

Field-theoretic simulations of colliding superconducting strings

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with Daisuke Yamauchi (Kanagawa), Daniele Steer (APC), Marc Lilley (IAP)

Field-theoretic simulations of colliding superconducting strings

(and gravitational waves from their network)

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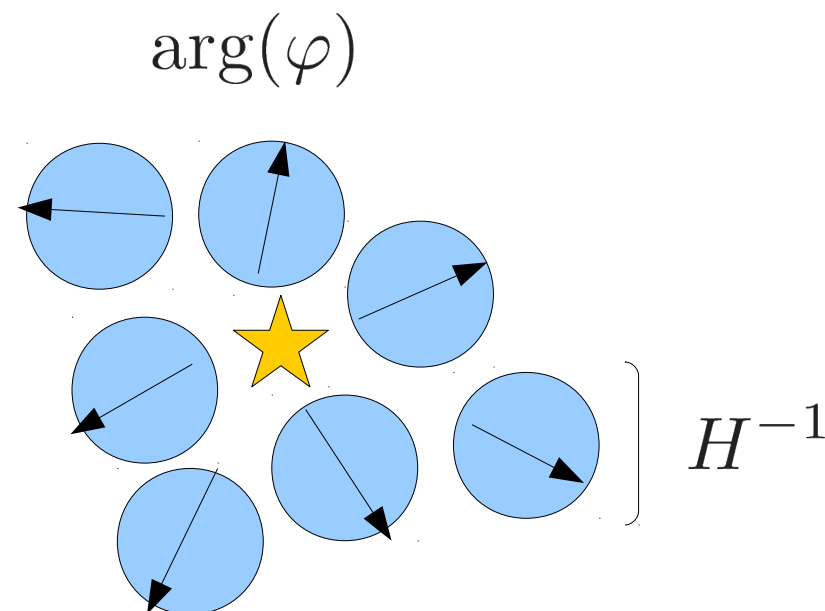
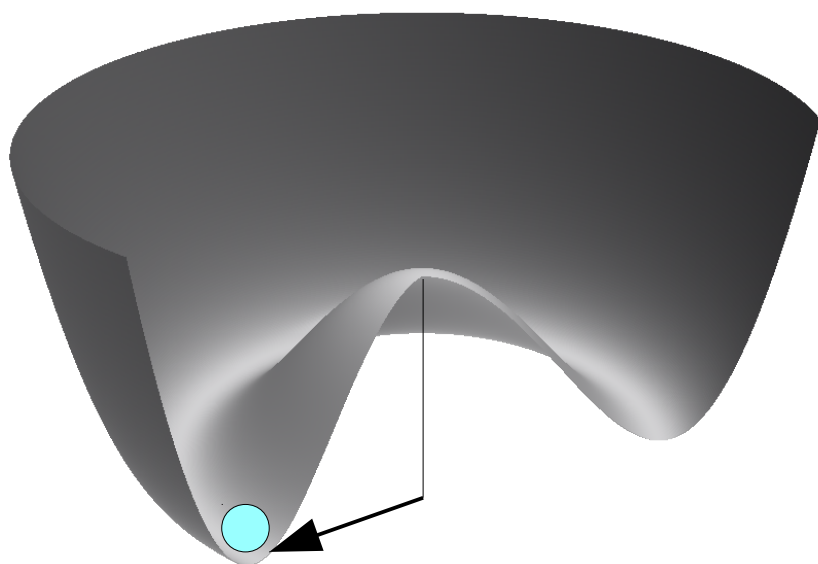
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[colliding strings]

with Daisuke Yamauchi (Kanagawa), Daniele Steer (APC), Marc Lilley (IAP)

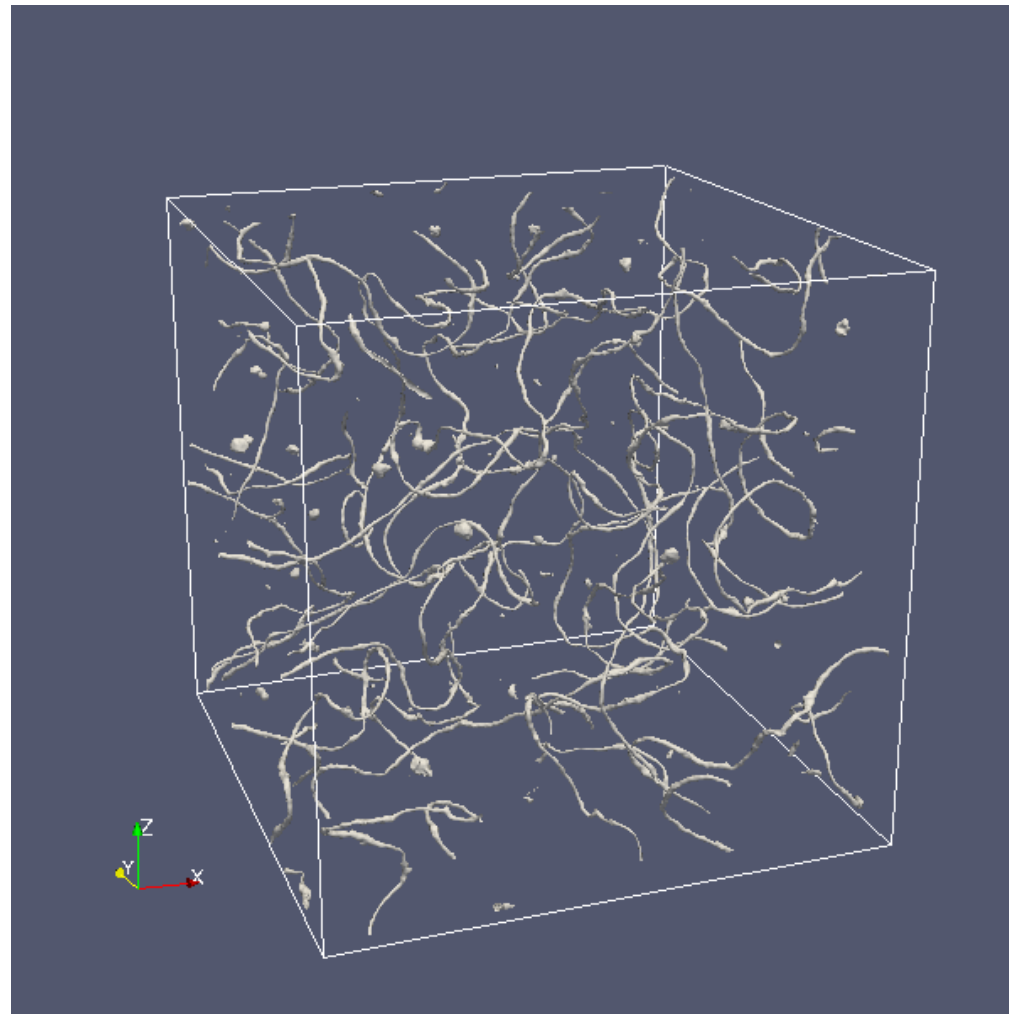
[GWs from strings]

with Daisuke Yamauchi (Kanagawa)



φ is aligned in a causal volume (=Hubble),
but is different in each volume in general.

If the above situation is realised,
a cosmic string appears at  .



$$\sim 10H^{-1}$$

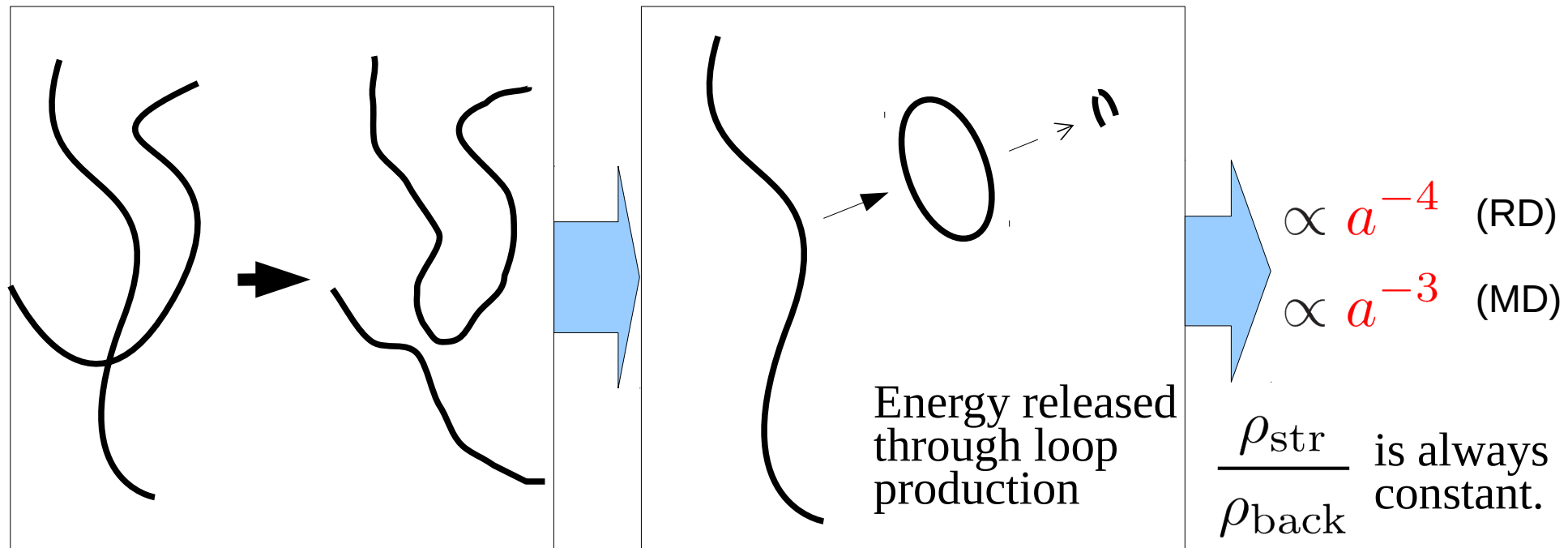
Interests : GUTs ? Superstrings ? Source of GWs ?

