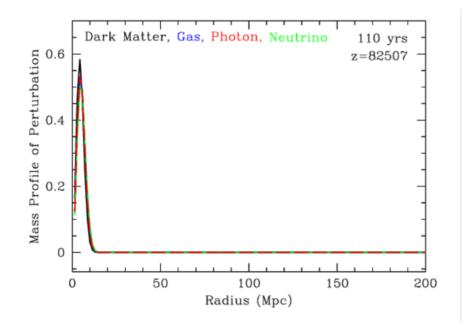


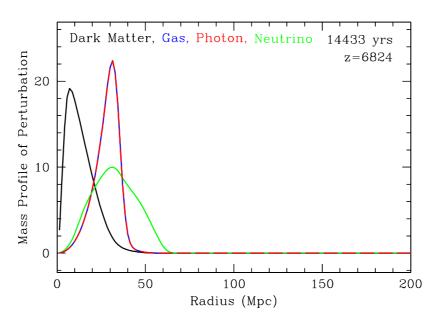
# Infrared Effects at the BAO Scale

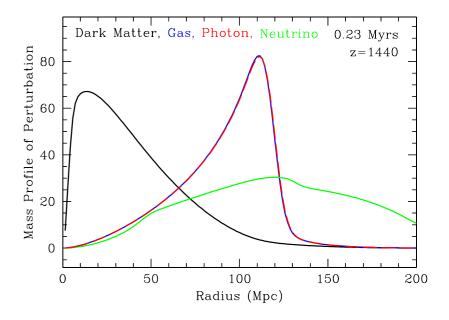
Gabriele Trevisan

CosPA 2017

## BAO in pictures

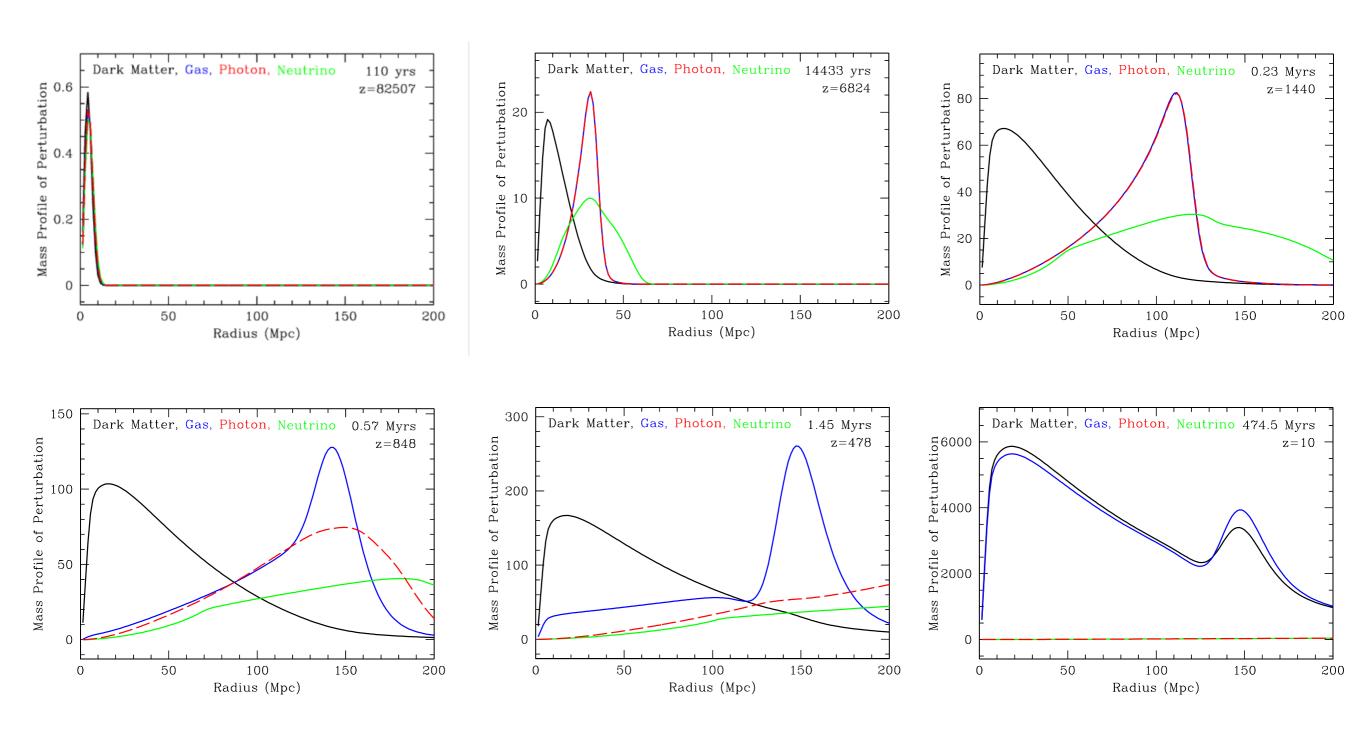






Eisenstein et al. 0604361

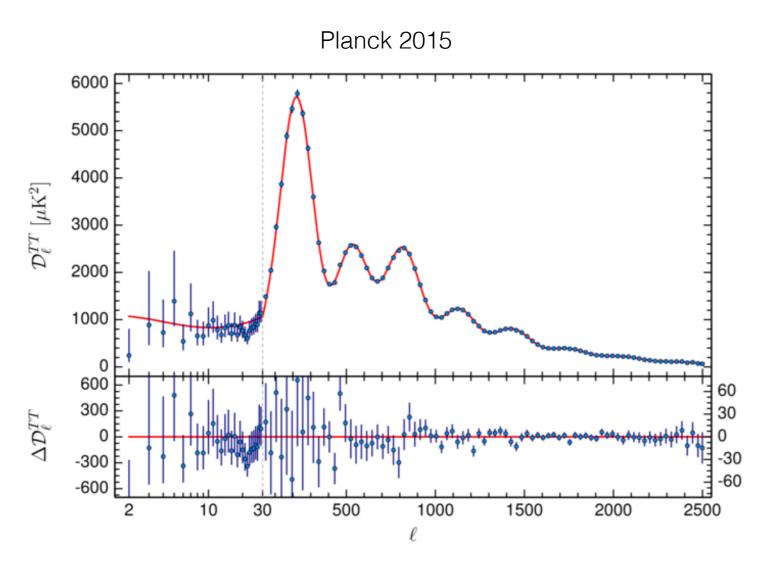
## BAO in pictures



Eisenstein et al. 0604361

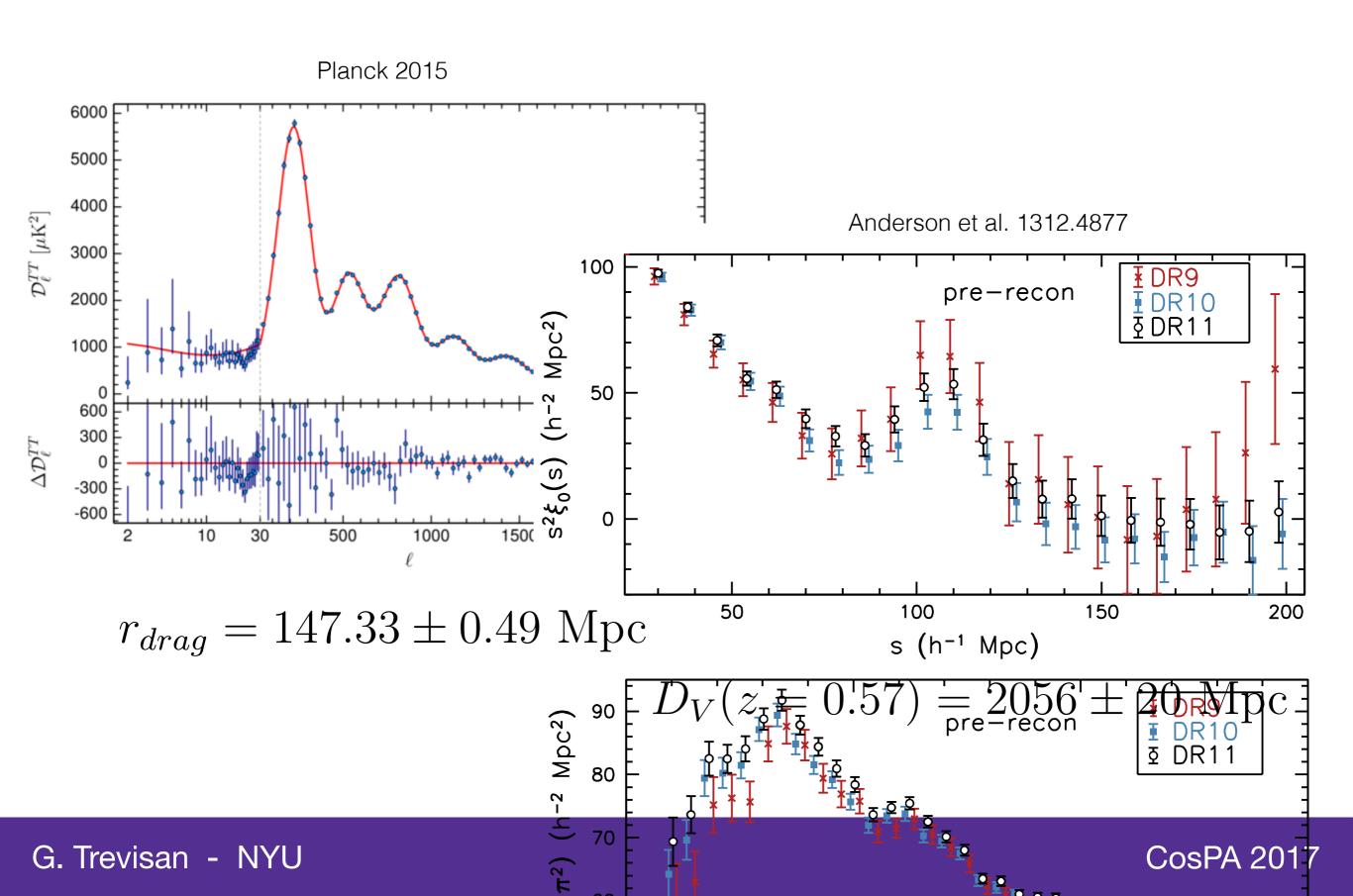


## BAO in the sky



$$r_{drag} = 147.33 \pm 0.49 \text{ Mpc}$$

