + 2c \right) g^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi - 2c \, g^{\mu\nu} \, T_\mu \, \Phi \partial_\nu \Phi \bigg) + \kappa_4 \lambda \Phi^2 \mathcal{L}_{\mathsf{m}} \bigg\} \,. \end{split}$$

■ In the weak field approximation $e^{i}{}_{\mu} = \delta^{i}_{\mu} + h^{i}{}_{\mu} (e_{i}{}^{\mu} = \delta^{\mu}_{i} - h_{i}{}^{\mu})$, the tensor $h_{\mu\nu}$ contains the anti-symmetric fluctuations:

$$h_{\mu\nu} = \underbrace{\frac{1}{2}\gamma_{\mu\nu}}_{\text{symmetric}} + \underbrace{a_{\mu\nu}}_{\text{anti-symmetric}} \text{ and } \left|h^{i}_{\mu}\right| \ll 1 \,.$$

• The metric tensor $g_{\mu\nu} = \eta_{ij} e^i{}_{\mu} e^j{}_{\nu} \approx \eta_{\mu\nu} + \gamma_{\mu\nu}$ contains no anti-symmetric part of $h_{\mu\nu}$.

The torsion tensor and torsion vector are

$$\begin{cases} T^{\rho}{}_{\mu\nu} = \delta^{\rho}_{i}(\partial_{\mu}h^{i}{}_{\nu} - \partial_{\nu}h^{i}{}_{\mu}) + \mathcal{O}(h^{2}_{\mu\nu}), \\ T_{\nu} = \partial_{\mu}h^{\mu}{}_{\nu} - \partial_{\nu}h + \mathcal{O}(h^{2}_{\mu\nu}). \end{cases}$$

• $e = 1 + h + \mathcal{O}(h_{\mu\nu}^2)$ with $h \equiv \delta_i^{\mu} h^i{}_{\mu} = h^{\mu}{}_{\mu} = (1/2)\gamma$.

Lagrangian in the lowest order

$$\begin{aligned} \mathcal{L}_g \approx &\frac{1}{2} \, \Phi^2 \, T_{\mathsf{NGR}} - \frac{a \kappa^2 \kappa_4^2}{8} \, \Phi^6 \eta^{\mu\rho} \eta^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} \\ &+ \left(4a + 2b + 2c\right) \eta^{\mu\nu} \partial_\mu \Phi \partial_\nu \Phi - \frac{2c}{2} \eta^{\mu\nu} T_\mu \, \Phi \partial_\nu \Phi \end{aligned}$$

• The current-vector interaction $\eta^{\mu\nu}T_{\mu}\Phi\partial_{\nu}\Phi$:

$$T_{\mu} \approx (1/2)\partial_{\rho}\gamma^{\rho}{}_{\mu} + \frac{\partial_{\rho}a^{\rho}{}_{\mu}}{-(1/2)\partial_{\mu}\gamma}$$

$$\Phi \partial_{\mu}\Phi \longrightarrow \int d^{4}x - 2c \partial_{\rho}a^{\rho\nu} \Phi \partial_{\nu}\Phi = \int d^{4}x c \Phi^{2} \frac{\partial_{\nu}\partial_{\rho}}{\partial_{\rho}a^{\rho\nu}} \longrightarrow 0$$
symmetric

No contribution from $a_{\mu\nu}$

$$T_{\text{NGR}} = \frac{1}{4} \left((2a+b)\partial_{\mu}\gamma_{\nu\rho}\partial^{\mu}\gamma^{\nu\rho} - (2a+b)\partial_{\mu}\gamma_{\nu\rho}\partial^{\rho}\gamma^{\mu\nu} + c\,\partial^{\rho}\gamma_{\rho\mu}\partial_{\sigma}\gamma^{\sigma\mu} - 2c\,\partial_{\mu}\gamma\,\partial_{\rho}\gamma^{\rho\mu} + c\,\partial_{\mu}\gamma\,\partial^{\mu}\gamma \right) \\ + (2a+b)\partial_{\mu}\gamma_{\nu\rho}\partial^{\nu}a^{\mu\rho} + c\,\partial^{\rho}\gamma_{\rho\mu}\partial_{\sigma}a^{\sigma\mu} - c\,\partial_{\mu}\gamma\,\partial_{\rho}a^{\rho\mu} \\ + (2a-b)\partial_{\mu}a_{\nu\rho}\partial^{\mu}a^{\nu\rho} + (2a-3b)\partial_{\mu}a_{\nu\rho}\partial^{\rho}a^{\mu\nu} + c\,\partial^{\rho}a_{\rho\mu}\partial_{\sigma}a^{\sigma\mu} \right)$$

<u>Note</u>: T_{NGR} becomes the well-known Fierz-Pauli Lagrangian (*Fierz and Pauli, 1939*) by setting $(a, b, c) = (\frac{1}{4}, \frac{1}{2}, -1)$ and $a_{\mu\nu} = 0$.