Non-trivial signals on CMB ? etc...

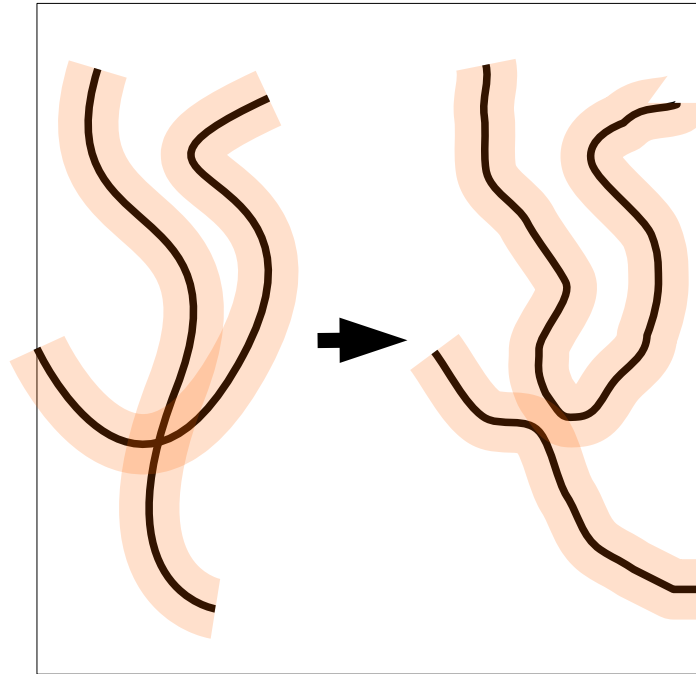
Decay rate of energy density

Naively, $\rho_{\text{str}} \propto a^{-2}$ \longrightarrow Eventually dominated, so highly suppressed.

But,



The fate of strings highly depends on the efficiency of reconnection process.



Reconnection process works even if strings couple with matter ?

Extend a past numerical study by Laguna and Matzner. [Laguna and Matzner, PRD 41 \(1990\) 1751](#)

Model Lagrangian

Abelian-Higgs model ($U(1)$ gauge theory) + **additional scalar field**

$$S = - \int dx^4 \sqrt{-g} \left(\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + (D_\mu \phi)^* (D^\mu \phi) + (\partial_\mu \sigma)^* (\partial^\mu \sigma) + V(\phi, \sigma) \right)$$

$$V(\phi, \sigma) = \frac{\lambda_\phi}{4} (|\phi|^2 - \eta^2)^2 + \boxed{\lambda_{\phi\sigma} (|\phi|^2 - \eta^2) |\sigma|^2} + \frac{\lambda_\sigma}{4} |\sigma|^4 + \frac{m_\sigma^2}{2} |\sigma|^2$$

$$D_\mu \equiv \partial_\mu - ieA_\mu \quad F_{\mu\nu} = \partial_\mu A_\nu - \partial_\nu A_\mu$$



Conserved current : $j_\mu = 2\text{Im}(\sigma^* \partial_\mu \sigma)$

* Its realisability and observability in cosmological context is discussed by Witten.

Ansatz and potential

$$V(\phi, \sigma) = \frac{\lambda_\phi}{4} (|\phi|^2 - \eta^2)^2 + \lambda_{\phi\sigma} (|\phi|^2 - \eta^2) |\sigma|^2 + \frac{\lambda_\sigma}{4} |\sigma|^4 + \frac{m_\sigma^2}{2} |\sigma|^2$$

$$D_\mu = \partial_\mu - ieA_\mu$$

$$\phi(\mathbf{r}) = \eta f(r) e^{in\theta} \quad A_\theta(\mathbf{r}) = \frac{n}{e} \alpha(r) \quad \sigma(\mathbf{r}) = \eta g(r) e^{i(kz - \omega t)}$$

Model parameters

winding number : $n = 1$

gauge coupling : e

self-coupling of ϕ : $\beta_\phi \equiv \frac{\lambda_\phi}{2e^2} = 1$

bear mass of σ : $\mu^2 \equiv \frac{m_\sigma^2}{2e^2\eta^2} = 0.01$

self-coupling of σ : $\beta_\sigma \equiv \frac{\lambda_\sigma}{2e^2} = 1$

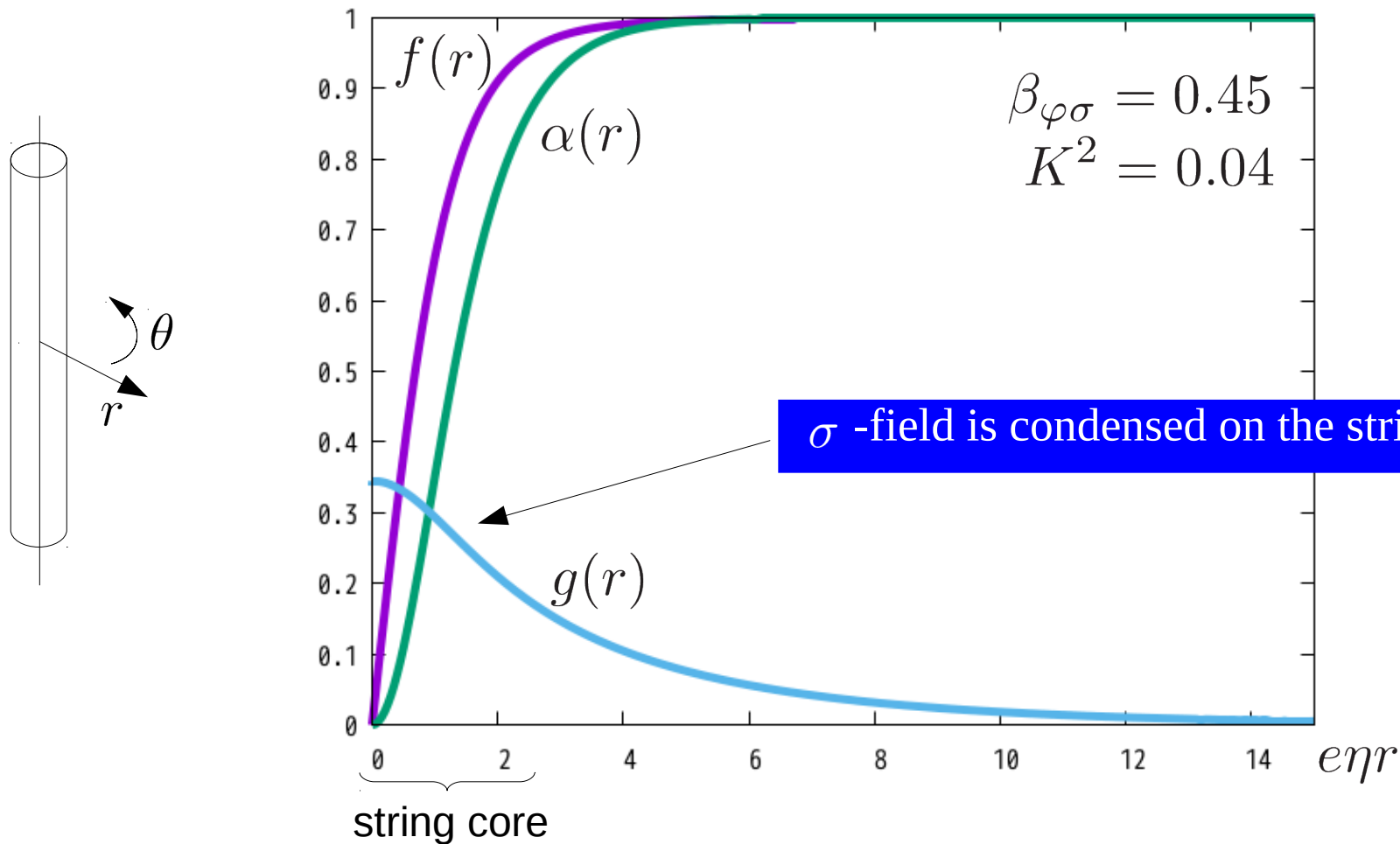
charge of σ : $\Omega^2 \equiv \frac{\omega^2}{e^2\eta^2} = 0$

