

# RH neutrino dark matter in a flavoured $B-L$ model

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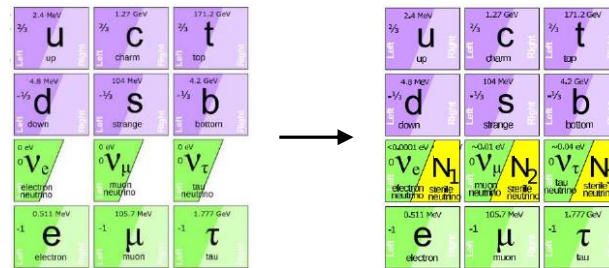
In collaboration with Chengcheng Han, Tsutomu T. Yanagida

arXiv:1710.01585, arXiv:1705.03858



# RH Neutrino Dark Matter

- Simple, well-motivated extension of Standard Model is addition of 3  $\nu_R$



- Very massive  $\nu_R$  ( $M \gtrsim 10^9$  GeV) can explain:
  - Smallness of active neutrino masses via seesaw mechanism
  - Observed baryon asymmetry via thermal leptogenesis
- However, *two*  $\nu_R$  are sufficient for both see-saw and leptogenesis (Frampton, Glashow, Yanagida '02)
- Why include the 3<sup>rd</sup>  $\nu_R$  at all? **→ Dark Matter!**

# $\nu_R$ DM *and* high-scale see-saw

- Many studies of RH/sterile neutrino DM ...

Often with  $\sim$ keV DM, and all 3  $\nu_R$  below the EW scale (e.g. “vMSM” Asaka, Shaposhnikov '05)

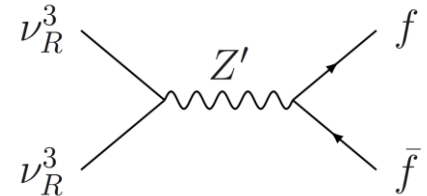
**Alternative approach:** assume two  $\nu_R$  remain very massive

- Keep full benefit of see-saw, and (non-resonant) leptogenesis

- But, need a production mechanism for lightest  $\nu_R$

Obvious choice: a new gauge symmetry

$$SU(3)_C \times SU(2)_L \times U(1)_Y \times U(1)'$$



- $\nu_R^3$  produced via thermal freeze-out (i.e. WIMP)
- Majorana mass for  $\nu_R^3$  generated by spontaneous symmetry breaking (partially) explains hierarchy in  $\nu_R$  masses

# Choosing a $U(1)$ symmetry

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Interested in vectorial  $U(1)$  symmetries that allow:

- Two  $\nu_R$  have large Majorana masses for leptogenesis
- Lightest  $\nu_R$  charged under  $U(1)$  and is DM

→ Completely fixes lepton charge

$$Q_l = (0, 0, -1)$$

In quark sector, assume

- Suppression of FCNCs ( $K - K$  and  $D^0 - D^0$  mixing)

Anomaly cancellation (SM+3 $\nu_R$ ) then restricts to a *single-parameter* class:

$$Q_q = \left( a, a, \frac{1}{3} - 2a \right)$$

# Choosing a U(1) symmetry

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One parameter family of U(1) symmetries:

$$Q_l = (0, 0, -1) \quad Q_q = \left( a, a, \frac{1}{3} - 2a \right)$$

- Particularly interesting case is  $a = 0$ 
  - *Flavoured B – L* symmetry
  - Anomaly cancellation within each generation (like SM)
  - Likely the least constrained choice, since only 3<sup>rd</sup> gen charged

# Flavoured B-L

- $U(1)_{(B-L)_3}$  :  $B - L$  gauge symmetry under which only 3<sup>rd</sup> generation charged

	$q_L^3, b_R, t_R$	$\ell_L^3, \tau_R, \nu_R^3$	$H$	$\Phi$
$Q_{(B-L)_3}$	+1/3	-1	0	+2

- Introduce  $\Phi(+2)$  to spontaneously break symmetry

- Also generates Majorana mass for dark matter  $\chi = (-\varepsilon\nu_R^{3*}, \nu_R^3)^T$

$$\mathcal{L} = \frac{i}{2}\bar{\chi}\not{\partial}\chi + \frac{g}{2}Z'_\mu\bar{\chi}\gamma^5\gamma^\mu\chi - \left(\frac{y}{2}\bar{\chi}\Phi P_R\chi + h.c.\right)$$