Gauge conditions

$$\begin{cases} \partial_{\mu}\gamma^{\mu}{}_{\nu}=0 & \text{transversed condition,} \\ \partial_{\mu}a^{\mu}{}_{\nu}=0 & \text{transversed condition,} \Longrightarrow \text{torsion vector } T_{\mu} \text{ vanishes.} \\ \gamma=0 & \text{traceless condition.} \end{cases}$$

 \blacksquare We define $j_{\mu}:=\Phi\partial_{\mu}\Phi$

$$\text{and} \quad \begin{cases} T_{\gamma}^{\mu\nu} := \frac{1}{2} T^{(\mu\nu)} = -2 \frac{\delta \mathcal{L}_{\mathsf{m}}}{\delta \gamma_{\mu\nu}} \,, \\ \\ T_{a}^{\mu\nu} := T^{[\mu\nu]} = -2 \frac{\delta \mathcal{L}_{\mathsf{m}}}{\delta a_{\mu\nu}} \,. \end{cases}$$

EoM

$$\begin{aligned} \bullet \quad & \text{EoM of } \gamma_{\mu\nu}: \\ & -j_{\rho} \Big\{ \frac{2a+b}{2} \partial^{\rho} \gamma^{\mu\nu} - \frac{2a+b}{4} \Big(\partial^{\nu} \gamma^{\rho\mu} + \partial^{\mu} \gamma^{\rho\nu} \Big) \\ & + \frac{c}{4} \left(\eta^{\rho\mu} \partial_{\sigma} \gamma^{\sigma\nu} + \eta^{\rho\nu} \partial_{\sigma} \gamma^{\sigma\mu} \right) \\ & - \frac{c}{2} \left(\frac{1}{2} \eta^{\rho\mu} \partial^{\nu} \gamma + \frac{1}{2} \eta^{\rho\nu} \partial^{\mu} \gamma + \eta^{\mu\nu} \partial_{\sigma} \gamma^{\sigma\rho} \right) \\ & + \frac{c}{2} \eta^{\mu\nu} \partial^{\rho} \gamma + \frac{2a+b}{2} \left(\partial^{\mu} a^{\rho\nu} + \partial^{\nu} a^{\rho\mu} \right) \\ & + \frac{c}{2} \left(\eta^{\rho\mu} \partial_{\sigma} \alpha^{\sigma\nu} + \eta^{\rho\nu} \partial_{\sigma} \alpha^{\sigma\mu} \right) - c \eta^{\mu\nu} \partial_{\sigma} a^{\sigma\rho} \Big\} \\ & - \Phi^{2} \Big\{ \frac{2a+b}{4} \Box \gamma^{\mu\nu} - \frac{2a+b}{8} \left(\partial_{\rho} \partial^{\nu} \gamma^{\mu\rho} + \partial_{\rho} \partial^{\mu} \gamma^{\nu\rho} \right) \\ & + \frac{c}{8} \left(\partial^{\mu} \partial_{\sigma} \gamma^{\sigma\nu} + \partial^{\nu} \partial_{\sigma} \gamma^{\sigma\mu} \right) - \frac{c}{4} \left(\partial^{\mu} \partial^{\nu} \gamma + \partial_{\rho} \partial_{\sigma} \gamma^{\sigma\rho} \eta^{\mu\nu} \right) \\ & + \frac{c}{4} \eta^{\mu\nu} \Box \gamma + \frac{2a+b}{4} \left(\partial_{\rho} \partial^{\mu} a^{\rho\nu} + \partial_{\rho} \partial^{\nu} a^{\rho\mu} \right) \\ & + \frac{c}{4} \left(\partial^{\mu} \partial_{\sigma} a^{\sigma\nu} + \partial^{\nu} \partial_{\sigma} a^{\sigma\mu} \right) \Big\} + \frac{c}{2} \partial^{\mu} j^{\nu} + \frac{c}{2} \partial^{\nu} j^{\mu} - c \partial_{\rho} j^{\rho} \eta^{\mu\nu} = \frac{1}{2} \kappa_{4} \lambda \Phi^{2} T_{\gamma}^{\mu\nu} , \\ \text{where } \Box := \eta^{\mu\nu} \partial_{\mu} \partial_{\nu}. \end{aligned}$$

,

• EoM of $a_{\mu\nu}$:

$$\begin{split} -j_{\rho} \bigg\{ \frac{2a+b}{2} \bigg(\partial^{\mu} \gamma^{\rho\nu} - \partial^{\nu} \gamma^{\rho\mu} \bigg) + \frac{c}{2} \bigg(\eta^{\rho\mu} \partial_{\sigma} \gamma^{\sigma\nu} - \eta^{\rho\nu} \partial_{\sigma} \gamma^{\sigma\mu} \bigg) \\ &- \frac{c}{2} \bigg(\eta^{\rho\mu} \partial^{\nu} \gamma - \eta^{\rho\nu} \partial^{\mu} \gamma \bigg) + 2 \bigg(2a-b \bigg) \partial^{\rho} a^{\mu\nu} \\ &+ \bigg(2a-3b \bigg) \bigg(\partial^{\mu} a^{\nu\rho} + \partial^{\nu} a^{\rho\mu} \bigg) + c \bigg(\eta^{\rho\mu} \partial_{\sigma} a^{\sigma\nu} - \eta^{\rho\nu} \partial_{\sigma} a^{\sigma\mu} \bigg) \bigg\} \\ &- \Phi^{2} \bigg\{ \frac{2a+b}{4} \bigg(\partial_{\rho} \partial^{\mu} \gamma^{\rho\nu} - \partial_{\rho} \partial^{\nu} \gamma^{\rho\mu} \bigg) + \frac{c}{4} \bigg(\partial^{\mu} \partial_{\sigma} \gamma^{\sigma\nu} - \partial^{\nu} \partial_{\sigma} \gamma^{\sigma\mu} \bigg) \\ &+ \bigg(2a-b \bigg) \Box a^{\mu\nu} + \frac{2a-3b-c}{2} \bigg(\partial_{\rho} \partial^{\mu} a^{\nu\rho} + \partial_{\rho} \partial^{\nu} a^{\rho\mu} \bigg) \bigg\} = \frac{1}{2} \kappa_{4} \lambda \Phi^{2} T_{a}^{\mu\nu} \, . \end{split}$$