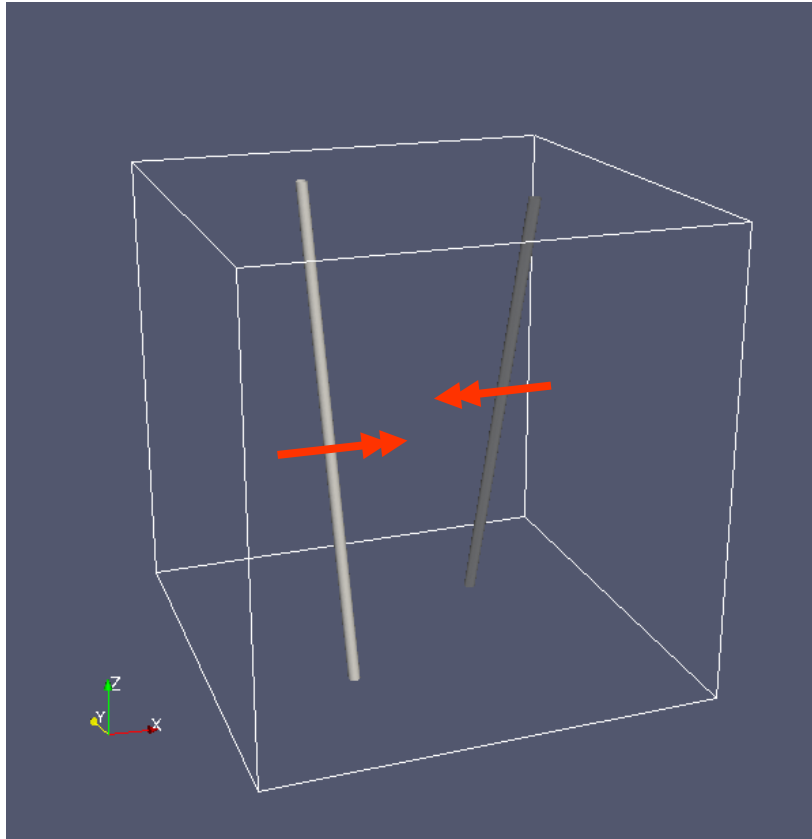
coupling between ϕ & σ : $\beta_{\phi\sigma} \equiv \frac{\lambda_{\phi\sigma}}{2e^2}$

current of σ : $K^2 \equiv \frac{k^2}{e^2\eta^2}$

Find straight vortex solutions

$$\phi(\mathbf{r}) = \eta f(r) e^{i\theta} \quad A_\theta(\mathbf{r}) = \frac{1}{e} \alpha(r) \quad \sigma(\mathbf{r}) = \eta g(r) e^{ikz}$$





Strategy

- Prepare 2 stable straight strings.
- Lorentz boost (velocity+rotation)

$$x^{\mu'} = \Lambda^{\mu}_{\nu} x^{\nu}$$

- Superposition

$$\phi = \frac{1}{\eta} \phi^{(1)} \phi^{(2)}$$

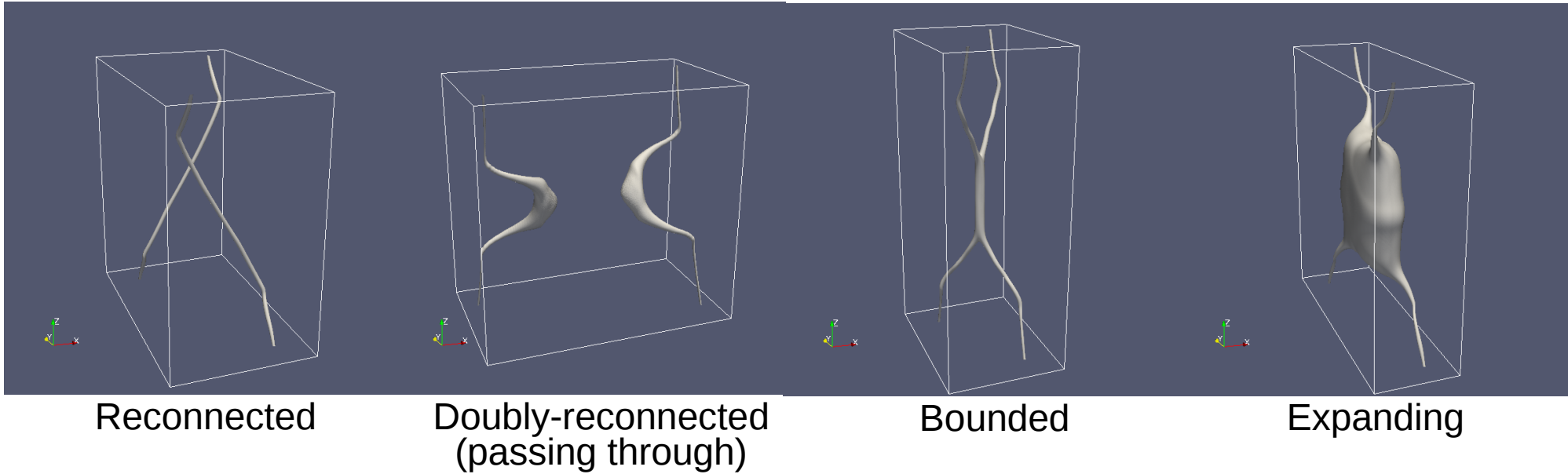
$$A_{\mu} = A_{\mu}^{(1)} + A_{\mu}^{(2)}$$

$$\sigma = \sigma^{(1)} + \sigma^{(2)}$$

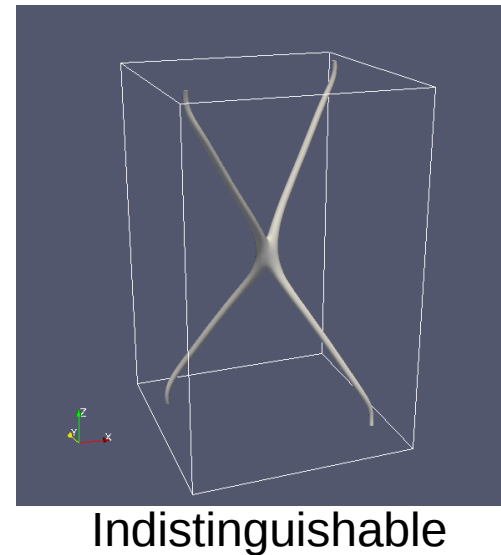
Numerical methods

- Leap-Frog scheme
- 2nd-order finite difference
- Adaptive box size depending on velocity and angle, roughly $200^3 \sim 800^3$