# Yukawa Structure

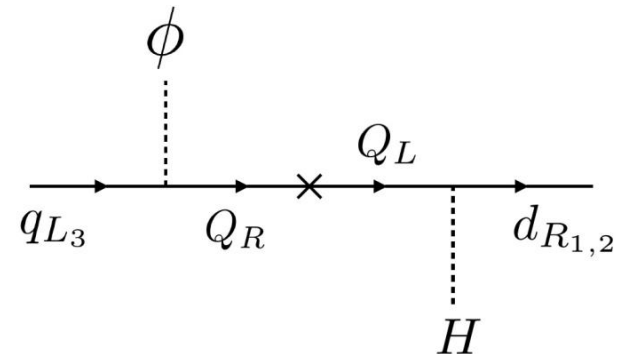
- However, off-diagonal Yukawa couplings involving 3<sup>rd</sup> generation are now forbidden

$$Y_d = \begin{pmatrix} \hat{Y}_d^{2 \times 2} & 0 \\ 0 & Y_b \end{pmatrix}$$

→ Require a mechanism to generate these upon  $U(1)_{(B-L)_3}$  breaking

Two general possibilities:

- Additional Higgs doublets charged under  $U(1)_{(B-L)_3}$
- New vector-like fermions



# DM Stability

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- Impose  $\mathbb{Z}_2$  symmetry to guarantee DM stability

RH  $\nu$  DM (-), everything else (+)

- Forbids dangerous Yukawa couplings

$$Y_\nu = \begin{pmatrix} \hat{Y}_\nu^{2 \times 2} & 0 \\ -\frac{Y'_L \phi_l}{M_L} Y_L^T & 0 \end{pmatrix}$$

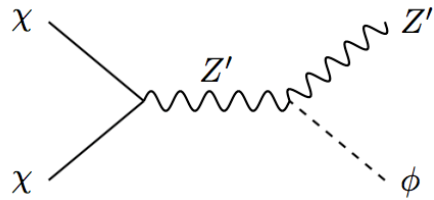
- Simultaneously solves potential problem with light RH neutrino raising active neutrino masses, and washing out lepton asymmetry

$$\frac{1}{M_{\nu_R}} \simeq \begin{pmatrix} \frac{1}{\hat{M}_{\nu_R}^{2 \times 2}} & 0 \\ 0 & \frac{1}{m_\chi} \end{pmatrix}$$

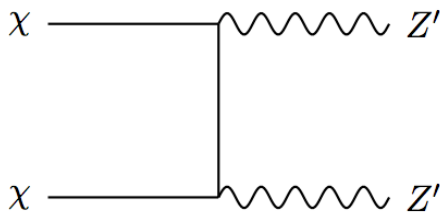


# DM Annihilation

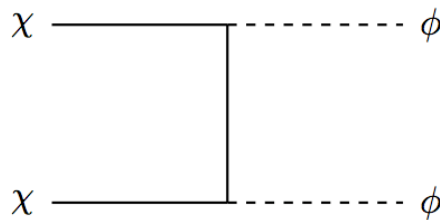
Four main annihilation channels:



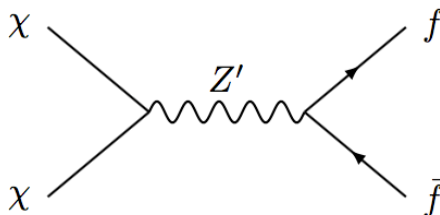
s-wave, dominates when kinematically open



s-wave, p-wave enhanced if  $m_\chi > m_{Z'}$



p-wave



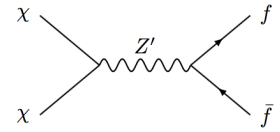
p-wave

# Relic density

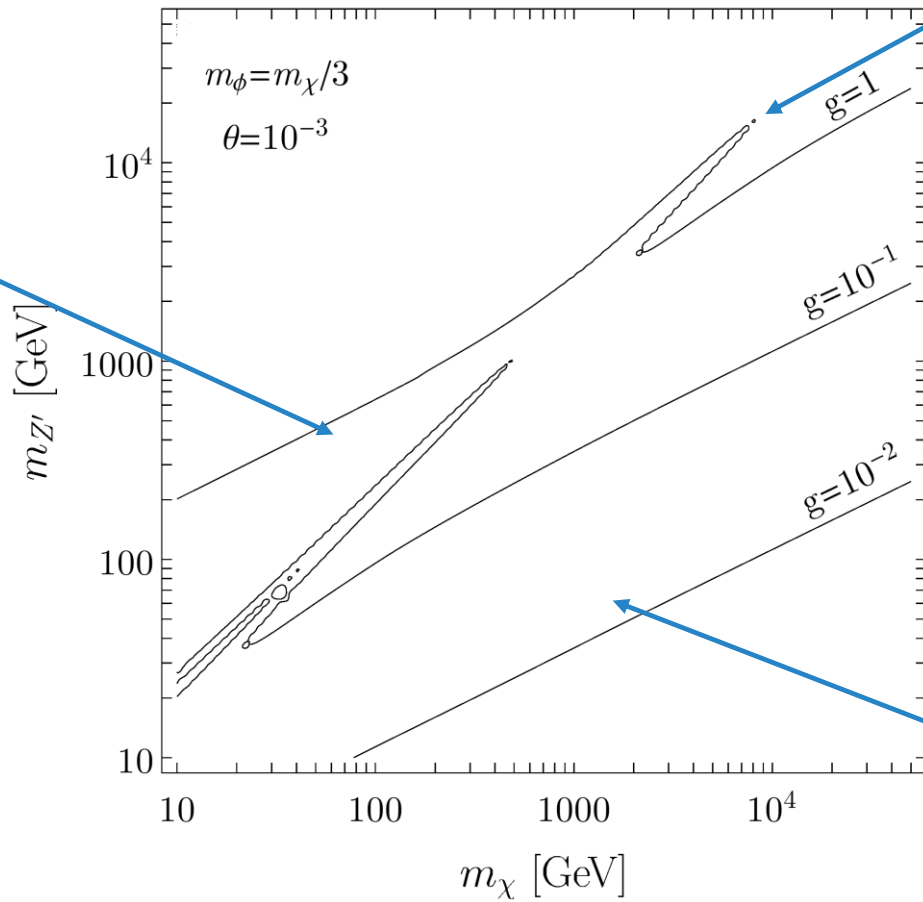
➤ Fix gauge coupling to satisfy correct relic density