• EoM of A_{μ} :

$$\kappa_4 \lambda \Phi^2 \frac{\delta \mathcal{L}_{\mathsf{m}}}{\delta A_{\mu}} + 3a\kappa^2 \kappa_4^2 \Phi^5(\partial_{\nu} \Phi) F^{\mu\nu} + \frac{a\kappa^2 \kappa_4^2}{2} \Phi^6 \partial_{\nu} F^{\mu\nu} = 0.$$

• EoM of Φ :

$$\begin{split} \Phi T_{\rm NGR} &- \frac{3a\kappa^2\kappa_4^2}{4} \Phi^5 F_{\mu\nu} F^{\mu\nu} + 2\lambda\kappa_4 \Phi \mathcal{L}_{\rm m} \\ + \kappa_4 \lambda \Phi^2 \frac{\delta \mathcal{L}_{\rm m}}{\delta \Phi} &- \left(8a + 4b + 4c\right) \Box \Phi + 2c \Phi \partial_\mu T^\mu = 0 \,. \end{split}$$

We consider the case of Weyl gauge invariance

$$2a + b + 4c = 0$$
 or $2a + b = 0$

along with the gauge conditions $\partial_{\mu}\gamma^{\mu\nu} = 0$, $\partial_{\mu}a^{\mu\nu} = 0$ and $\gamma = 0$ • The EoM of $\gamma_{\mu\nu}$ and $a_{\mu\nu}$ for 2a + b + 4c = 0

$$\begin{cases} cj_{\rho} \left\{ 2\partial^{\rho}\gamma^{\mu\nu} - \partial^{\mu}\gamma^{\rho\nu} - \partial^{\nu}\gamma^{\rho\mu} + 2\partial^{\mu}a^{\rho\nu} + 2\partial^{\nu}a^{\rho\mu} \right\} \\ +c \Phi^{2} \Box \gamma^{\mu\nu} + \frac{c}{2} \partial^{\mu}j^{\nu} + \frac{c}{2} \partial^{\nu}j^{\mu} - c \partial_{\rho}j^{\rho}\eta^{\mu\nu} = \frac{1}{2}\kappa_{4}\lambda\Phi^{2}T_{\gamma}^{\mu\nu}, \\ j_{\rho} \left\{ 2c \left(\partial^{\mu}\gamma^{\rho\nu} - \partial^{\nu}\gamma^{\rho\mu} \right) + 4(b+2c)\partial^{\rho}a^{\mu\nu} \\ +4\left(b+c\right) \left(\partial^{\mu}a^{\nu\rho} + \partial^{\nu}a^{\rho\mu} \right) \right\} + 2\left(b+2c\right)\Phi^{2} \Box a^{\mu\nu} = \frac{1}{2}\kappa_{4}\lambda\Phi^{2}T_{a}^{\mu\nu}. \end{cases}$$

Assuming that the scalar field varies slowly

 $\Phi \approx \Phi_c$ a constant field, $\implies j_\mu \approx 0$,

■ The eqs. reduce to

$$\Box \gamma^{\mu\nu} = \frac{\kappa_4 \lambda}{2c} T^{\mu\nu}_{\gamma} ,$$
$$\Box a^{\mu\nu} = \frac{\kappa_4 \lambda}{4(b+2c)} T^{\mu\nu}_a$$

• The EoM of $\gamma_{\mu\nu}$ and $a_{\mu\nu}$ for 2a + b = 0

$$\begin{cases} \frac{c}{2} \partial^{\mu} j^{\nu} + \frac{c}{2} \partial^{\nu} j^{\mu} - c \partial_{\rho} j^{\rho} \eta^{\mu\nu} = \frac{1}{2} \kappa_4 \lambda \Phi^2 T_{\gamma}^{\mu\nu} ,\\ -8a j_{\rho} f^{\rho\mu\nu} - 4a \Phi^2 \Box a^{\mu\nu} = \frac{1}{2} \kappa_4 \lambda \Phi^2 T_a^{\mu\nu} , \end{cases}$$

where $f^{\rho\mu\nu} := \partial^{\rho}a^{\mu\nu} + \partial^{\mu}a^{\nu\rho} + \partial^{\nu}a^{\rho\mu}$ is the field strength of $a^{\mu\nu}$. • Only the anti-symmetric tensor $a^{\mu\nu}$ survives!!