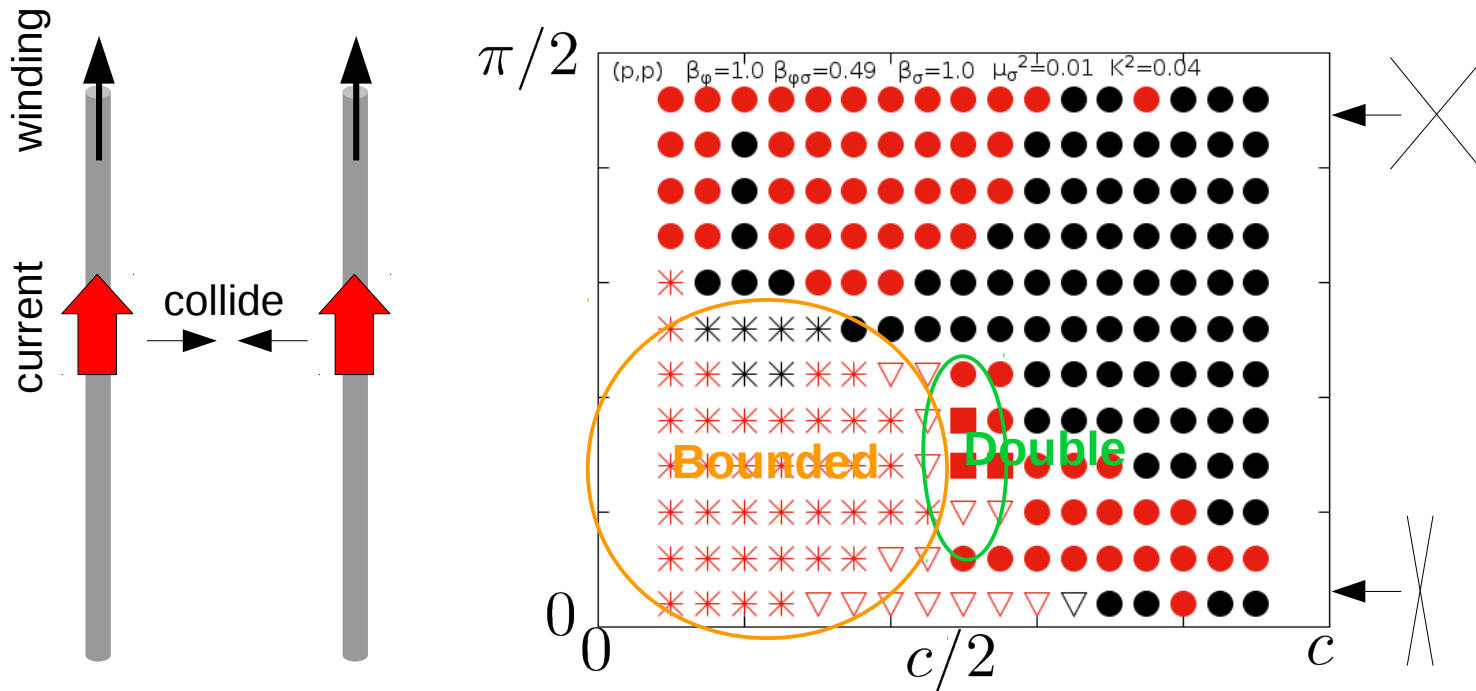
Define 4 kinds of final states



current flow		no current flow
●	reconnected	●
⊙	doubly-reconnected	⊙
+	bounded	+
×	expanding	×
?	indistinguishable	?



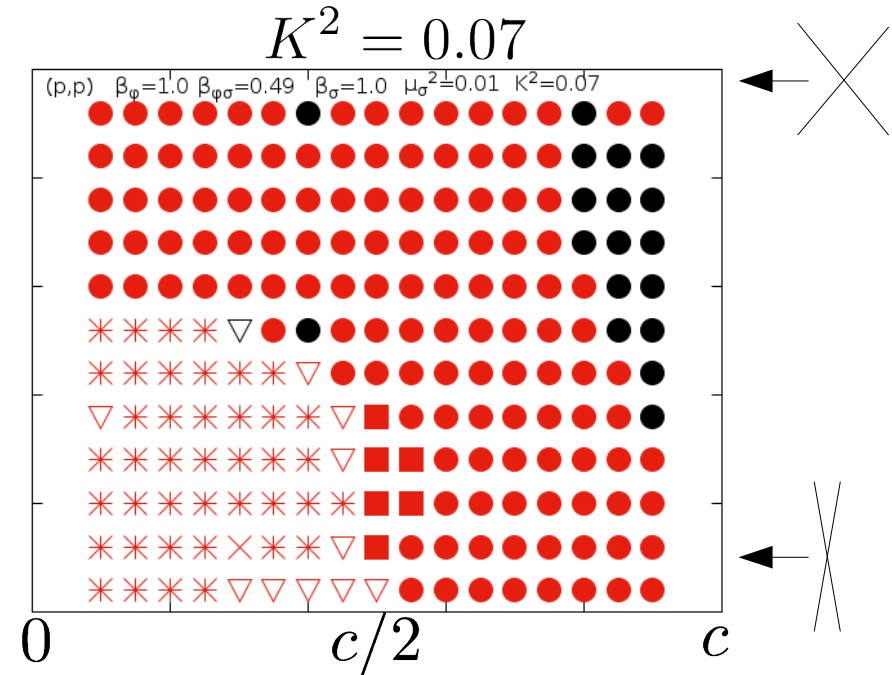
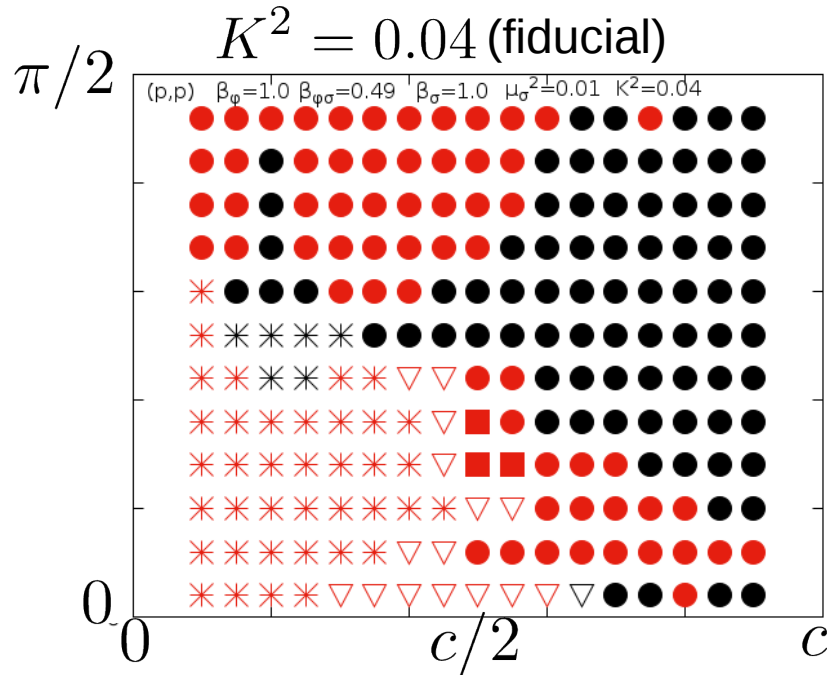
'Phase diagram' : fiducial setup



- Basically they reconnect with each other as usual
- Low-angle and low-velocity collisions leads to the bound state formation.
- The double reconnection takes place for intermediate-velocity collisions.
- The matter current tends to disappear for high-speed collisions.

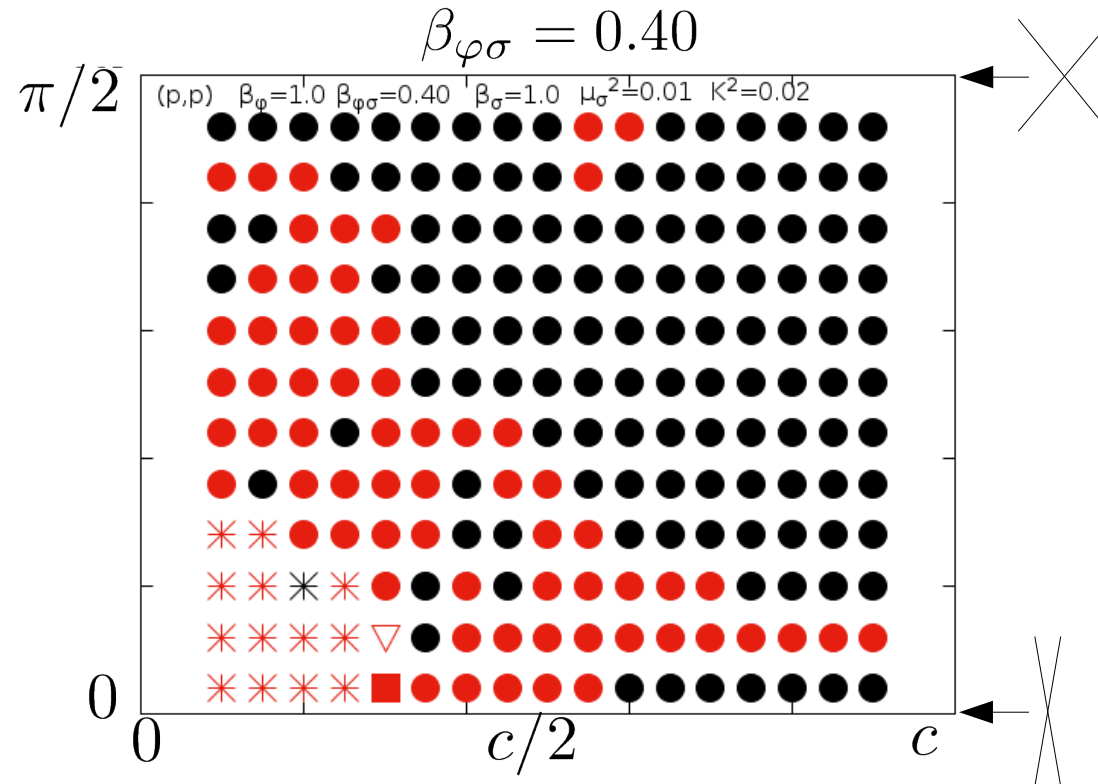
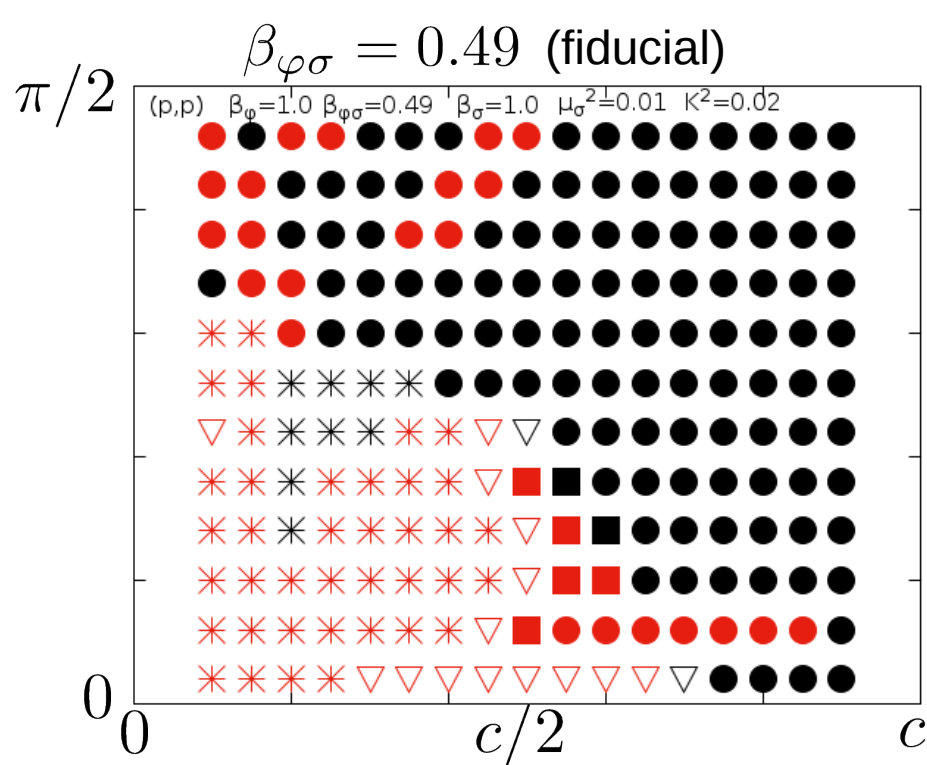
→ Discussion on the bound states : Steer, Lilley, Yamauchi, TH,
submitted to PRD, 1710.07475