## BAO in the sky



## Standard Perturbation Theory (SPT)

Fluid Equations:

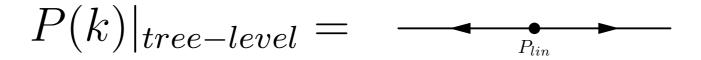
$$\partial_{\tau}\delta + \vec{\nabla} \cdot [(1+\delta)\vec{v}] = 0,$$

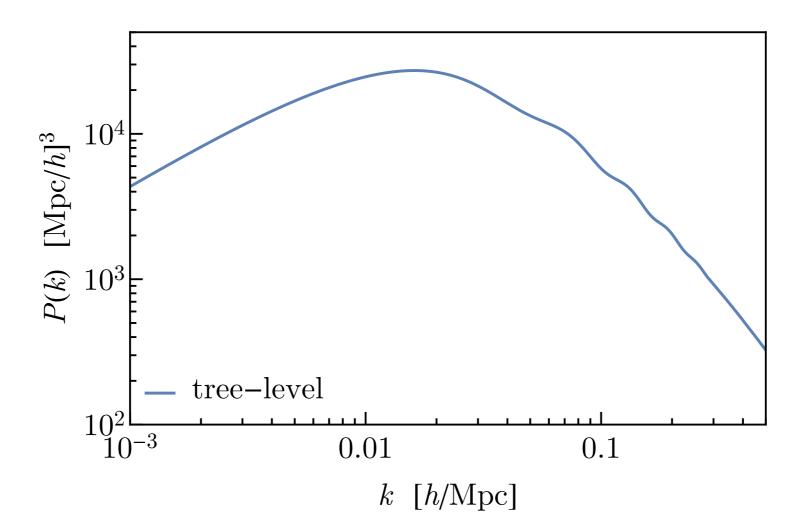
$$\partial_{\tau}\vec{v} + \mathcal{H}\vec{v} + \vec{v} \cdot \vec{\nabla}\vec{v} + \vec{\nabla}\phi = 0$$

$$\Delta\phi = \frac{3}{2}\mathcal{H}^{2}\Omega_{m}\delta$$

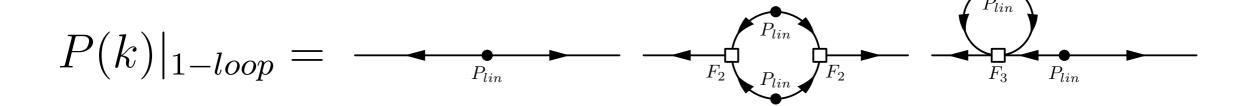
Perturbative solution:

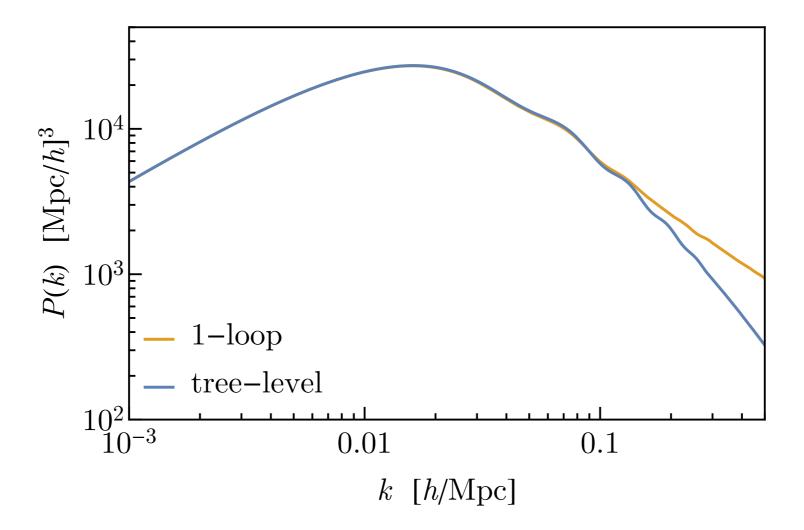
$$\delta_n(\vec{k})$$
  $F_n$   $\delta_{\mathrm{linear}}^n(\vec{q})$ 





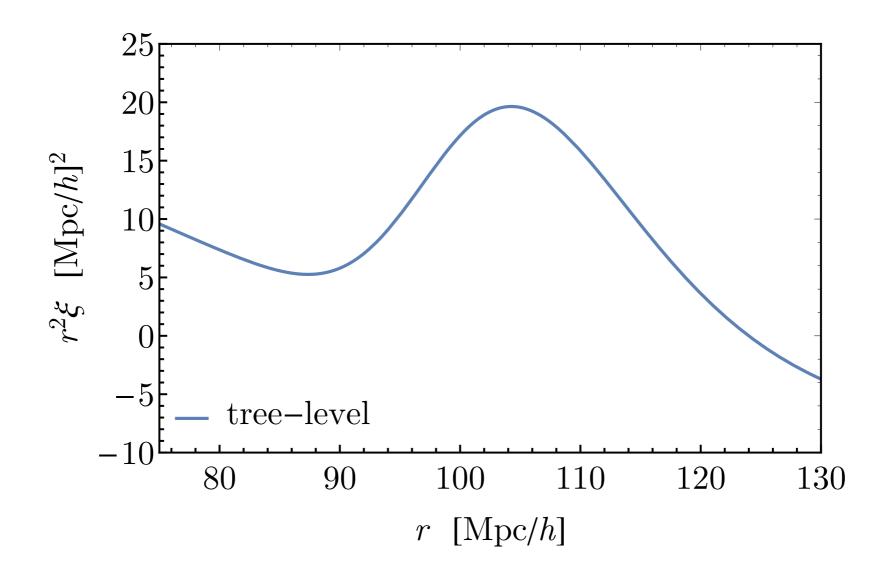
Where is the BAO in Fourier space?
The BAO signal is ~5% oscillation with freq. 1/150 Mpc



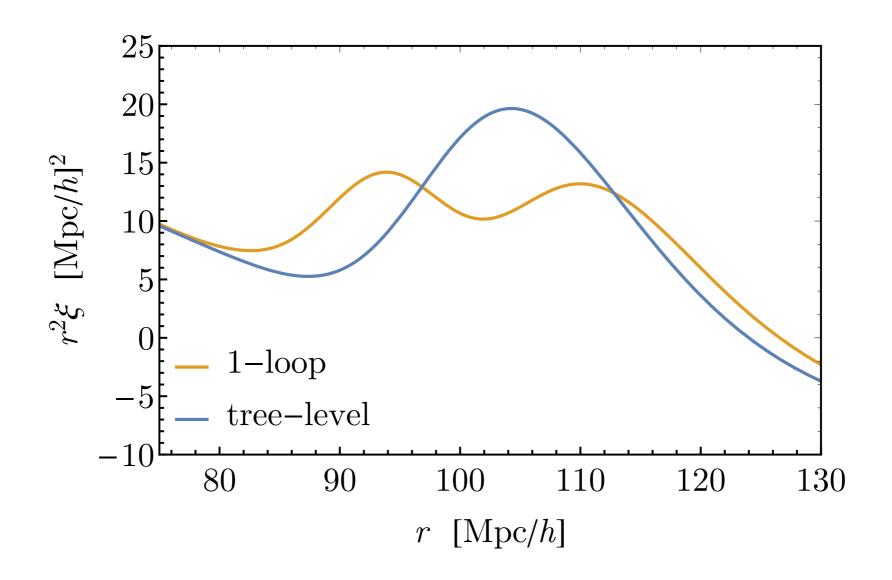


Perturbation Theory should recover the BAO, so the more loops the better, right?

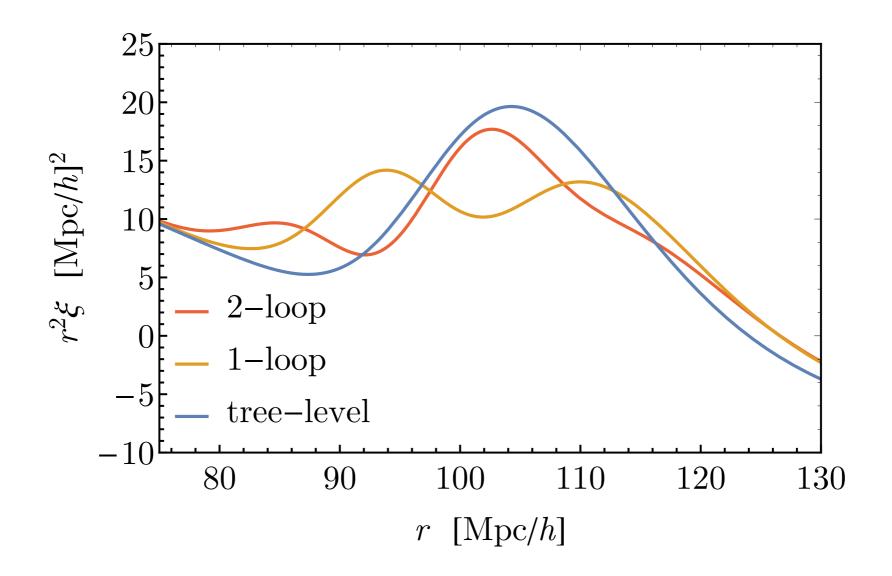
SPT completely fails around the BAO scale



SPT completely fails around the BAO scale



SPT completely fails around the BAO scale



## Why SPT fails?

The wiggly component (BAO) receive large infrared (IR-enhanced) contribution from loop integrals

$$P_{1-\text{loop}}^{w}(k) \sim k^{2} P_{\text{lin}}^{w}(k) \int_{\ell_{\text{BAO}}^{-1} \lesssim p \lesssim k} \frac{\mathrm{d}^{3} p}{(2\pi)^{3}} \frac{P_{\text{lin}}(p)}{p^{2}}$$
$$= P_{\text{lin}}^{w}(k) \epsilon_{s}$$

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$$= P_{\mathrm{lin}}^w(k) \epsilon_{s_<}$$
 and  $\epsilon_{s_<} \approx 1$ 