■ For slowly varying scalar field, the eq. reduce to

$$\begin{cases} 0 = T_{\gamma}^{\mu\nu} \,, \\ \Box \, a^{\mu\nu} = -\frac{\kappa_4\lambda}{8a} \, T_a^{\mu\nu} \,. \end{cases}$$

<u>Outline</u>

1 Teleparallel Gravity

- 2 Five-Dimensional Geometry
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Summary

- We have summarized the possible choices of the coefficients (a, b, c) on the torsion scalar as shown by TABLE.
- The Einstein-frame can be achieved by taking 2a + b + c = 0 with $c \le 0$.
- We have obtained new classes of conformal invariant theories of gravity without the electromagnetic field A_μ.
- We provide a conformal invariant gravity in teleparallelism with the condition 2a + b = 0 with c = 0, which gives rise to the existence of the Einstein-frame.
- The Weyl gauge theory under the ghost-free constraints 2a + b + c > 0 and $c \neq 0$ can be obtained with the requirements either 2a + b + 4c = 0 or 2a + b = 0.
- For the conformal invariant models with 2a + b = 0, we found that only the anti-symmetric tensor field is allowed rather than the symmetric one.



Thank You!!!



7 Backup Slides

Ling-Wei Luo

Geometrical Meaning of Torsion

■ Torsion free: a tangent vector does not rotate when we parallel transport it. (*P.371, John Baez and Javier P. Muniain, "Gauge Fields, Knots and Gravity," 1994*)

$$T(u,v) = \nabla_u v - \nabla_v u - [u,v]$$

vanished in coordinate space



Notation in 5D

- In orthonormal frame, 5D metric is $\bar{g}_{MN} = \bar{\eta}_{IJ} e^{I}{}_{M} e^{J}{}_{N}$, $\bar{\eta}_{IJ} = \text{diag}(+1, -1, -1, -1, \varepsilon)$ with $\varepsilon = -1$.
- Coordinate frame $M, N = 0, 1, 2, 3, 5, \quad \mu, \nu = 0, 1, 2, 3, \quad \alpha, \beta = 1, 2, 3.$
- Orthonormal frame (anholonomic frame) $A, B, I, J, K = \hat{0}, \hat{1}, \hat{2}, \hat{3}, \hat{5}, \quad i, j, k = \hat{0}, \hat{1}, \hat{2}, \hat{3}, \quad a, b = \hat{1}, \hat{2}, \hat{3}.$

Affine Connection and Lorentz Connection

Consider noncoordinate basis (orthonormal frame)

$$\begin{split} e_{i}{}^{\nu}D_{\mu}V^{i} &= e_{i}{}^{\nu}(\partial_{\mu}V^{i} + \omega^{i}{}_{j\mu}V^{j}) \\ &= e_{i}{}^{\nu}\left(\partial_{\mu}(e^{i}{}_{\rho}V^{\rho}) + \omega^{i}{}_{j\mu}V^{j}\right) \\ &= e_{i}{}^{\nu}\left((\partial_{\mu}e^{i}{}_{\rho})V^{\rho} + e^{i}{}_{\rho}(\partial_{\mu}V^{\rho}) + \omega^{i}{}_{j\mu}e^{j}{}_{\rho}V^{\rho}\right) \\ &= (e_{i}{}^{\nu}\partial_{\mu}e^{i}{}_{\rho})V^{\rho} + \underbrace{\delta^{\nu}_{\rho}\partial_{\mu}V^{\rho}}_{\partial_{\mu}V^{\nu}} + e_{i}{}^{\nu}\omega^{i}{}_{j\mu}e^{j}{}_{\rho}V^{\rho} \\ &= \partial_{\mu}V^{\nu} + (e_{i}{}^{\nu}\partial_{\mu}e^{i}{}_{\rho} + e_{i}{}^{\nu}\omega^{i}{}_{j\mu}e^{j}{}_{\rho})V^{\rho} \\ &\equiv \partial_{\mu}V^{\nu} + \Gamma^{\nu}{}_{\rho\mu}V^{\rho} = \nabla_{\mu}V^{\nu}. \end{split}$$

The relation between affine connection and Lorentz connection

$$\Gamma^{\nu}{}_{\rho\mu} \equiv e_i{}^{\nu}\partial_{\mu}e^i{}_{\rho} + e_i^{\nu}\,\omega^i{}_{j\mu}\,e_{\rho}^j$$

• We can define the total covariant derivative ∇_{μ}

$$\partial_{\mu}e^{i}{}_{\rho} - \Gamma^{\nu}{}_{\rho\mu}e^{i}{}_{\nu} + \omega^{i}{}_{j\mu}e^{j}{}_{\rho} = 0$$
$$\Longrightarrow \nabla_{\mu}e^{i}{}_{\rho} = 0 \text{ (vielbein postulate).} \qquad \blacktriangleleft \text{Absolute } p$$