'Phase diagram' : K^2 dependence



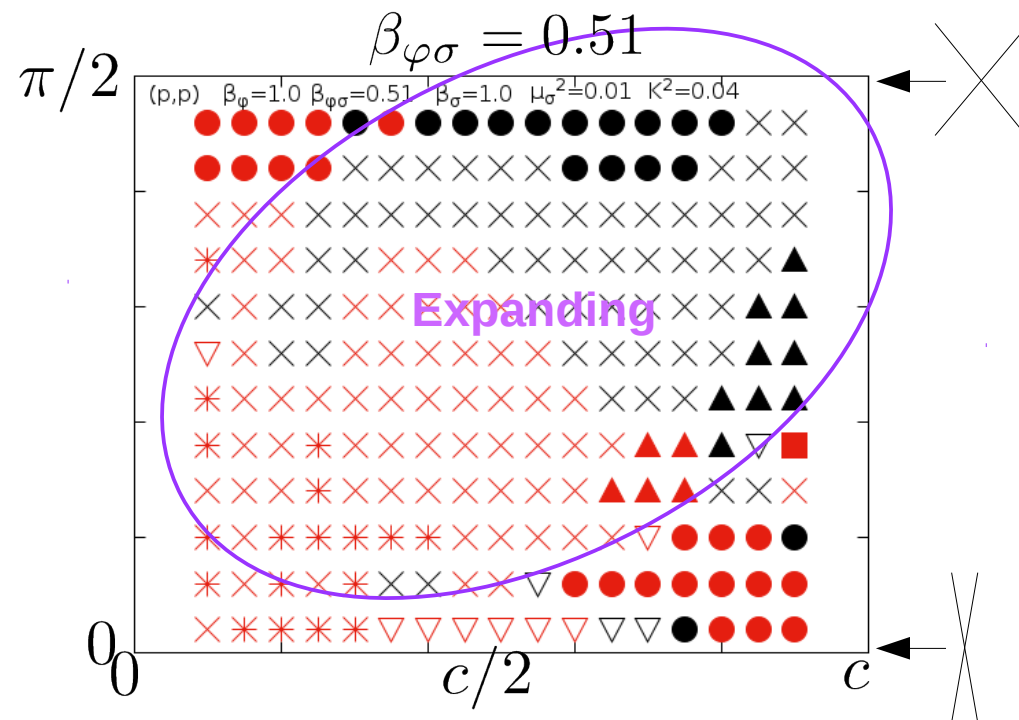
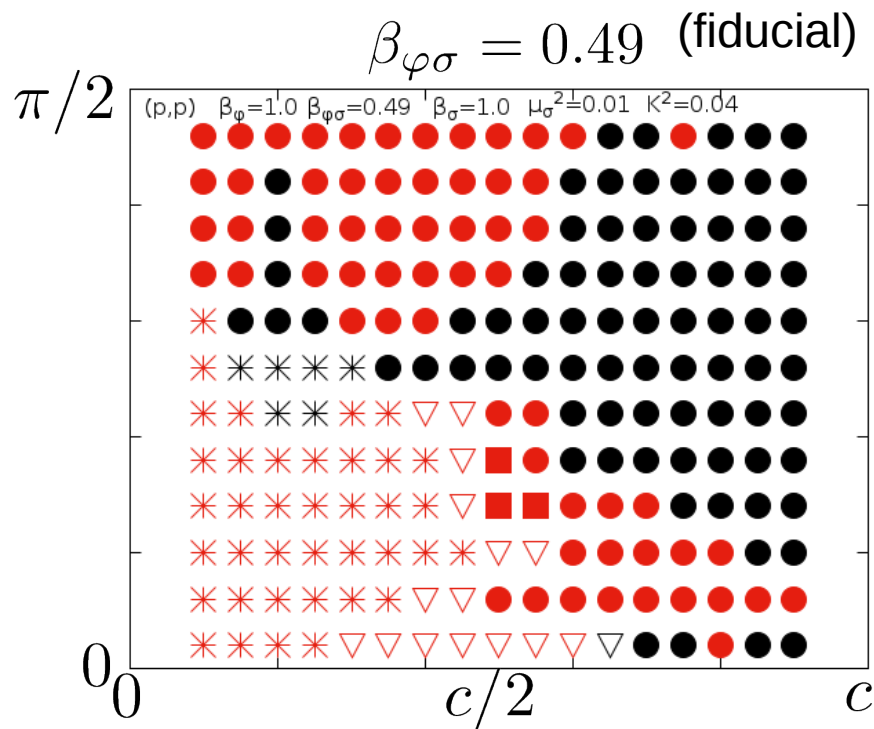
- The current after collisions is easy to survive for large K^2
- The initial current strength K^2 does not affect the final configuration.

'Phase diagram': $\beta_{\varphi\sigma}$ dependence



- The small $\beta_{\varphi\sigma}$ leads to less possibilities for forming bound states.
- $\beta_{\varphi\sigma}$ is not responsible for the final current strength.

'Phase diagram': $\beta_{\varphi\sigma}$ dependence



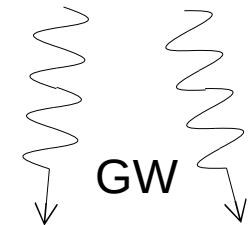
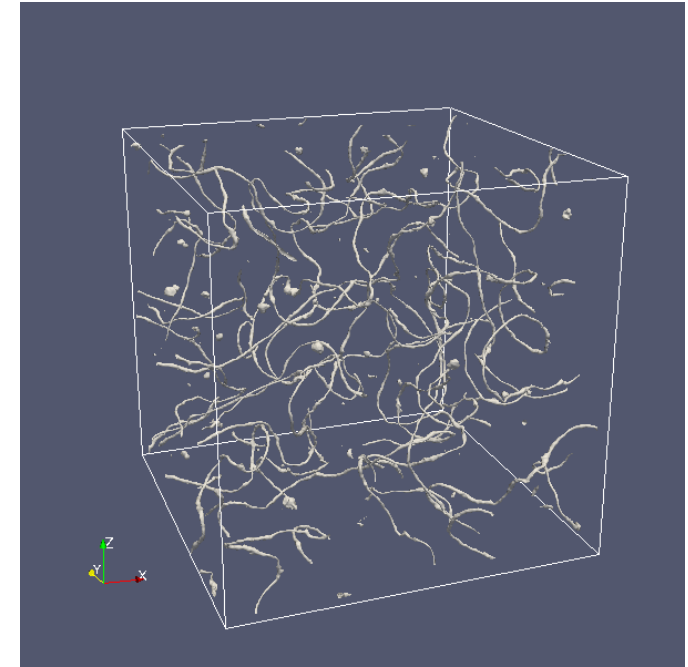
- The large $\beta_{\varphi\sigma}$ leads to the expanding bubble.
- Even so, high-speed collisions avoid the bubble nucleation.