Bad for doing PT! Better not to expand

### To fix just resum

Lagrangian PT (à la Zeldovich)

short modes should not be resummed, calculations are more cumbersome, especially EFT, Fourier space is numerically more challenging

- IR-resummation (à la EPT) Senatore, Zaldarriaga '15 numerically more demanding
- Consistency relations (using EP) Baldauf et. al '15 split  $P_{lin}$  into smooth and wiggly components, NLO corrections
- Time-Sliced PT (à la QFT) Blas et al. '15

need to split  $P_{iin}$  into smooth and wiggly components manually check the IR-enhanced contribution resummation relies on separation of scales calculations are more cumbersome, especially NLO and UV

The idea is to resum IR displacement modes

$$K_0(\boldsymbol{k}, \boldsymbol{q}) \equiv \exp\left[-\frac{1}{2}k_ik_jA_{ij}^{\mathrm{IR}}(\boldsymbol{q})\right]$$

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$$K_0(\mathbf{k}, \mathbf{q}) \equiv \exp\left[-\frac{1}{2}k_i k_j A_{ij}^{\mathrm{IR}}(\mathbf{q}) + \frac{i}{6}k_i k_j k_k B_{ijk}^{\mathrm{IR}}(\mathbf{q})\right]$$

$$\sim \epsilon_{s<}$$

The idea is to resum IR displacement modes

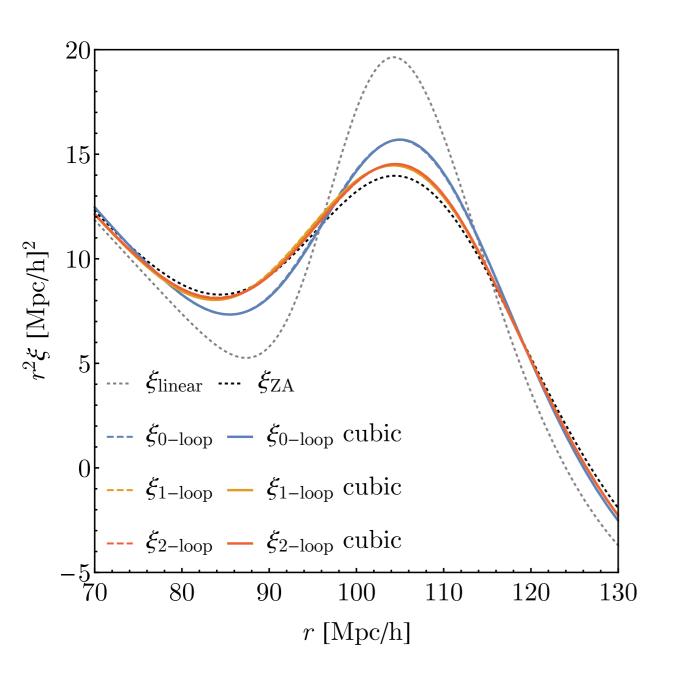
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$$\sim \epsilon_{s<}$$

For example at tree-level

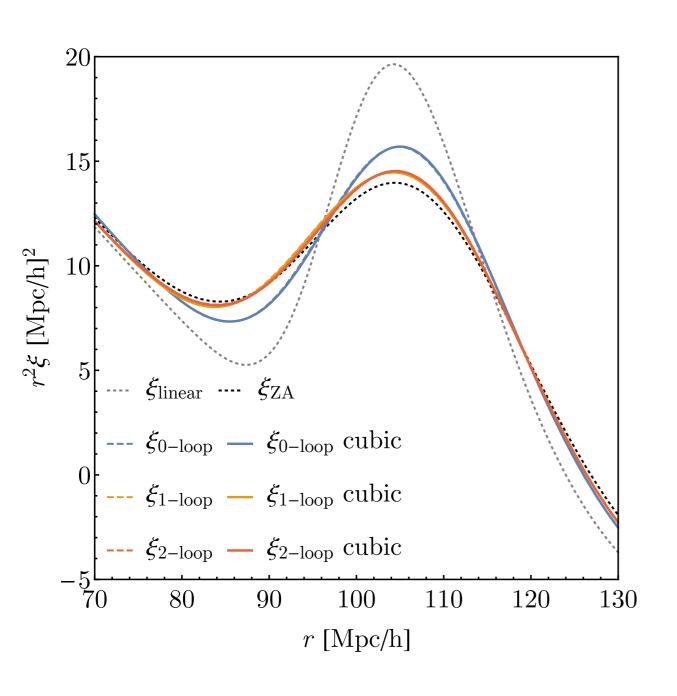
$$\xi_{\text{tree}}(r) = \frac{1}{|A_{ij}(\boldsymbol{r})|^{1/2}} \int d^3q \; \xi_{\text{tree}}^{\text{E}}(q) \; \exp\left[-\frac{1}{2}(\boldsymbol{r} - \boldsymbol{q})_i A_{ij}^{-1}(\boldsymbol{r})(\boldsymbol{r} - \boldsymbol{q})_j\right].$$