Brief History of 5-Dimensional Theories

- Kaluza-Klein (KK) theory: to unify electromagnetism and gravity by gauge theory
 - Cylindrical condition (Kaluza, 1921)
 - Compactification to a small scale (Klein, 1926)
- Generalization of KK: induced-matter theory ⇒ matter from the 5th-dimension (*Wesson*, 1998)
- Large Extra dimension (Arkani-Hamed, Dimopoulos and Dvali (ADD), 1998)
 - Solving hierarchy problem
 - SM particles confined on the 3-brane

- Randall-Sundrum model in AdS₅ spacetime (Randall and Sundrum, 1999)
 - RS-I (UV-brane and SM particles confined on IR-brane)
 ⇒ solving hierarchy problem
 - RS-II (only one UV brane)
 - \implies compactification to generate 4-dimensional gravity
- DGP brane model (*Dvali*, *Gabadadze and Porrati*, 2000) ⇒ accelerating universe
- Universal Extra Dimension (Appelquist, Cheng and Dobrescu, 2001)
 - Not only graviton but SM particles can propagate to the extra dimension ⇒ low compactification scale: reach to the electroweak scale

5D TEGR without Vector

In Gauss normal coordinate

$$\bar{g}_{MN} = \begin{pmatrix} g_{\mu\nu}(x^{\mu}, y) & 0\\ 0 & \varepsilon \phi^2(x^{\mu}, y) \end{pmatrix} \,.$$

■ The 5D torsion scalar in the orthonormal frame

$$\overset{(5)}{=} T = \underbrace{\bar{T}}_{i\,\hat{5}j} + \frac{1}{2} \left(\bar{T}_{i\hat{5}j} \, \bar{T}^{i\hat{5}j} + \bar{T}_{i\hat{5}j} \, \bar{T}^{j\hat{5}i} \right) + 2 \, \bar{T}^{j}{}_{j}{}^{i} \, \bar{T}^{\hat{5}}{}_{i\hat{5}} - \bar{T}^{j}{}_{\hat{5}j} \, \bar{T}^{k\hat{5}}{}_{k} \, .$$
 induced 4D torsion scalar

 \blacksquare The non-vanishing components of vielbein are $e^i{}_\mu$ and $e^{\hat{5}}{}_5$

Projection of the torsion tensor

$$\bar{T}^{\rho}{}_{\mu\nu} = T^{\rho}{}_{\mu\nu}$$
 (purely 4-dimensional object)



■ The 5D torsion scalar in the coordinate frame

$${}^{(5)}T = \bar{T} + \frac{1}{2} \left(\bar{T}_{\rho 5\nu} \, \bar{T}^{\rho 5\nu} + \bar{T}_{\rho 5\nu} \, \bar{T}^{\nu 5\rho} \right) + 2 \, \bar{T}^{\sigma}{}_{\sigma}{}^{\mu} \, \bar{T}^{5}{}_{\mu 5} - \bar{T}^{\nu}{}_{5\nu} \, \bar{T}^{\sigma 5}{}_{\sigma} \, .$$

Note:

In general, the induced torsion is $\bar{T}^{\rho}{}_{\mu\nu}=T^{\rho}{}_{\mu\nu}+\bar{C}^{\rho}{}_{\mu\nu}$, where

$$\bar{C}^{\rho}{}_{\mu\nu} = \bar{e}_{\hat{5}}{}^{\rho} (\partial_{\mu} e^{\hat{5}}{}_{\nu} - \partial_{\nu} e^{\hat{5}}{}_{\mu}) \,. \label{eq:constraint}$$

 $\bar{C}^{\hat{5}}{}_{\mu\nu} = \Gamma^{\hat{5}}{}_{\nu\mu} - \Gamma^{\hat{5}}{}_{\mu\nu} = h^{M}_{\mu} h^{N}_{\nu} T^{\hat{5}}{}_{MN} \sim \omega_{\mu\nu} \text{ is related to the extrinsic torsion or twist } \omega_{\mu\nu}.$

KK Reduction

The metric is reduced to

$$\bar{g}_{MN} = \begin{pmatrix} g_{\mu\nu}(x^{\mu}) & 0\\ 0 & -\phi^2(x^{\mu}) \end{pmatrix} \,.$$

• The residual components are $T^{\rho}_{\mu\nu}$ and $\bar{T}^{5}_{\mu5} = \frac{1}{\phi} \partial_{\mu} \phi$.

• The 5D torsion scalar with $\kappa_4 = \kappa_5/(2\pi r)$

$${}^{(5)}T = T + 2\,T^{\sigma}{}_{\sigma}{}^{\mu}\,\bar{T}^{5}{}_{\mu 5}$$

Effective action of 5D TEGR

$$S_{\text{eff}} = \frac{1}{2\kappa_4} \int d^4 x \, e \left(\phi \, T + 2 \, T^\mu \, \partial_\mu \phi \right)$$

(C.Q. Geng, LWL, H.H. Tseng, 2014)

Minimal and Non-Minimal Coupling

Minimal coupled case

$$T \sim -R$$
 (TEGR).