$$\left\{ \begin{aligned} \frac{1}{\sqrt{-g}} D_\mu (\sqrt{-g} g^{\mu\nu} D_\nu \varphi) &= V_{\varphi^*} \\ \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} F^{\mu\nu}) &= -2e_\varphi g^{\mu\nu} \text{Im}(\varphi^* D_\mu \varphi) \\ \frac{1}{\sqrt{-g}} \partial_\mu (\sqrt{-g} \mathcal{F}^{\mu\nu}) &= -2e_\sigma g^{\mu\nu} \text{Im}(\sigma^* D_\mu \sigma) \end{aligned} \right.$$

$$\ddot{h}_{ij} + 2\mathcal{H}\dot{h}_{ij} - \Delta h_{ij} = \frac{2}{m_{\text{pl}}^2} (T_{ij}^{\text{scalar}} + T_{ij}^{\text{gauge}})$$

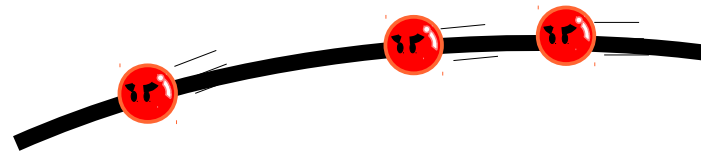
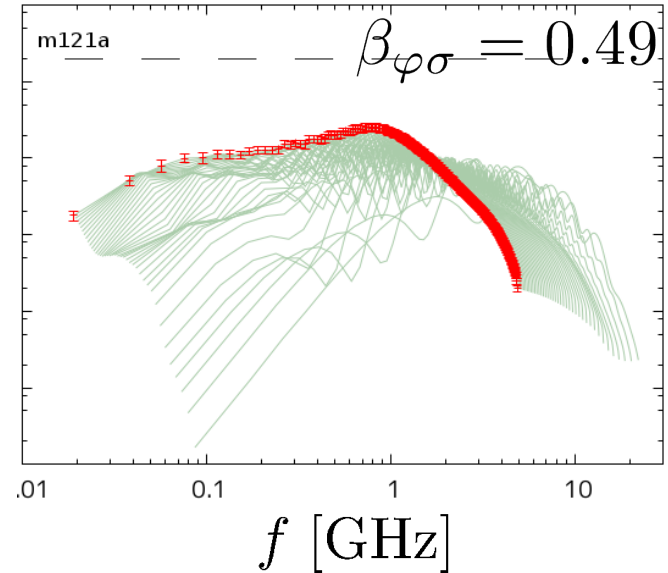
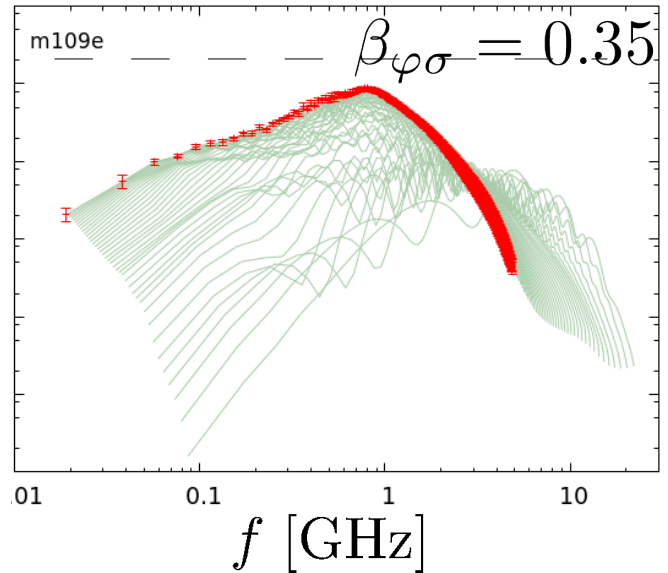
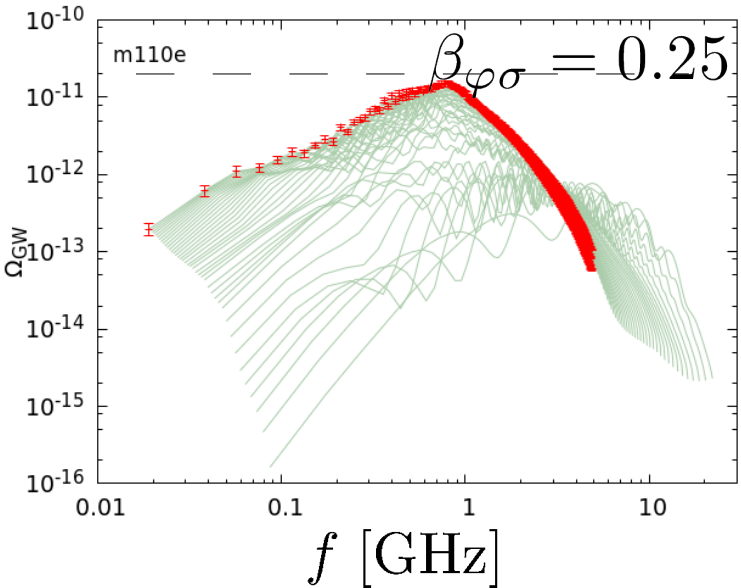
$$T_{ij}^{\text{scalar}} = 2\text{Re} [(D_\mu \varphi)^* (D_\nu \varphi)] + 2\text{Re} [(D_\mu \sigma)^* (D_\nu \sigma)]$$

$$T_{ij}^{\text{gauge}} = F_{\mu\alpha} F_\nu^\alpha + \mathcal{F}_{\mu\alpha} \mathcal{F}_\nu^\alpha$$

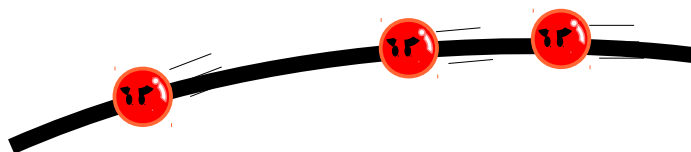
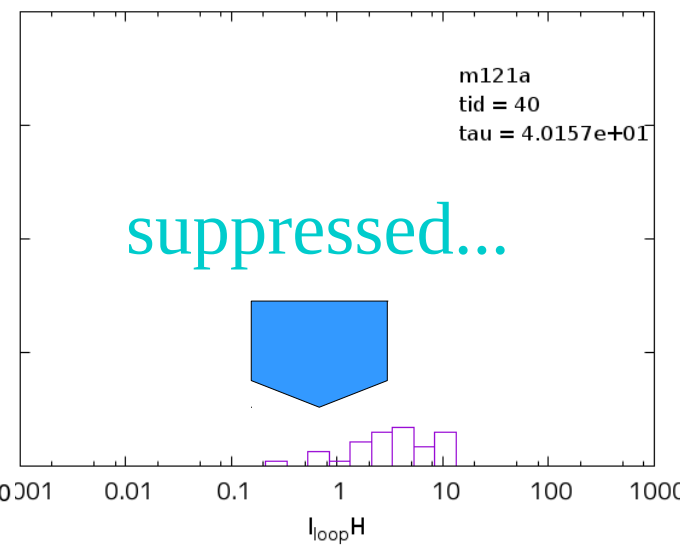
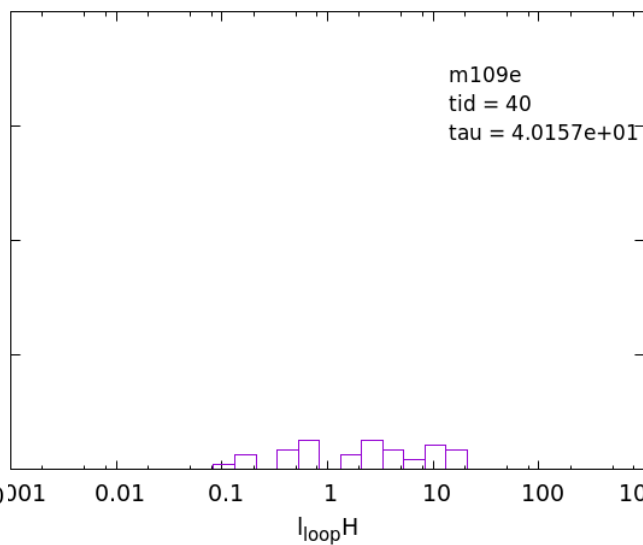
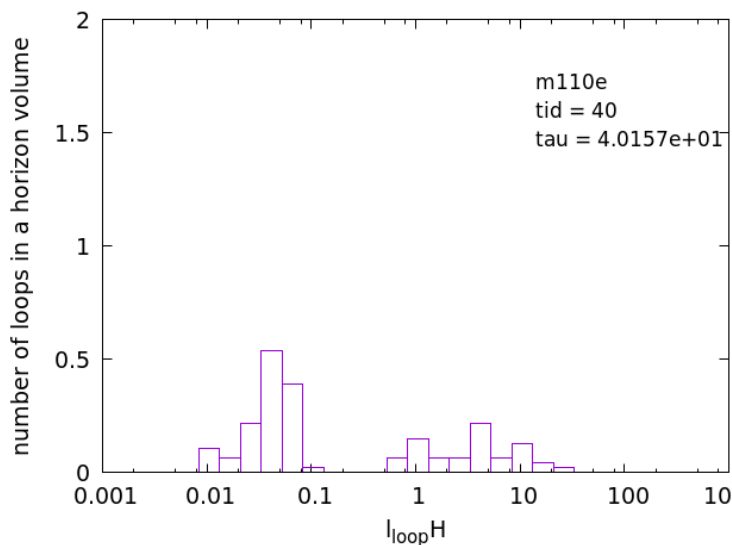