→ 5 free parameters:  $m_\chi$ ,  $m_{Z'}$ ,  $m_\phi$ ,  $\theta$  ( $\phi$ -h mixing),  $\epsilon$  ( $Z$ - $Z'$  mixing)

$$m_\chi \approx m_{Z'}/2$$



$\chi\chi \rightarrow \phi\phi$   
 $\chi\chi \rightarrow ff$



$\chi\chi \rightarrow \phi Z'$   
 $(\chi\chi \rightarrow Z' Z')$

# Unitarity/perturbativity

- Even in simplified models, partial wave perturbative unitarity gives strong bounds

$\phi\phi, Z'_L Z'_L$ :

$$m_\phi < \frac{\sqrt{\pi} m_{Z'}}{g}$$

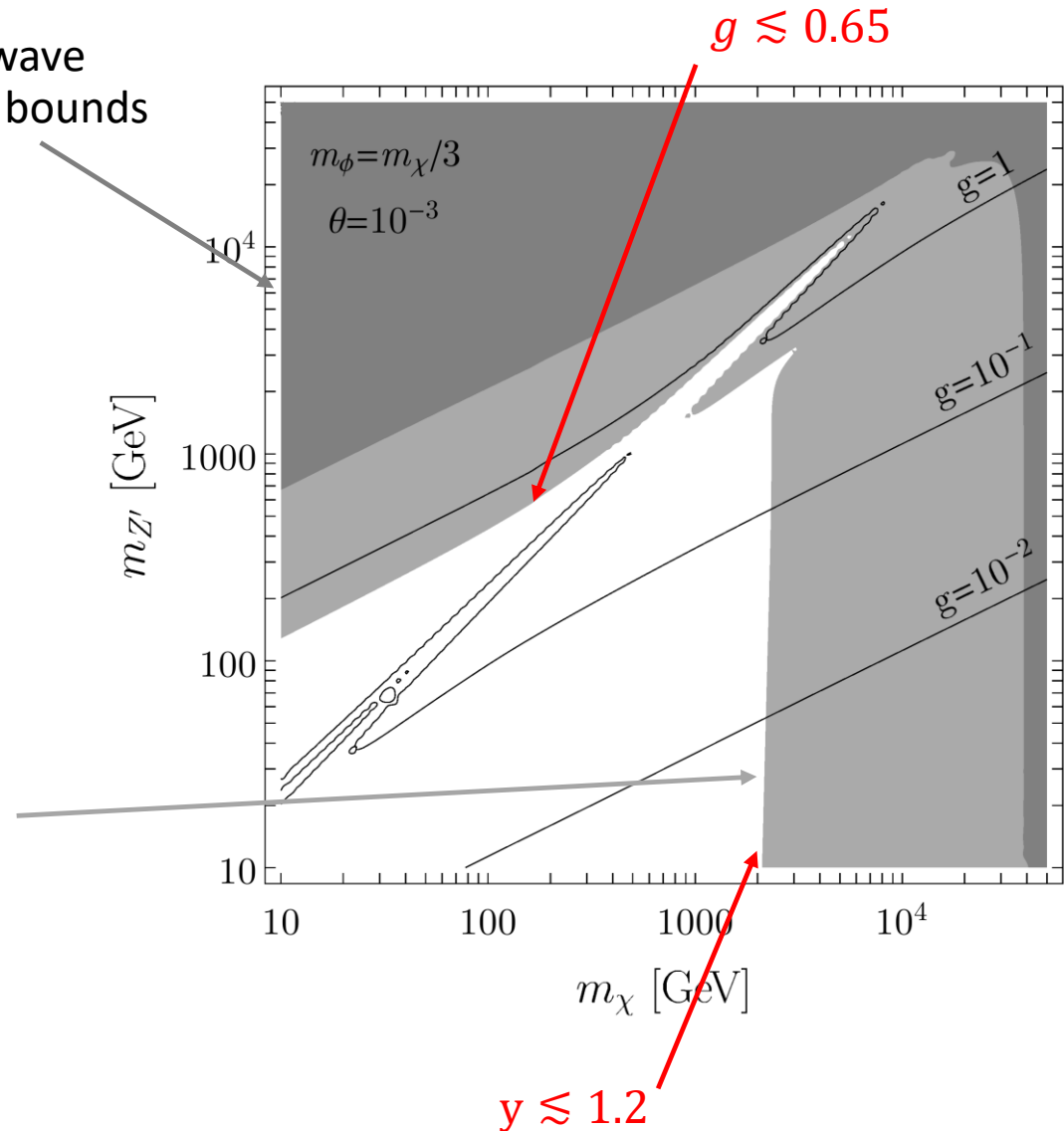
$\chi\chi \rightarrow \chi\chi$ :