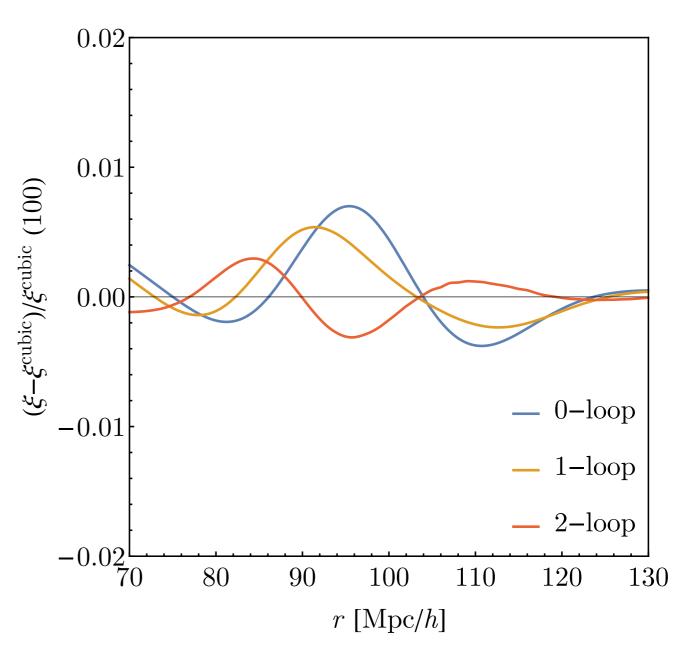
#### IR-resummation with NLO terms



Senatore, Trevisan: arXiv 1710.02178

#### IR-resummation with NLO terms

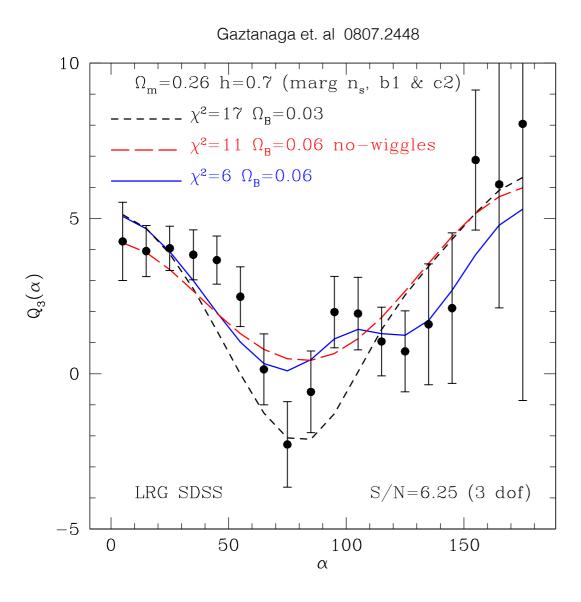




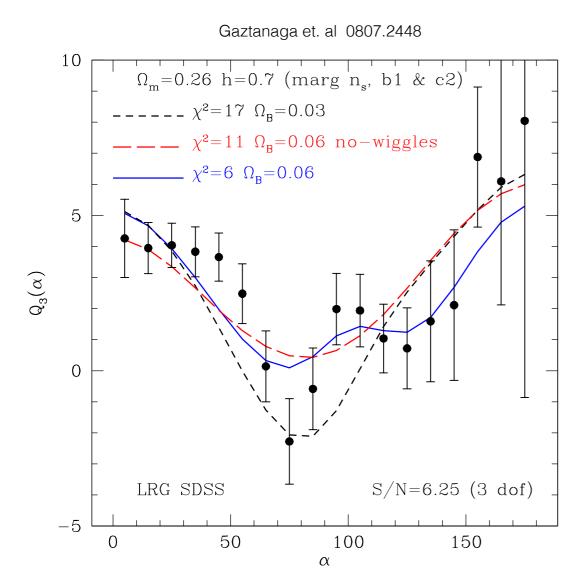
Senatore, Trevisan: arXiv 1710.02178



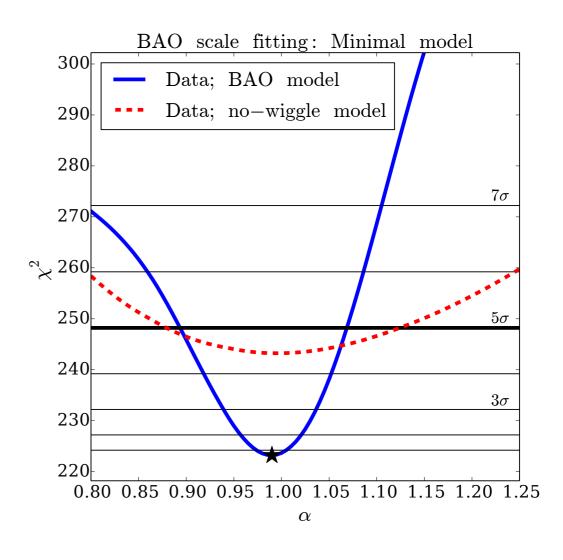
#### Detection of the BAO in the 3-PF



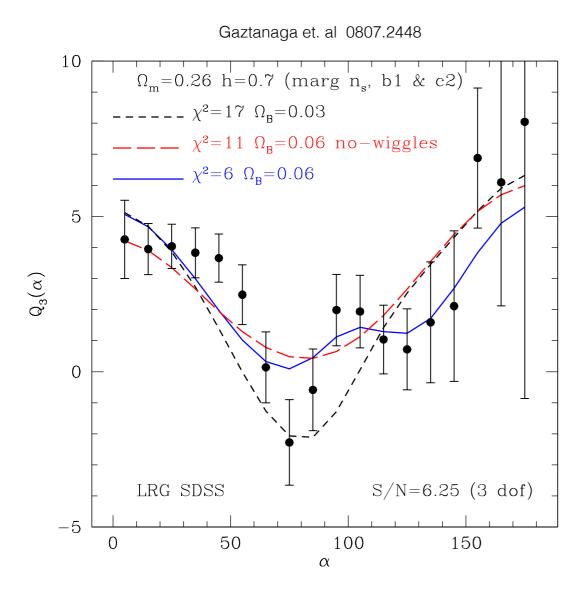
#### Detection of the BAO in the 3-PF



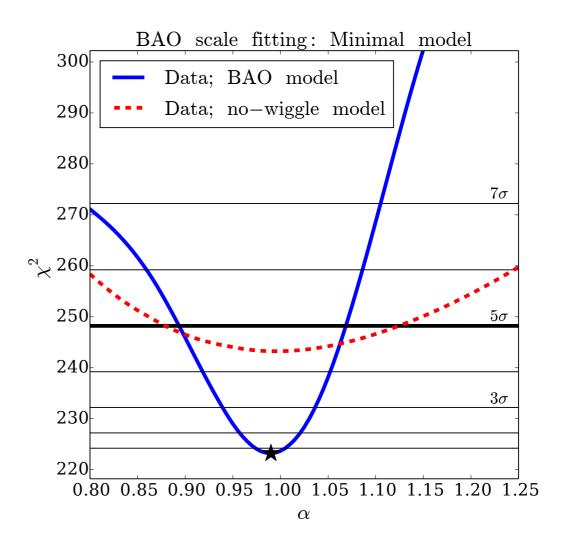
Slepian et al. 1607.06097



#### Detection of the BAO in the 3-PF



Slepian et al. 1607.06097

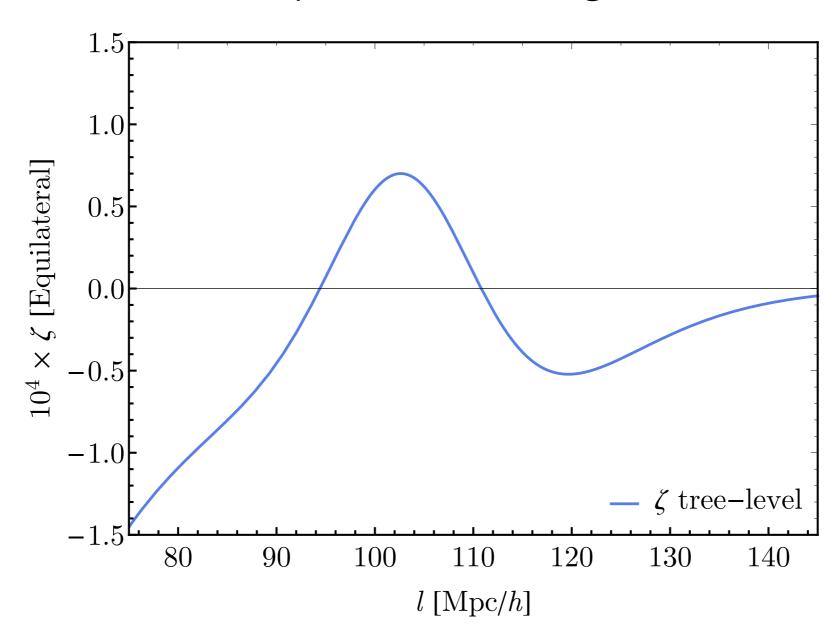


Analyses use  $P_{\text{phys}}(k) = [P(k) - P_{\text{nw}}(k)] \exp [-k^2 \Sigma_{\text{nl}}^2 / 2] + P_{\text{nw}}(k)$ ,