■ TEGR in 5D KK scenario with the metric given by

$$\bar{g}_{MN} = \begin{pmatrix} g_{\mu\nu} - k^2 A_{\mu} A_{\nu} & k A_{\mu} \\ k A_{\nu} & -1 \end{pmatrix} \text{ with } k^2 = \kappa_4,$$

The effective Lagrangian is

$$\mathcal{L}_{\rm eff} = e \left(\frac{1}{2\kappa_4} T - \frac{1}{4} F_{\mu\nu} F^{\mu\nu} \right) \quad \text{(coincides with the form of GR)}.$$

(de Andrade, Guillen, Pereira, 2000)

Non-minimal coupled case

$$\phi T \not\sim -\phi R$$

Remark:

The curvature-torsion relation in TEGR: $-\tilde{R}(e) = T - 2 \tilde{\nabla}_{\mu} T^{\mu}$.

5D GR vs. 5D TEGR in KK Scenario

The dimensional reduction of 5D GR

• Brans-Dicke theory with
$$\omega_{\rm BD} = 0$$

$$-\sqrt{-(5)g} {}^{(5)}\tilde{R} \longrightarrow -\sqrt{-g} \left(\phi \tilde{R} \underbrace{-\Box \phi}_{\text{surface term}} \right)$$

Remark:

Brans-Dicke theory (Brans & Dicke, 1961):

$$\int d^4x \sqrt{-g} \left\{ \frac{-1}{2\kappa} \phi \tilde{R} + \frac{\omega_{\rm BD}}{\phi} g^{\mu\nu} \partial_\mu \phi \, \partial_\nu \phi \right\}.$$

The dimensional reduction of 5D TEGR

$$^{(5)}e^{(5)}T \longrightarrow e\left(\phi T + \underbrace{2T^{\mu}\partial_{\mu}\phi}_{\text{no analogue}}\right)$$

• Substituting the relation $-\tilde{R}(e) = T - 2\tilde{\nabla}_{\mu}T^{\mu}$ into the 4D effective Lagrangian

Equivalence

$$\frac{-1}{2\kappa_4}\int d^4x\,e\!\left(\phi\tilde{R}(e)\underbrace{-2\,\tilde{\nabla}_{\mu}(\phi\,T^{\mu})}_{\text{surface term}}\right).$$

Einstein-frame

• By conformal transformation $(\tilde{g}_{\mu\nu} = \Omega^2(x) g_{\mu\nu})$:

$$\begin{split} T &= \Omega^2 \, \widetilde{T} - 4 \, \Omega \, \widetilde{g}^{\mu\nu} \, \widetilde{T}_{\mu} \partial_{\nu} \Omega + 6 \, \widetilde{g}^{\mu\nu} \, \partial_{\mu} \Omega \, \partial_{\nu} \Omega \,, \\ T_{\mu} &= \widetilde{T}_{\mu} + 3 \, \Omega^{-1} \, \partial_{\mu} \Omega \,. \end{split}$$

• Choosing $\phi = \Omega^2$, the action reads

$$S_{\rm eff} = \int d^4 x \, \widetilde{e} \, \left[\frac{1}{2 \, \kappa_4} \, \widetilde{T} + \frac{1}{2} \, \widetilde{g}^{\mu\nu} \partial_\mu \psi \, \partial_\nu \psi \right] \,,$$

(C.Q. Geng, Chang Lai, LWL, H.H. Tseng, 2014)

where $\psi = (6/\kappa_4)^{1/2} \ln \Omega$ is the dilaton field.

■ There exists an Einstein-frame for such a non-minimal coupled effective Lagrangian in teleparallel gravity.

Equation of Motion of the TEGR Effective Action

The gravitational equation of motion

$$\frac{1}{2} e_i{}^{\mu} \left(\phi T + 2 T^{\sigma} \partial_{\sigma} \phi \right) - e_i{}^{\rho} \left(\phi T^j{}_{\rho\nu} S_j{}^{\mu\nu} \right) - e_i{}^{\nu} \left(\partial_{\sigma} \phi T^{\mu}{}_{\nu}{}^{\sigma} + \partial_{\nu} \phi T^{\mu} + \partial^{\mu} \phi T_{\nu} \right) + \frac{1}{e} \partial_{\nu} \left(e \left(\phi S_i{}^{\mu\nu} + e_i{}^{\mu} \partial^{\nu} \phi - e_i{}^{\nu} \partial^{\mu} \phi \right) \right) = \kappa_4 \Theta_i^{\mu}$$

with $\Theta^{\mu}_{\nu} = {\rm diag}(\rho,-P,-P,-P)$

The modified Friedmann equations in flat FLRW universe are

$$3\phi H^2 + 3H\dot{\phi} = \kappa_4 \rho,$$

$$3\phi H^2 + 2\dot{\phi} H + 2\phi \dot{H} + \ddot{\phi} = -\kappa_4 P,$$

where $H=\dot{a}/a$ is the Hubble parameter (here $\rho=P=0$ is assumed.)