$$m_\chi < \frac{\sqrt{\pi} m_{Z'}}{g}, \quad g < \sqrt{4\pi}$$

- Can also impose perturbativity of couplings to high scales (i.e. no Landau pole below  $M_{Pl}$ )

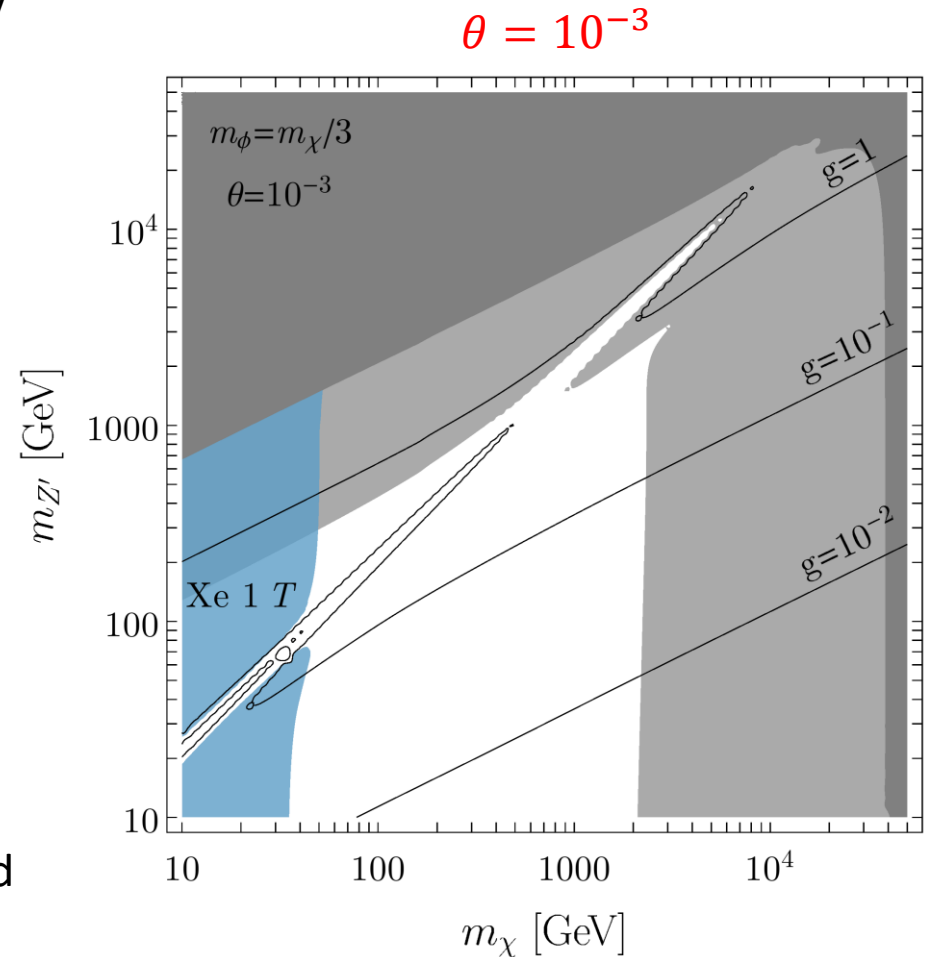
→ *Significantly* stronger bounds

$$m_\chi \lesssim 2 \text{ TeV}$$



# DM Direct Detection

- SI DM-nucleon scattering mediated by  $Z'$  is velocity suppressed
- Further suppression due to lack of  $Z'$  coupling to light quarks
- Higgs- $\phi$  mixing leads to spin-independent scattering
 
$$\mathcal{L} \supset (\phi^\dagger \phi)(H^\dagger H)$$
- At the very least, radiatively generated at two-loop