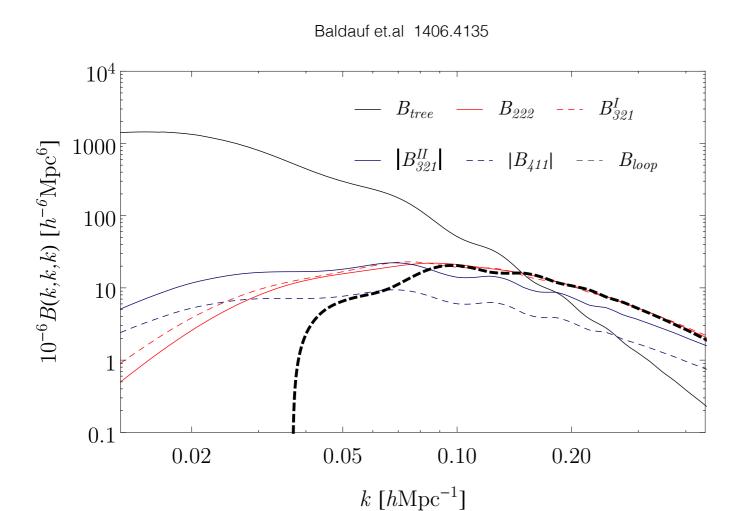
plugged into the tree-level SPT 3-PF

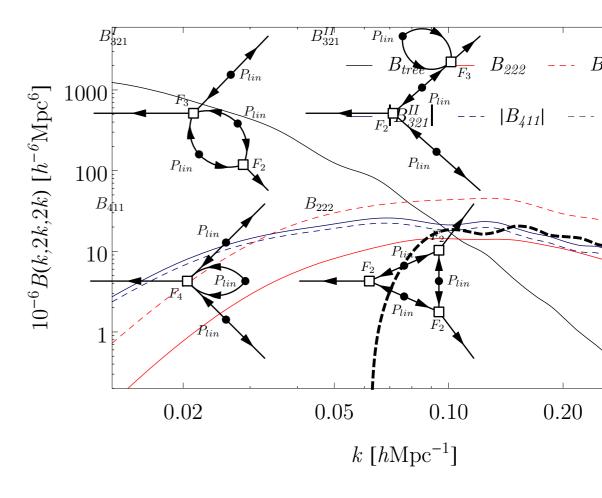
## SPT predictions for the 3-PF

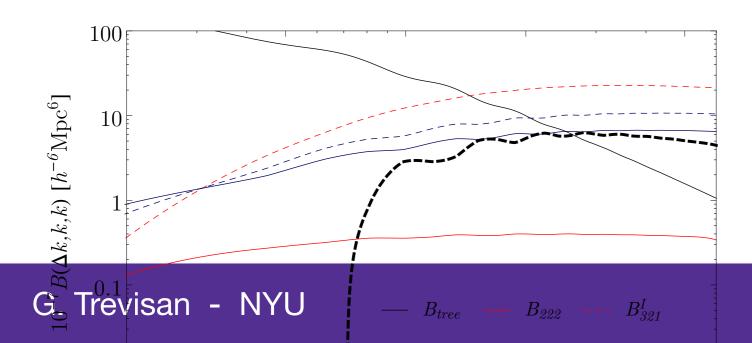
#### Equilateral triangles



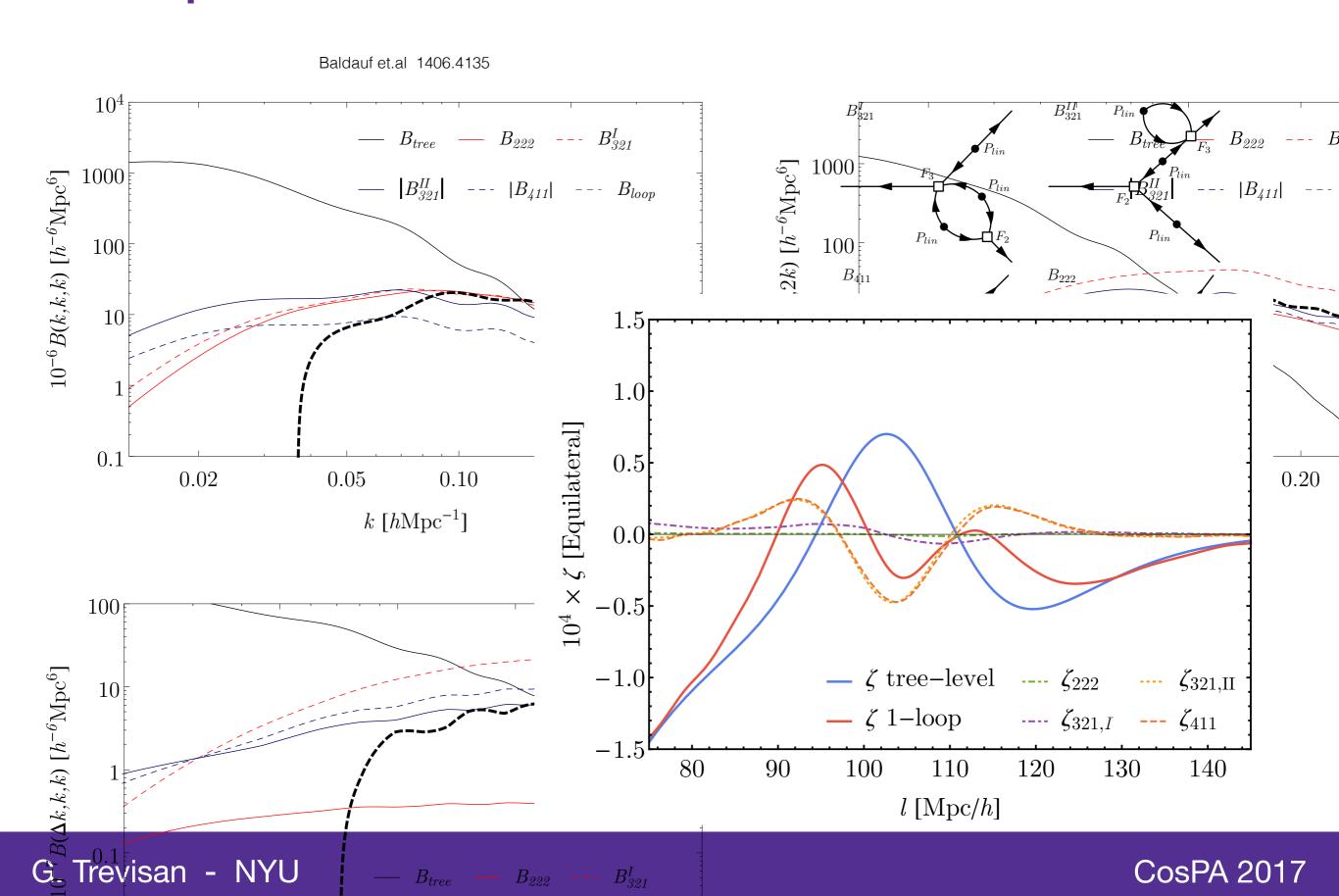
## SPT predictions for the 3-PF







## SPT predictions for the 3-PF



#### IR-resummation for the 3-PF

Similar to the resummation of the 2PF For example at tree-level

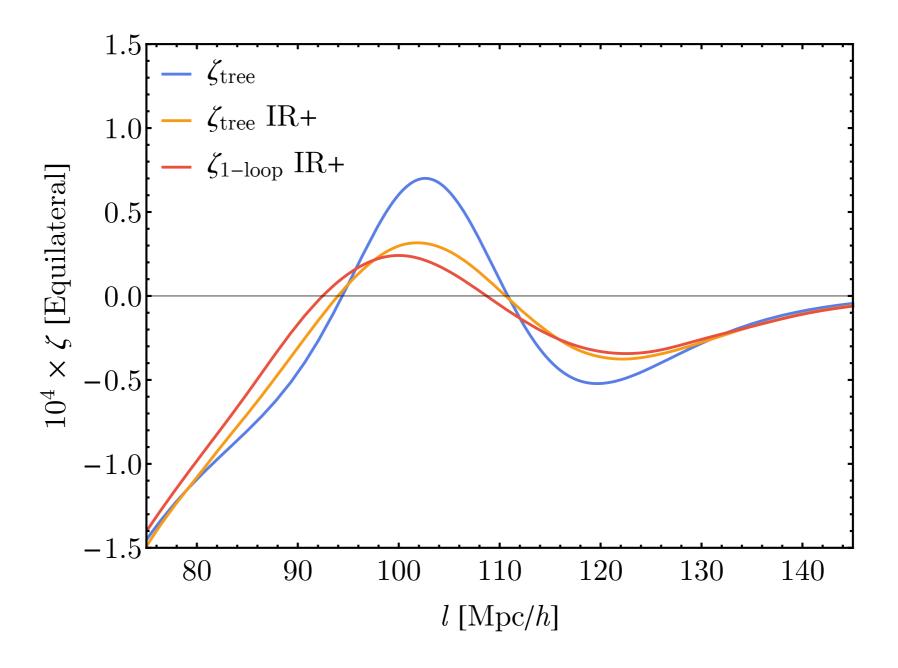
$$\zeta_{tree}^{\mathrm{IR}+}(\vec{x}_1, \vec{x}_2, \vec{x}_3) = \int \mathrm{d}^6 r \, \zeta_{tree}^{\mathrm{E}}(\boldsymbol{r}) G(\boldsymbol{x}, \boldsymbol{x} - \boldsymbol{r})$$

Gaussian-like kernel induced by long displacement modes

Scoccimarro, Trevisan: in preparation

#### IR-resummation for the 3-PF

Similar to the resummation of the 2PF



Scoccimarro, Trevisan: in preparation

#### Conclusions

- A simpler approx. for the IR-resummation
- Although IR-enhanced, IR mode-coupling leads to <1% in the 2-PF</p>
- IR-resummation fixes also the 3-PF
- Tree-level and 1-loop already agree quite well
- Analysis of 2+3 PF may be an alternative to reconstruction

## Thanks!

Instead of using comoving coordinates, use fluid coordinates

$$\vec{x}(\vec{q},t) = \vec{q} + \vec{s}(\vec{q},t)$$

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$$1 + \delta(\vec{x}, t) = \int d^3q \, \delta_D^3(\vec{x} - \vec{q} - \vec{s}(\vec{q}, t))$$

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to obtain the 2-PF

$$P(k) = \int d^3q_{12} e^{-i\vec{k}\cdot\vec{q}_{12}} \left\langle e^{-i\vec{k}\cdot(\vec{s}(q_1)-\vec{s}(q_2))} \right\rangle$$