GW spectrum [preliminary]

$$\beta_{\varphi\sigma} \equiv \lambda_{\varphi\sigma}/2e^2$$



Loop distribution



Q. Can strings reconnect in a usual way even if they couple to matter ?

Yes, but they have rich diversity of their final states

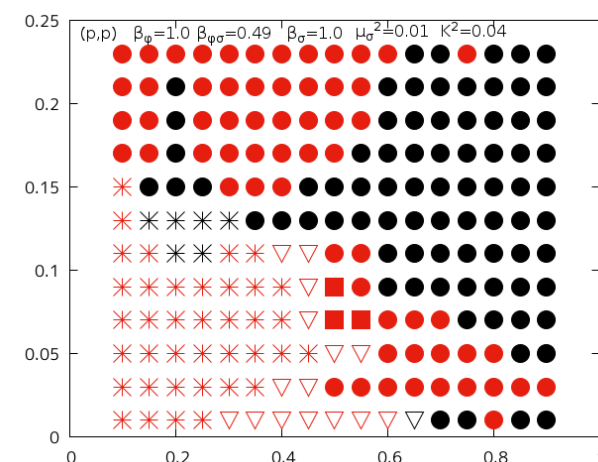
↔ Always successful
for critical AH strings

Main results:

- Stable pairs can **form a bound state** like Type-I AH strings.
- They can **pass through** each other by double-reconnection like Type-II AH strings.
- The final configuration depends on $\beta_{\varphi\sigma}$, not K^2 .
- K^2 is responsible only for the final current strength.

Preliminary results:

- GW signal is suppressed for superconducting cases
- strong coupling leads to the less population of loops ?

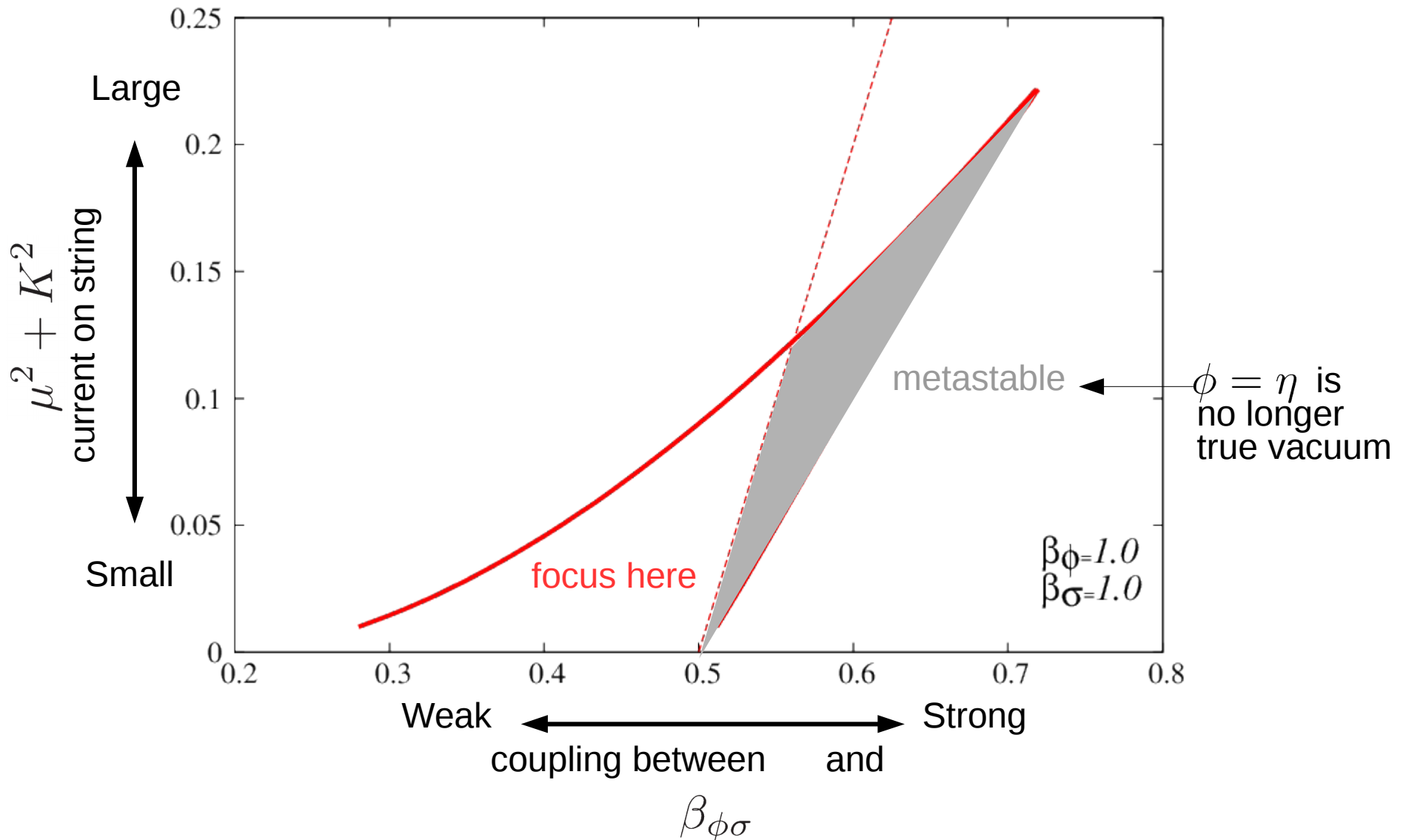


work in progress....

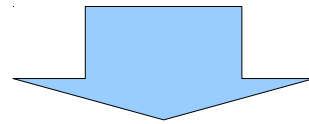


Result I : viable parameter region

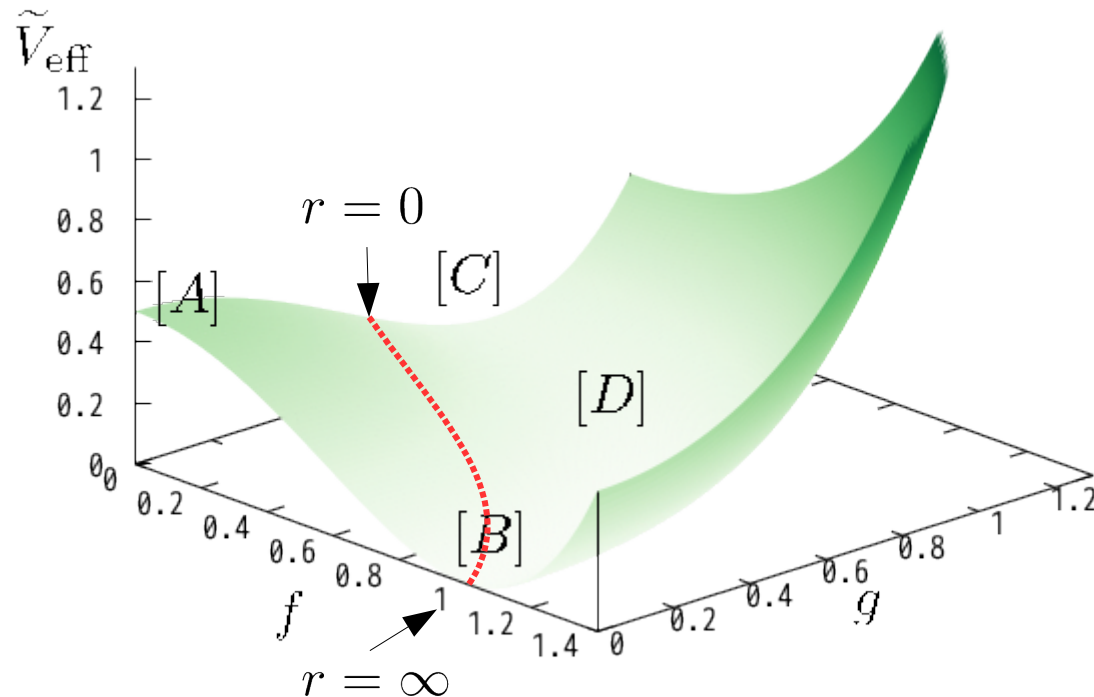
Superconducting string configuration is available only in the triangle region.