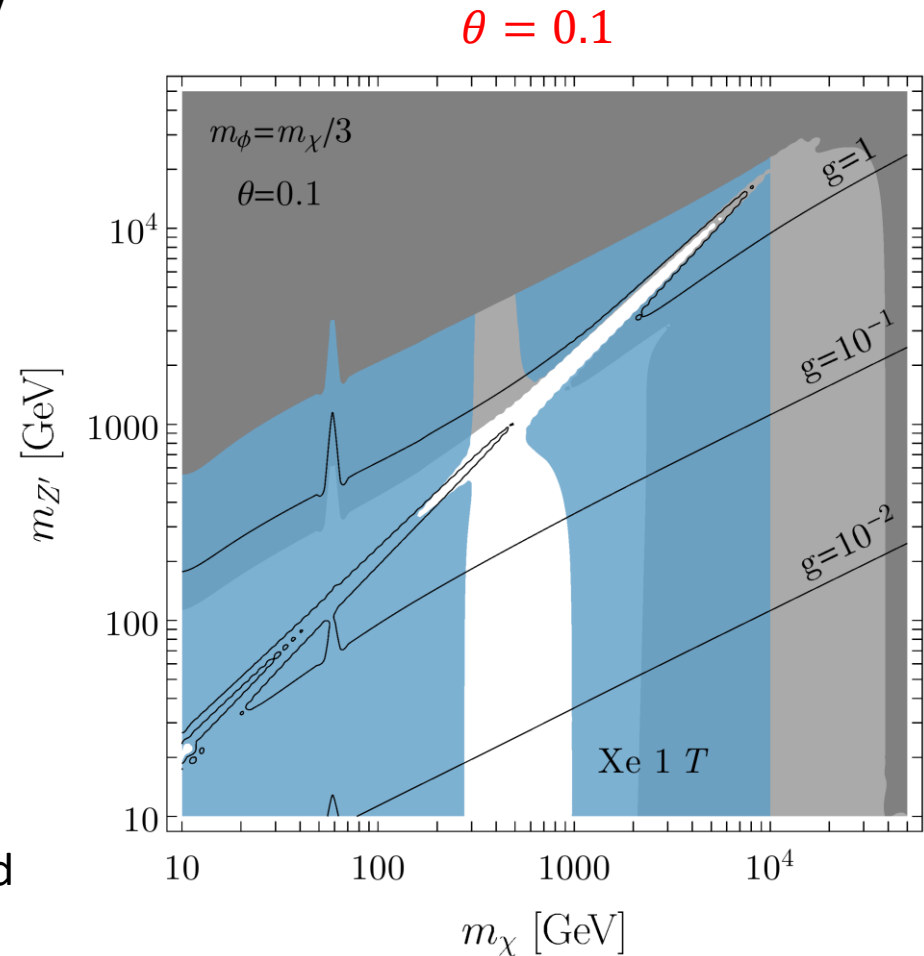
Limits even for very small mixing angles

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Larger mixing angles *strongly* constrained!

# DM Indirect Detection

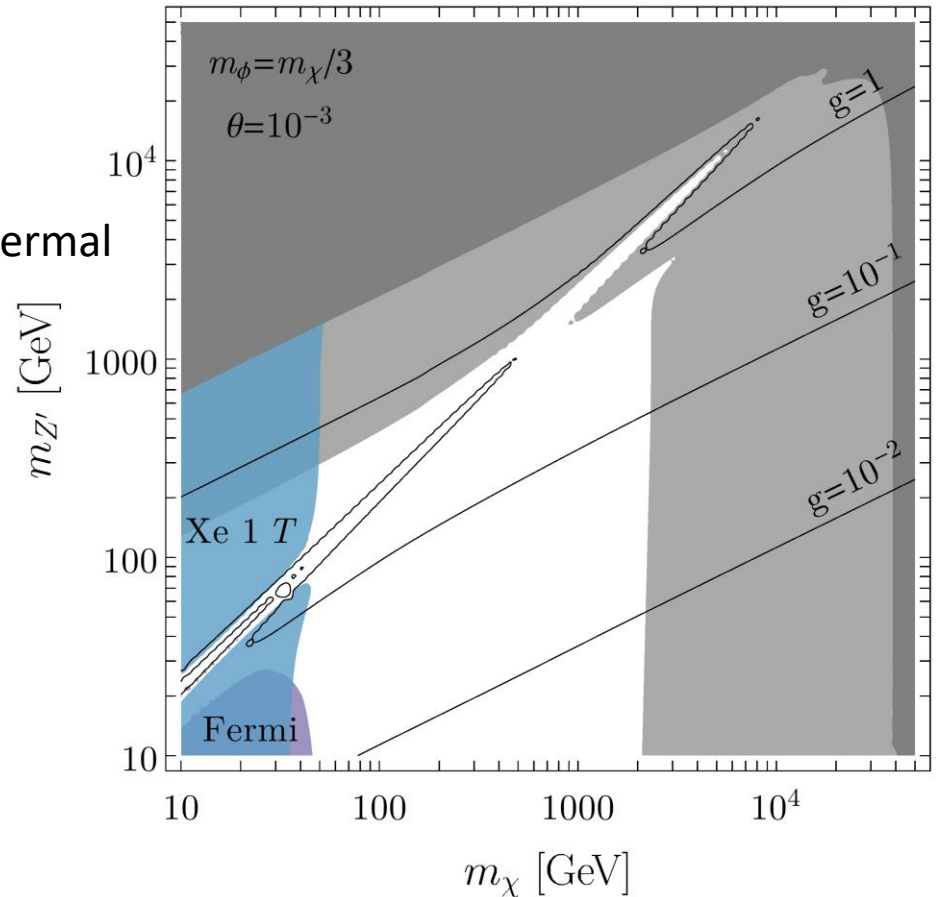
➤  $\chi\chi \rightarrow \phi Z'$  and  $\chi\chi \rightarrow Z'Z'$  annihilation are s-wave  
→ in principle lead to gamma-ray signals