The linear order solution for the displacement field is

$$\vec{s}(p) \simeq \vec{s}_1(p) = i \frac{\vec{p}}{p^2} \delta_{\text{lin}}(p),$$

and leads to the Zel'dovich approximation:

### A brief excursus on Lagrangian PT (LPT)

The linear order solution for the displacement field is

$$\vec{s}(p) \simeq \vec{s}_1(p) = i \frac{\vec{p}}{p^2} \delta_{\text{lin}}(p),$$

and leads to the Zel'dovich approximation:

$$P(k) = \int d^3q_{12} e^{-i\vec{k}\cdot\vec{q}_{12}} e^{-\frac{1}{2}k_ik_j\langle s_is_j\rangle(q_{12})}$$

where

$$\langle s_i s_j \rangle \sim \int \frac{\mathrm{d}^3 p}{(2\pi)^3} \frac{P_{\mathrm{lin}}(p)}{p^2}$$

Lets go back to the 1-loop expression

$$P_{1-\text{loop}}(k) \sim \frac{1}{2} \int_{p \ll \Lambda} \frac{\mathrm{d}^3 p}{(2\pi)^3} \frac{(\boldsymbol{p} \cdot \boldsymbol{k})^2}{p^4} \left[ P_{\text{lin}}(|\boldsymbol{k} - \boldsymbol{p}|) + P_{\text{lin}}(|\boldsymbol{k} + \boldsymbol{p}|) - 2P_{\text{lin}}(|\boldsymbol{k}|) \right] P_{\text{lin}}(p).$$

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$$P_{\text{lin}} \propto k^n$$

For a smooth component  $\sim P_{\text{lin}}(k) \frac{p^2}{l \cdot 2}$ ,

$$\sim P_{\rm lin}(k) \frac{p^2}{k^2}$$

Very long modes ( $p \ll k$ ) do not contribute to the loop

Lets go back to the 1-loop expression

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$$P_{\text{lin}} \propto \sin(k/k_{osc})$$

For the BAO 
$$\sim P_{\rm lin}^w(k) \left(\cos\left(p\ell_{\rm BAO}\right) - 1\right)$$

So there is an IR-enhancement for modes  $\ell_{\mathrm{BAO}}^{-1} \lesssim p \lesssim k$ 