• The equations of motion of scalar field ϕ in the

$$T - 2 \,\partial_{\mu}T^{\mu} - 2T^{\mu}\Gamma^{\nu}_{\nu\mu} + e \,L_m = 0 \xrightarrow[\Gamma^{\nu}_{\nu\mu} = \Gamma^{\alpha}_{\alpha 0} = 3\frac{\dot{a}}{a} a \ddot{a} + \dot{a}^2 = 0 \,.$$

• Suppose the solution of a(t) is proportional to t^m , the solution is

$$a(t) = a_s + b\sqrt{t} \,.$$

- The constraint of the coefficient: $a_s b = 0$
- b = 0 case:
 - $a(t) = a_s \Rightarrow$ the static universe.
- $a_s = 0$ case:
 - The Hubble parameter H = 1/(2t) > 0
 - The the acceleration of scale factor $\ddot{a} = -b/(4t^{2/3}) > 0$ for b < 0⇒ accelerated expanding universe.
- In general relativity, the equation of motion of ϕ is $\tilde{R}(e) = 0$ \implies the same solution for the scale factor in vacuum.

Assuming that ω = ω(φ) under the conformal transformation
 Due to dφ̃/dφ = ω' exp(λ_φω)(1 + λ_φω'φ) with ω' := dω/dφ

$$\widetilde{\phi} = \widetilde{\phi}(\phi) \implies \omega(\widetilde{\phi})$$

• By setting $\partial_{\mu}\omega = \partial_{\mu}\ln\widetilde{\phi}$.

• Lagrangian density (*) without A_{μ} becomes

$$\begin{split} \mathcal{L}_{g} &= \, \widetilde{e} \frac{1}{2\kappa_{4}} \left\{ \widetilde{\phi} \, \widetilde{T}_{\mathsf{NGR}} + \left(4 \, a + 2 \, b \right) \widetilde{g}^{\mu\nu} \, \widetilde{T}_{\mu} \, \partial_{\nu} \widetilde{\phi} \\ &+ \left(24a + 12b \right) \frac{1}{\widetilde{\phi}} \, \widetilde{g}^{\mu\nu} \partial_{\mu} \widetilde{\phi} \, \partial_{\nu} \widetilde{\phi} \right\}. \end{split}$$

• Comparing with the effective Lagrangian

$$2a+b=-c,$$

$$24a+12b=2a+b+c,$$

$$\implies \begin{cases} 2a+b=0,$$

$$c=0.$$

Other Conformal Invariant Model

The conformal invariant model investigated by Maluf and Faria is

$$\mathcal{L} = ke \bigg[-\phi^2 \bigg(\frac{1}{4} T^{abc} T_{abc} + \frac{1}{2} T^{abc} T_{cba} - \frac{1}{3} T^a T_a \bigg) + k' g^{\mu\nu} D_{\mu} \phi D_{\nu} \phi \bigg],$$

(Maluf and Faria, 2012)

where $k = 1/(16\pi G)$, k' = 6 and $\eta_{ab} = (-1, +1, +1, +1)$ as well as $D_{\mu} := \partial_{\mu} - \frac{1}{3}T_{\mu}$, which gives the conformal invariant condition

2a + b + 3c = 0.

In their discussion, a new arbitrary parameter k' for the scalar kinetic term is introduced, resulting in a four-parameters model.

Note:

In our model, the coefficient of the kinetic term of ϕ is 2a + b + c so that the conformal invariant theory is totally determined by three parameters only.

Comparison Table

• Minimal coupled models with $\frac{1}{2\kappa}T_{\rm NGR}$

Class	Additional condition	Reference
2a + b + c = 0,	-	Einstein, 1929
$(a, b, c) = (\frac{1}{4}, \frac{1}{2}, -1)$	-	Cho, 1976
2a + b + c = 0	-	Hehl et al., 1978
c = -1	-	Nitsch and Hehl, 1980
C – 1	-	Hayashi and Shirafuji, 1979
2a + b + c = 0,	Static isotropic	Hehl et al., 1978
$(a, b, c) = (\frac{1}{2}, 0, -1)$	metric by Scherrer	Nitsch and Hehl, 1980
$\begin{aligned} 2a+b+c &= 0,\\ (a,b,c) &= \left(\frac{1}{4},\frac{1}{2},-1\right) \end{aligned}$	Einstein-frame	Geng et al., 2014
$2a + b + c = 0,$ $c \le 0$		\checkmark

Non-minimal coupled models with $\frac{1}{2\kappa}\phi T_{\rm NGR}$ (conformal invarience)

Class	Additional condition	Reference
2a+b+3c=0	$k'g^{\mu u}D_{\mu}\phi D_{\nu}\phi,$ where $D_{\mu}:=\partial_{\mu}-rac{1}{3}T_{\mu}$ with arbitrary k'	Maluf and Faria, 2012
2a + b + c = 0, $c = 0$	-	\checkmark
2a + b + 4c = 0 $2a + b = 0$	$\begin{array}{l} 2a+b+c>0,\\ c\neq 0 \end{array}$	\checkmark