➤  $\chi\chi \rightarrow Z'Z'$  : p-wave enhanced by  $m_\chi^4/m_{Z'}^4$   
→ xsec today suppressed compared to thermal relic xsec

➤ Only need to consider  $\chi\chi \rightarrow \phi Z'$

- Multibody SM final-state (eg.  $\phi Z' \rightarrow bbbb$ )
- Significant  $Z'$  BR to  $\nu$

→ Current Fermi dwarf limits very weak

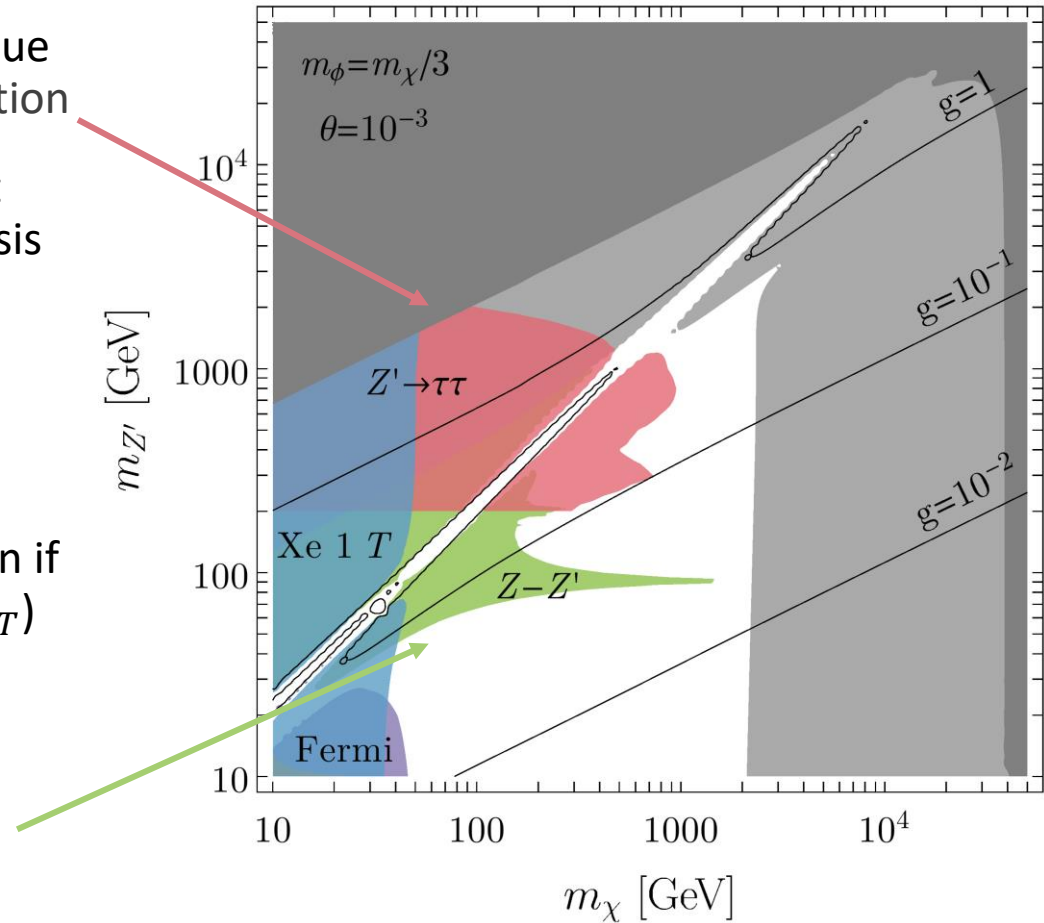


# Collider bounds

- Collider bounds relatively weak due to suppressed  $bb \rightarrow Z'$  cross-section
- Stronger limits if coupling to light leptons after rotation to mass basis
- Also have  $Z - Z'$  kinetic mixing
- Generated by RGE evolution, even if vanishing at high scales (e.g.  $\Lambda_{GUT}$ )