LPT calculations involve the average of an exponential

$$P(k) = \int d^3q_{12} e^{-i\vec{k}\cdot\vec{q}_{12}} \left\langle e^{-i\vec{k}\cdot(\vec{s}(q_1)-\vec{s}(q_2))} \right\rangle$$

which can be done as

$$\left\langle e^{-i\boldsymbol{k}\cdot\boldsymbol{\Delta}(\boldsymbol{q})}\right\rangle = \exp\left[\sum_{n=1}^{\infty}\frac{(-i)^n}{n!}\left\langle (\boldsymbol{k}\cdot\boldsymbol{\Delta}(\boldsymbol{q}))^n\right\rangle_c\right] = K(\boldsymbol{k},\boldsymbol{q}),$$

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Once expanded to some order N in  $P_{lin}$ 

$$LPT = SPT$$

and the idea is to resum IR modes ( $\sim \epsilon_{s<}$ )

$$X|_{N} \\ \text{means up to order } \\ N$$

$$K_0$$
 contains only IR-displacements

$$K^{\text{IR+}}|_{N} = K_{0} \cdot \frac{K}{K_{0}}|_{N}$$

$$= \sum_{i=0}^{N} R|_{N-i} \cdot K_{j}$$

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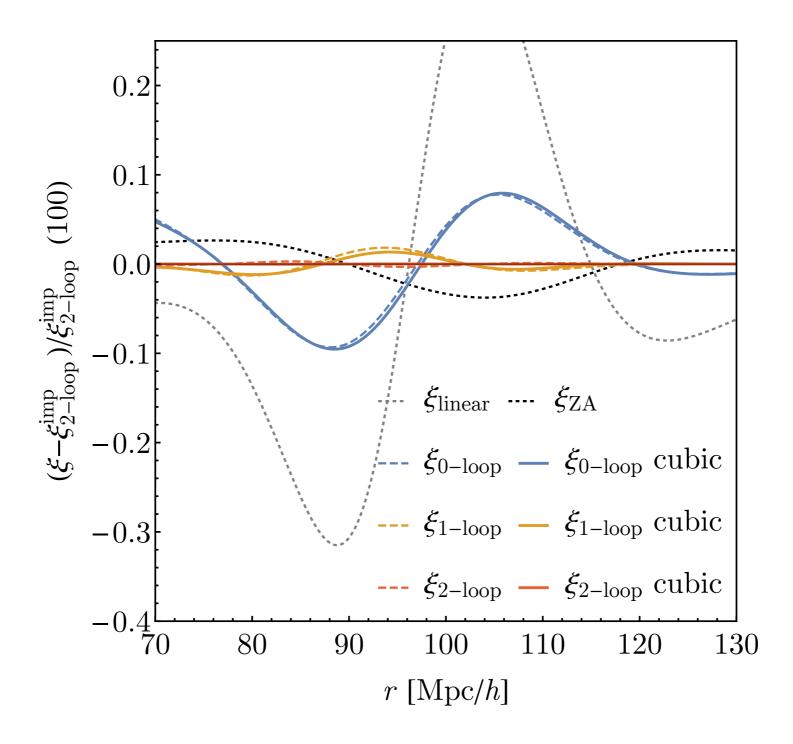
to get\*

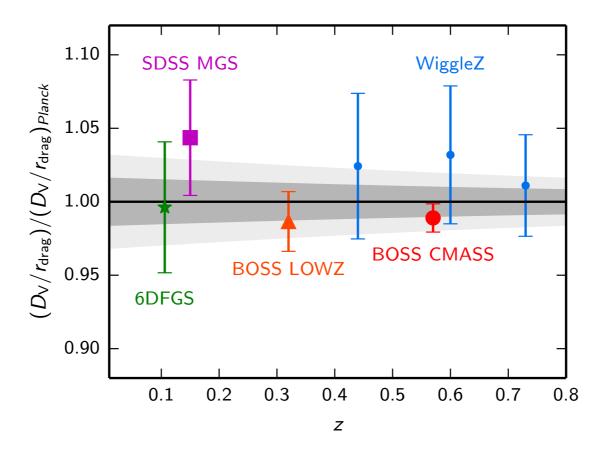
$$P(k) = \int d^3q \, e^{-i\boldsymbol{q}\cdot\boldsymbol{k}} \sum_{j=0}^{N} R(\boldsymbol{k}, \boldsymbol{q})||_{N-j} \cdot \xi_{||j}^{\mathrm{E}}(q)$$

\*we actually use a much simpler and intelligible approximation wrt the original paper which is parametrically justified

$$\zeta(\vec{x}_1, \vec{x}_2, \vec{x}_3) = \frac{10}{7} \xi_2(\vec{x}_{12}) \xi_2(\vec{x}_{13}) + \nabla_i^{-1} \xi_2(\vec{x}_{12}) \nabla_i \xi_2(\vec{x}_{13}) + \nabla_i \xi_2(\vec{x}_{12}) \nabla_i^{-1} \xi_2(\vec{x}_{13}) 
+ \frac{4}{7} \nabla_i \nabla_j^{-1} \xi_2(\vec{x}_{12}) \nabla_i \nabla_j^{-1} \xi_2(\vec{x}_{13}) + \text{cyc.}$$

### Back-up





**Fig. 14.** Acoustic-scale distance ratio  $D_{\rm V}(z)/r_{\rm drag}$  in the base  $\Lambda {\rm CDM}$  model divided by the mean distance ratio from *Planck* TT+lowP+lensing. The points with  $1\,\sigma$  errors are as follows: green star (6dFGS, Beutler et al. 2011); square (SDSS MGS, Ross et al. 2015); red triangle and large circle (BOSS "LOWZ" and CMASS surveys, Anderson et al. 2014); and small blue circles (WiggleZ, as analysed by Kazin et al. 2014). The grey bands show the 68 % and 95 % confidence ranges allowed by *Planck* TT+lowP+lensing.

$$D_{\rm V}(z) = \left[ (1+z)^2 D_{\rm A}^2(z) \frac{cz}{H(z)} \right]^{1/3}.$$

### 1 loop TSPT

$$P_{w,\delta\delta}^{1-loop}(\eta;k)\Big|_{hard} = \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_4^s \\ \end{array}}_{\overline{\Gamma}_4^s} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{K_3} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{K_3} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{K_3} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{\overline{\Gamma}_5^s} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{\overline{\Gamma}_5^s} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{\overline{\Gamma}_5^s} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_3^s \\ \end{array}}_{\overline{\Gamma}_7^s} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_7^s \\ \end{array}}_{\overline{\Gamma}_7^w} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^s \\ \overline{\Gamma}_7^w \\ \end{array}}_{\overline{\Gamma}_7^w} + \underbrace{\begin{array}{c} \overline{\Gamma}_3^w \\ \overline{\Gamma}_7^w \\ \end{array}}_{\overline{\Gamma}_7^w} + \underbrace{\begin{array}{c} \overline{\Gamma}_7^w \\ \overline{\Gamma}_7^w \\ \end{array}}_{\overline{\Gamma}_7^w} +$$

The perturbative solution parametrically goes as

$$s_n \sim s_1 \delta_1^{n-1}$$