Equations of Motion of the NGR Effective Action • Varying the full action $S = S_q + S_m$ with respect to e^i_{μ} , A_{μ} and ϕ $\frac{1}{2} \, e_i{}^{\mu} \Big(\phi \, T_{\mathsf{NGR}} - \frac{a \kappa^2}{4} \phi^3 \, g^{\lambda \rho} g^{\nu \sigma} F_{\lambda \nu} \, F_{\rho \sigma} + \frac{2a + b + c}{\phi} \, g^{\lambda \nu} \partial_{\lambda} \phi \, \partial_{\nu} \phi$ $-2cg^{\lambda\nu}T_{\lambda}\partial_{\nu}\phi\bigg) - e_{i}^{\rho}\bigg\{\phi T^{j}{}_{\rho\nu}\Sigma_{j}{}^{\mu\nu} - \frac{a\kappa^{2}}{2}\phi^{3}g^{\mu\lambda}g^{\nu\sigma}F_{\lambda\nu}F_{\rho\sigma}$ $+\frac{2a+b+c}{\phi}g^{\mu\lambda}\partial_{\lambda}\phi\,\partial_{\rho}\phi-c\left(\partial_{\sigma}\phi\,T^{\mu}{}_{\rho}{}^{\sigma}+\partial_{\rho}\phi\,T^{\mu}+\partial^{\mu}\phi\,T_{\rho}\right)\right\}$ $+\frac{1}{e}\partial_{\nu}\left\{e\left(\phi\Sigma_{i}^{\mu\nu}-c\,e_{i}^{\mu}\,\partial^{\nu}\phi+c\,e_{i}^{\nu}\,\partial^{\mu}\phi\right)\right\}=\kappa_{4}\lambda\phi\Theta_{i}^{\mu}\,,$ $\frac{\lambda}{e} \frac{\delta \mathcal{L}_{\mathsf{m}}}{\delta A_{...}} + \frac{1}{e} \frac{a\kappa^2}{2\kappa_4} \phi^2 \,\partial_\mu \bigg(eF^{\mu\nu} \bigg) + \frac{3a\kappa^2}{2\kappa_4} \,\phi \,\partial_\mu \phi \,F^{\mu\nu} = 0 \,,$ $T_{\rm NGR} - \frac{3a\kappa^2}{4} \phi^2 g^{\mu\rho} g^{\nu\sigma} F_{\mu\nu} F_{\rho\sigma} + 2\kappa_4 \lambda \left(L_{\rm m} + \frac{\phi}{e} \frac{\delta \mathcal{L}_{\rm m}}{\delta \phi} \right)$ $+\frac{2a+b+c}{\phi^2}\left(g^{\mu\nu}\partial_{\mu}\phi\partial_{\nu}\phi-2\phi\,\Box\phi\right)+\frac{2c}{c}\,\partial_{\nu}\left(eg^{\mu\nu}T_{\mu}\right)=0\,,$ where $\hat{\Box} := \frac{1}{e} \partial_{\nu} (eg^{\mu\nu} \partial_{\mu})$ and $\Theta_i{}^{\mu} := -\frac{1}{e} \frac{\delta \mathcal{L}_m}{\delta e^i}$ is the dynamical

energy-momentum tensor with $\mathcal{L}_{m} := eL_{m}$.

Weyl Geometry

■ Parallel transport of the vector V from point $P(=x^{\mu})$ to point $P'(=x^{\mu} + dx^{\mu})$

$$\nabla V = dx^{\nu} (\nabla_{\nu} V^{\mu}) \partial_{\mu} = (dV^{\mu} - \delta V^{\mu}) \partial_{\mu} = 0.$$

• Weyl Geometry (*Weyl*, 1918): Define the measure $l := g_{\mu\nu}V^{\mu}V^{\nu}$ of V^{μ} , the variation of the measure l is proportional to l with the 1-form factor $\varphi = \varphi_{\mu}dx^{\mu}$

$$dl = -\varphi l = -(\varphi_{\rho} dx^{\rho})g_{\mu\nu}V^{\mu}V^{\nu} \,.$$

However the variation of l can be written as

$$\begin{split} dl &= d \left(g_{\mu\nu} V^{\mu} V^{\nu} \right) \\ &= \partial_{\rho} g_{\mu\nu} dx^{\rho} V^{\mu} V^{\nu} dx^{\rho} - g_{\mu\nu} \Gamma^{\mu}{}_{\sigma\rho} dx^{\rho} V^{\sigma} V^{\nu} - g_{\mu\nu} \Gamma^{\nu}{}_{\sigma\rho} dx^{\rho} V^{\mu} V^{\sigma} \\ &= (\partial_{\rho} g_{\mu\nu} - g_{\sigma\nu} \Gamma^{\sigma}{}_{\mu\rho} - g_{\mu\sigma} \Gamma^{\sigma}{}_{\nu\rho}) dx^{\rho} V^{\mu} V^{\mu} \,, \end{split}$$

where $dV^{\mu} = \delta V^{\mu} = -\Gamma^{\mu}{}_{\sigma\rho}V^{\sigma}dx^{\rho}.$

Identity:
$$\partial_{\rho}g_{\mu\nu} - g_{\sigma\nu}\Gamma^{\sigma}{}_{\mu\rho} - g_{\mu\sigma}\Gamma^{\sigma}{}_{\nu\rho} = -\varphi_{\rho}g_{\mu\nu} \neq 0.$$

◀ WGT