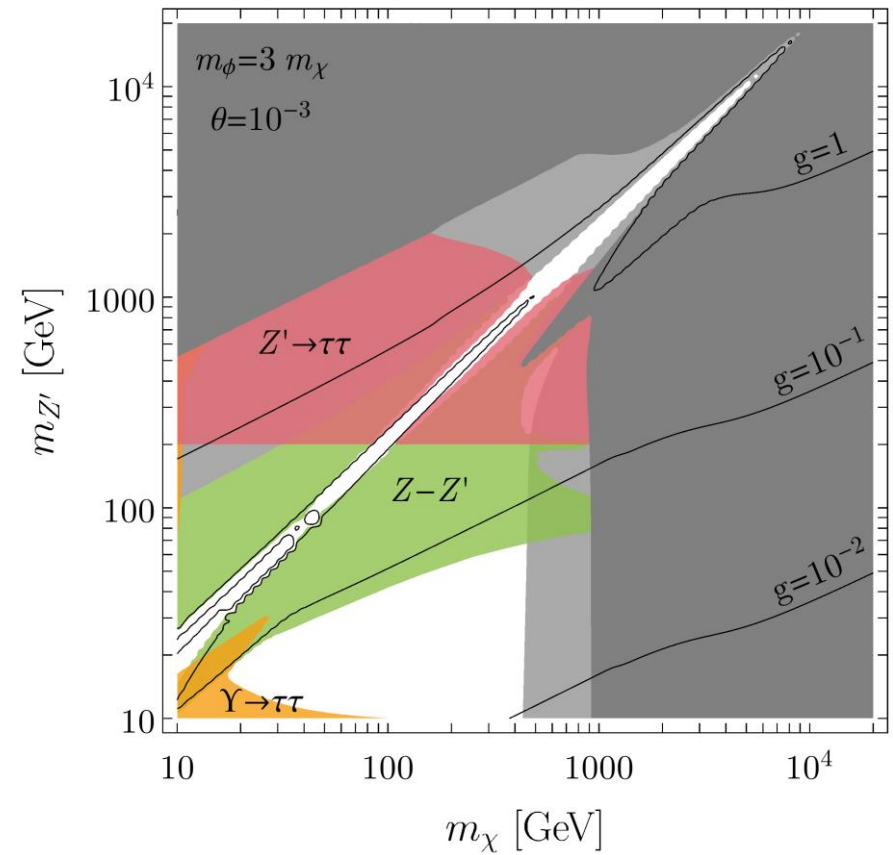
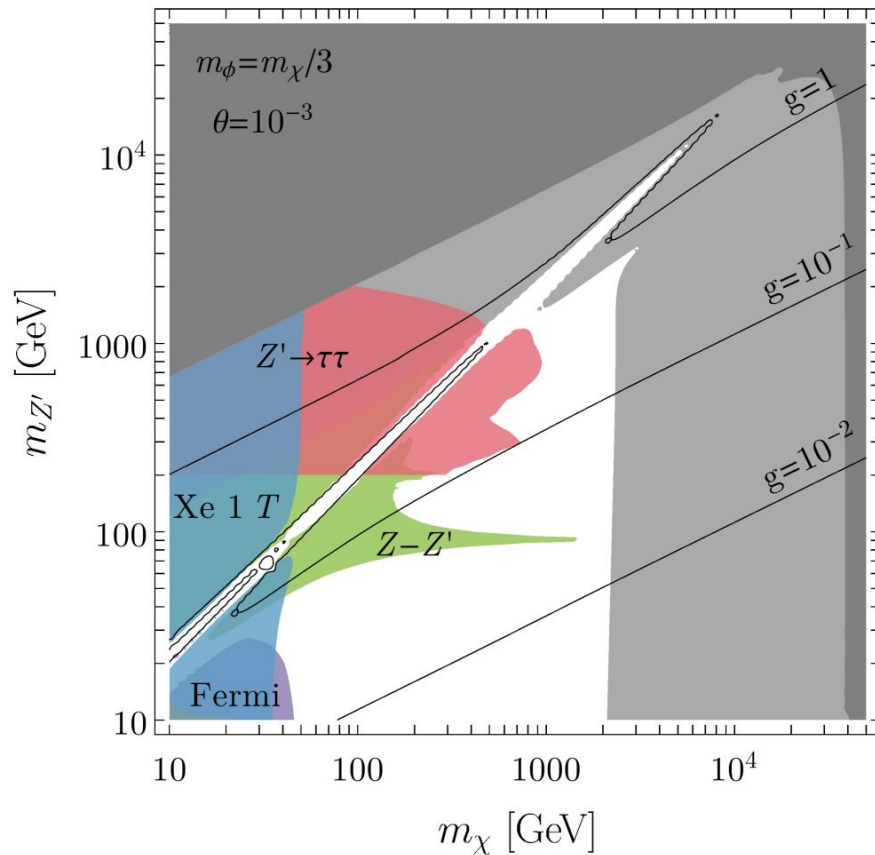
$$\varepsilon(m_{Z'}) \sim \frac{2g_Y g'}{9\pi^2} \log \frac{\Lambda}{m_{Z'}}$$