$$V(\phi, \sigma) = \frac{\lambda_\phi}{4} (|\phi|^2 - \eta^2)^2 + \lambda_{\phi\sigma} (|\phi|^2 - \eta^2) |\sigma|^2 + \frac{\lambda_\sigma}{4} |\sigma|^4 + \frac{m_\sigma^2}{2} |\sigma|^2$$



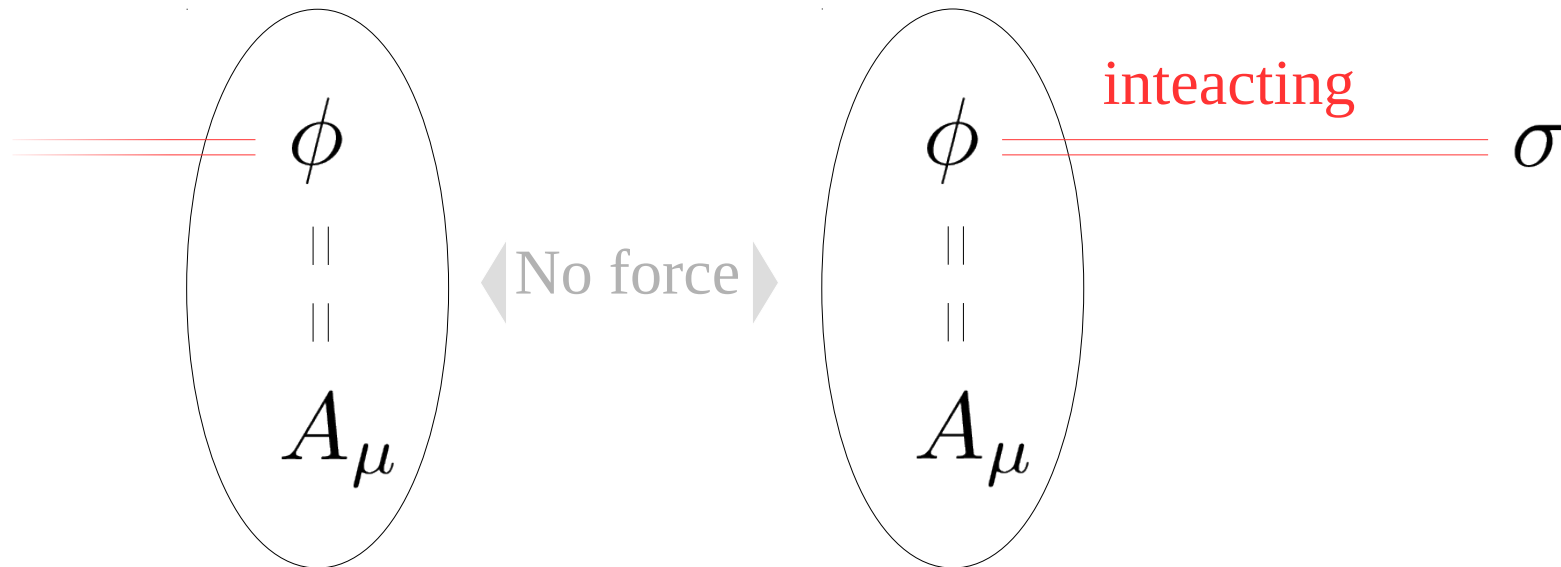
$$\tilde{V}_{\text{eff}}(\phi, \sigma) = \frac{\beta_\phi}{2} (f^2 - 1)^2 + 2\beta_{\phi\sigma} (f^2 - 1)g^2 + \frac{\beta_\sigma}{2} g^4 + \gamma g^2$$



Model : superconducting strings

string sector

“matter” sector



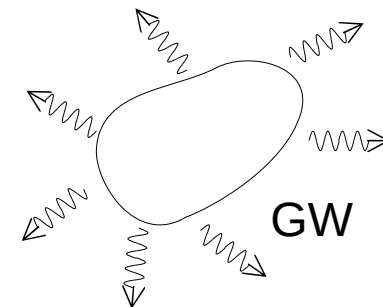
- ~~The seed of large-scale structure~~ → Inflation
- Massive strings can radiate GWs.

$$G\mu \approx \left(\frac{\eta}{M_{\text{pl}}} \right)^2 = 10^{-6} \Rightarrow \mu \sim 10^{22} \text{g/cm} \quad \mu H^{-1} \sim 10^{16} M_{\odot}$$

$$G\mu \leq 1.3 \times 10^{-7} \quad \text{Planck collaboration, A\&A 594 (2016) A13}$$

$$G\mu \leq 1.5 \times 10^{-8} \quad \text{Janet et al., ApJ 653 (2006) 1571}$$

- Probe for very early Universe
- A clue of (grand) unified theories ?
- Superstrings ?



Jones, Stoica, Tye, JHEP 07 (2002) 051
Sarangi, Tye, Phys.Lett. B536 (2002) 185

Scaling solution

String networks *without any interactions* between strings have the attractor solution of their characteristic length (= mean separation of strings), which is growing as

$$\xi(t) \propto t \propto H^{-1}$$



The number of strings in the cosmological horizon scale is always constant.

$$n_{\text{str}}(t) H^{-3} = \text{const}$$

Confirmed in numerical simulations.

Vincent, Antunes, Hindmarsh, PRL 80 (1998) 2277
Moore, Shellard, PRD 65 (2001) 023503

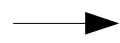
Energy density

Naively thinking ...

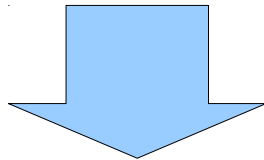
$$\rho_{\text{str}} = \frac{\mu a L}{(aL)^3} \propto a^{-2}$$



eventually dominated by strings !



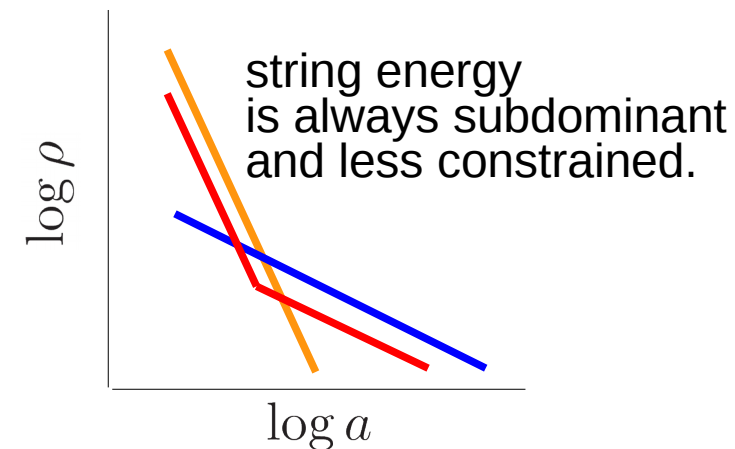
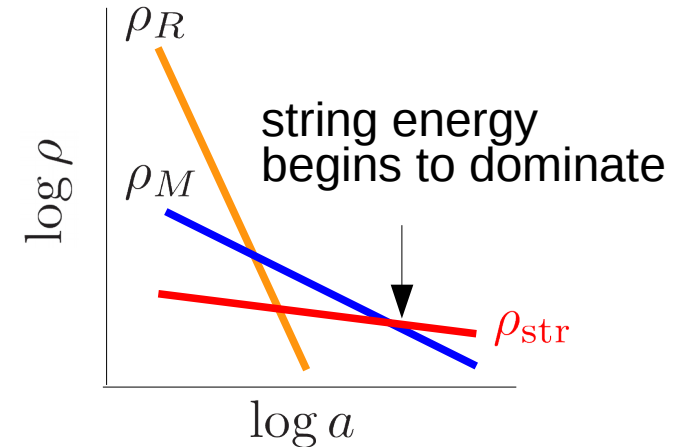
highly constrains amounts of strings



If strings follow scaling solutions,

$$\rho_{\text{str}} = \frac{\mu \xi}{\xi^3} \propto H^2 \propto a^{-4} \quad (\text{if radiation dominant})$$

$$\propto a^{-3} \quad (\text{if matter dominant})$$



String can survive at the present time !