EW precision  $\rightarrow \varepsilon \leq 10^{-1} - 10^{-2}$



$$m_\phi = 3m_\chi$$

- For heavy scalar,  $\chi\chi \rightarrow \phi Z'$  annihilation channel forbidden
- Larger gauge coupling required for correct relic density  
→ stronger perturbativity constraints
- Reduced direct detection bound due to increased  $m_\phi$





# Summary

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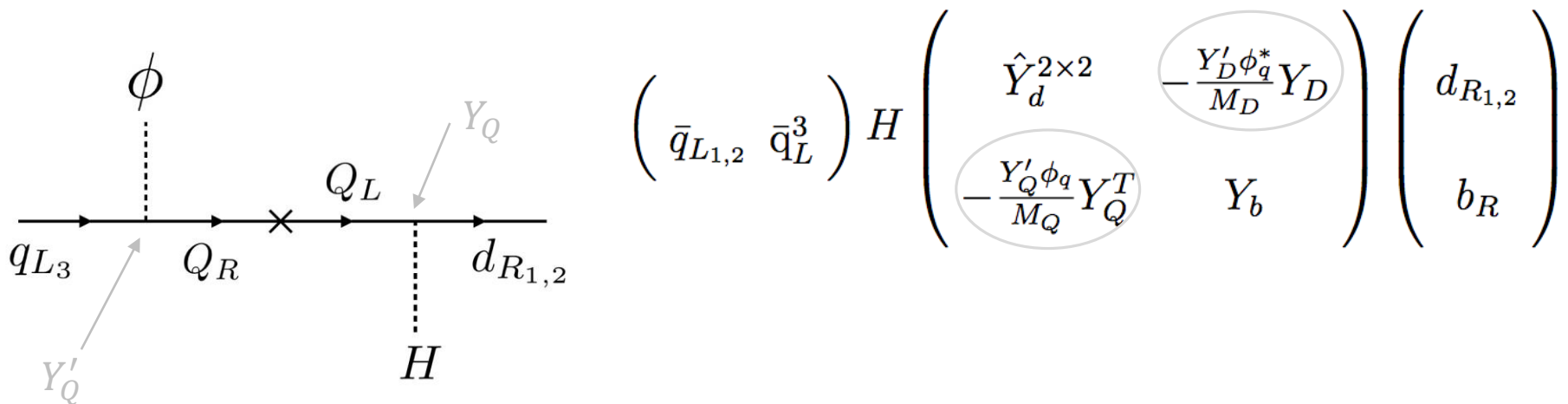
- Lightest RH neutrino can provide WIMP dark matter candidate in minimal gauged  $U(1)$  extensions
- Consistent with high-scale see-saw and leptogenesis
- Anomaly cancellation restricts possible symmetries:  
An interesting (and least constrained) model is *flavoured B-L* symmetry
- Significant viable parameter space remains to be explored
- Requiring perturbativity up to  $M_{Pl}$  gives a strong upper bound on DM mass ( $m_\chi \lesssim 2$  TeV)
- Also interesting connection to B-physics anomalies (see talk by Chengcheng Han)

# Backup

# Yukawa Structure

For general 3x3 Yukawa couplings, introduce:

- $U(1)_{(B-L)_3}$  neutral  $V$ - $L$  fermions:  $Q_{L,R}, U_{L,R}, D_{L,R}, L_{L,R}, E_{L,R}, N_{L,R}$
- SM singlet scalars ( $U(1)'$  breaking):  $\phi_l(+1), \phi_q(+\frac{1}{3})$



# Higgs- $\phi$ Mixing ( $\theta = 0.1$ )

- Large Higgs- $\phi$  mixing even more strongly constrained by direct detection
- Exception when  $m_\phi \approx m_h$  due to destructive interference

